

relensing

Reconstructing the mass profile of galaxy clusters from gravitational lensing

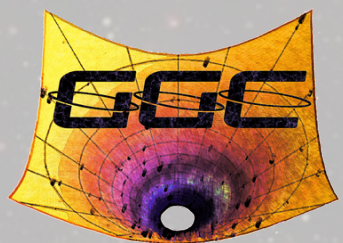
Based on [arXiv:2201.10076](https://arxiv.org/abs/2201.10076)

Daniel Alexdy Torres Ballesteros (daatorresba@unal.edu.co)

Supervisor: Prof. Leonardo Castañeda Colorado (lcastanedac@unal.edu.co)

Universidad Nacional de Colombia

Cosmology from home 2022



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Reconstructing the mass profile of galaxy clusters from gravitational lensing

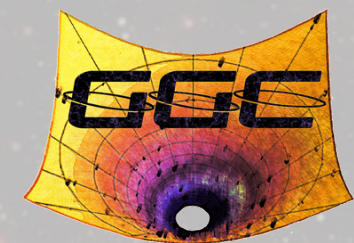
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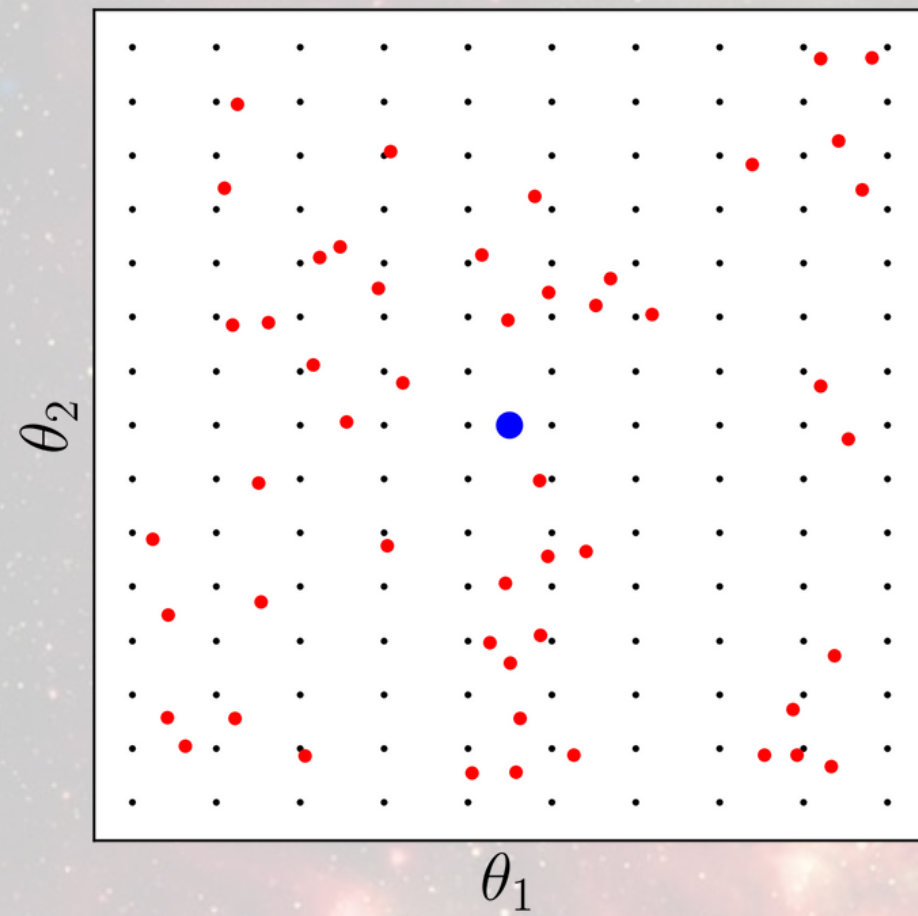
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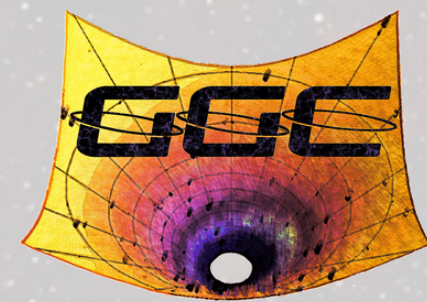
Cosmology from home 2022



GRID

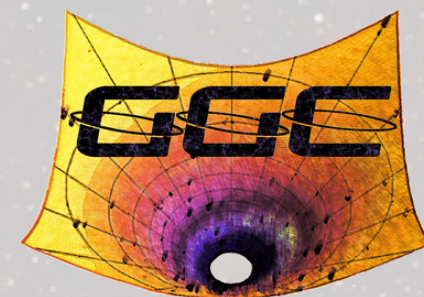
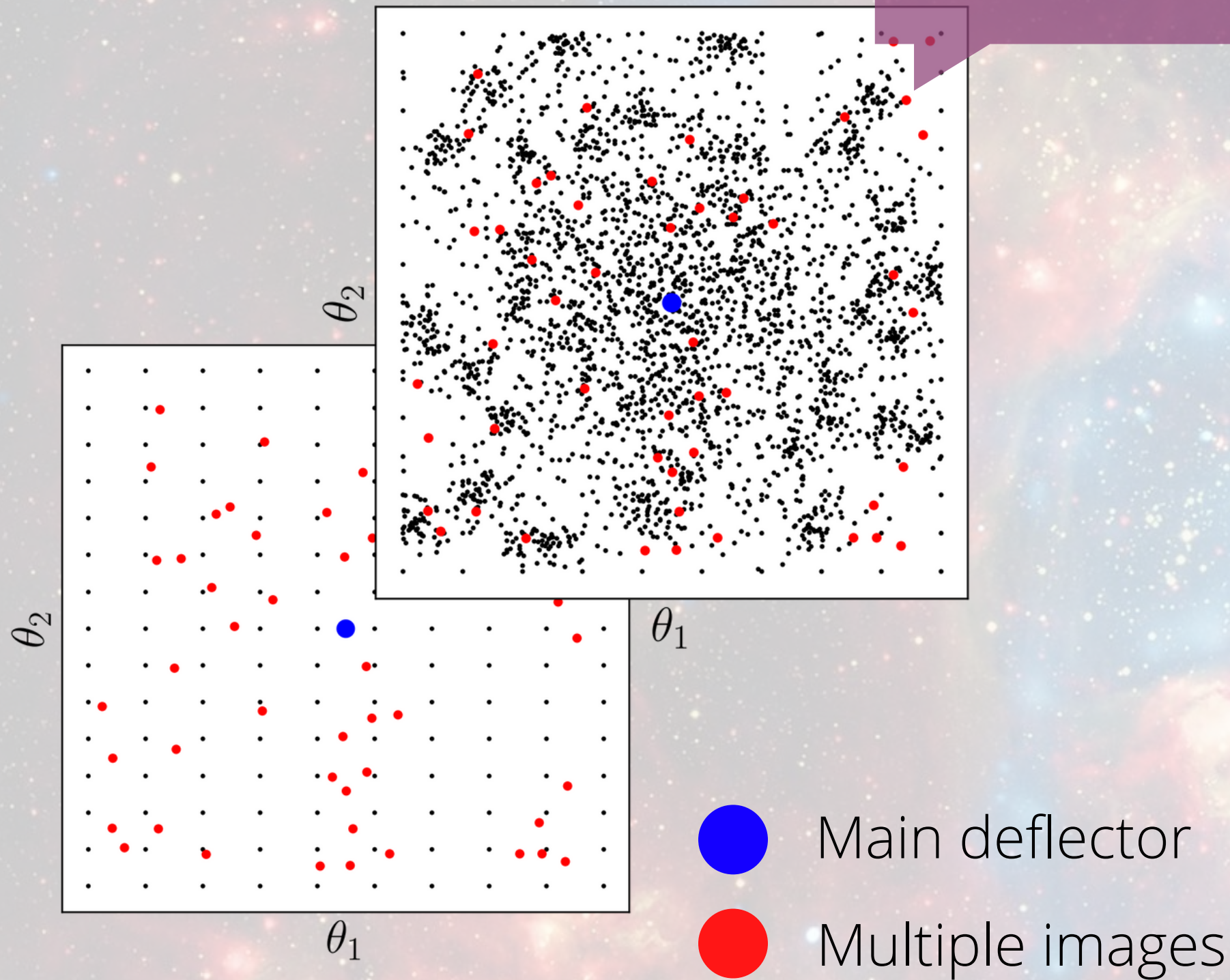


-  Main deflector
-  Multiple images



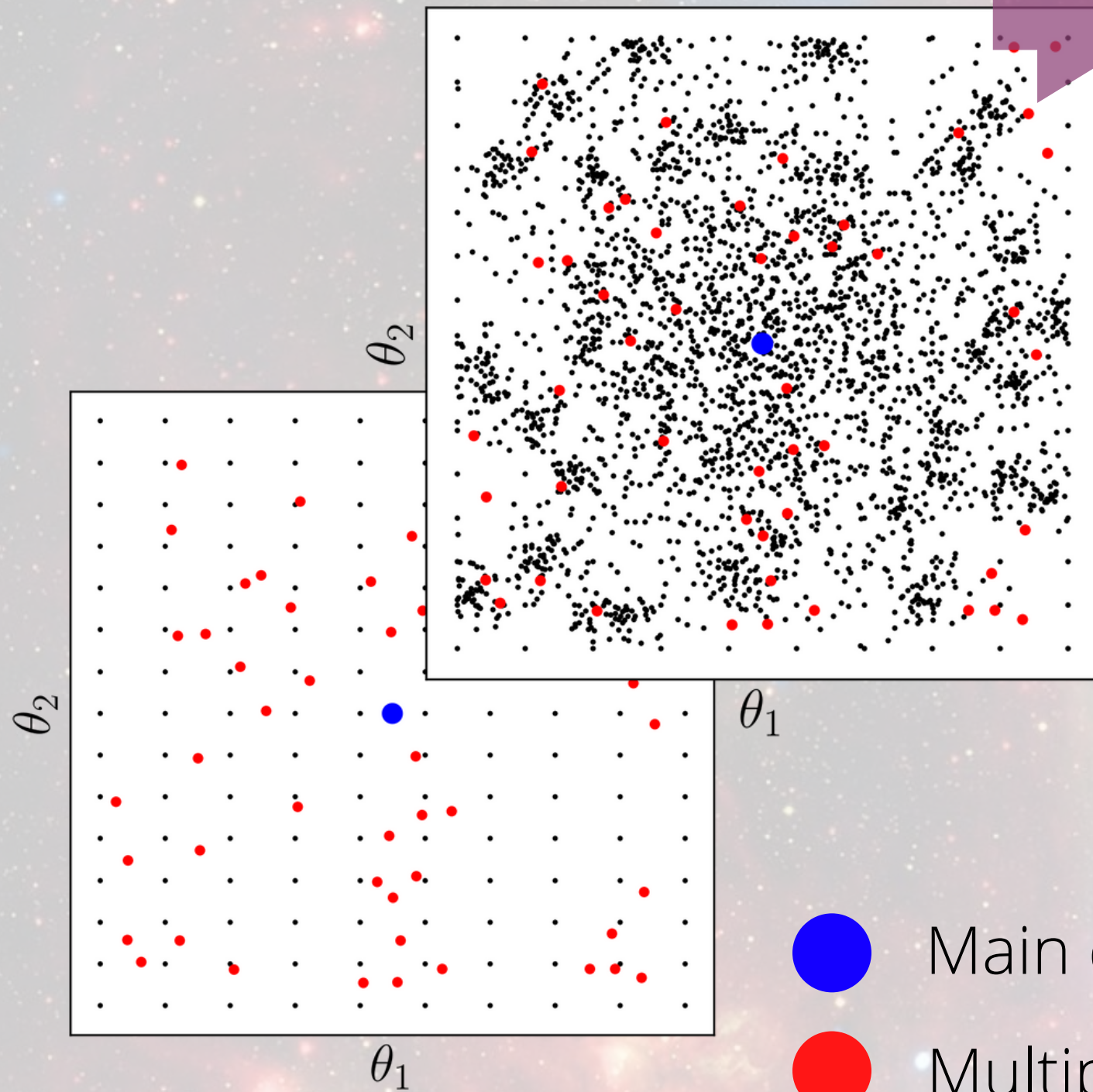
GRID

N nodes in total



GRID

N nodes in total





For an arbitrary angular position θ_i

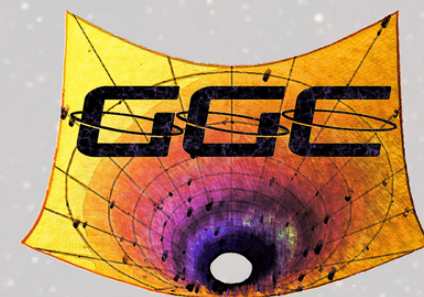
$$\psi_i = \sum_{j=1}^N \mathcal{P}_{ij} \psi_j;$$

$$\kappa_i = \sum_{j=1}^N \mathcal{K}_{ij} \psi_j,$$

$$\gamma_{n,i} = \sum_{j=1}^N \mathcal{G}_{n,ij} \psi_j;$$

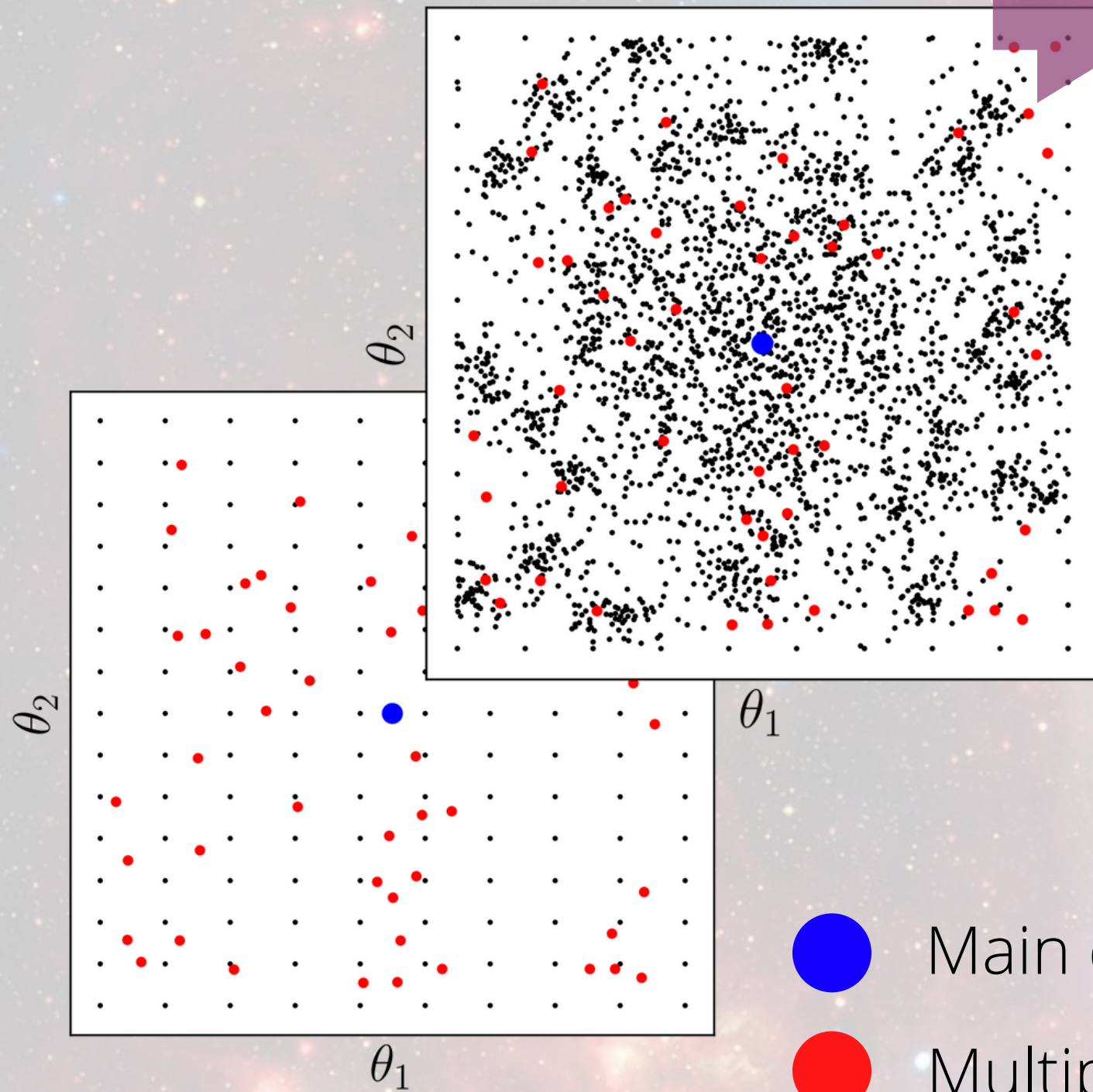
$$\alpha_{n,i} = \sum_{j=1}^N \mathcal{D}_{n,ij} \psi_j \quad (n = 1, 2)$$

-  Main deflector
-  Multiple images



GRID

N nodes in total



- Main deflector
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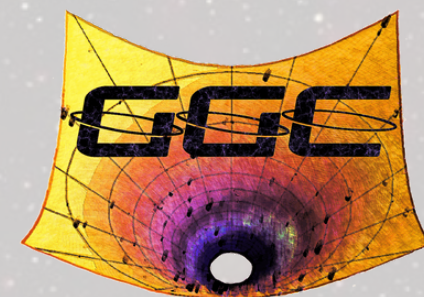
For an arbitrary angular position θ_i

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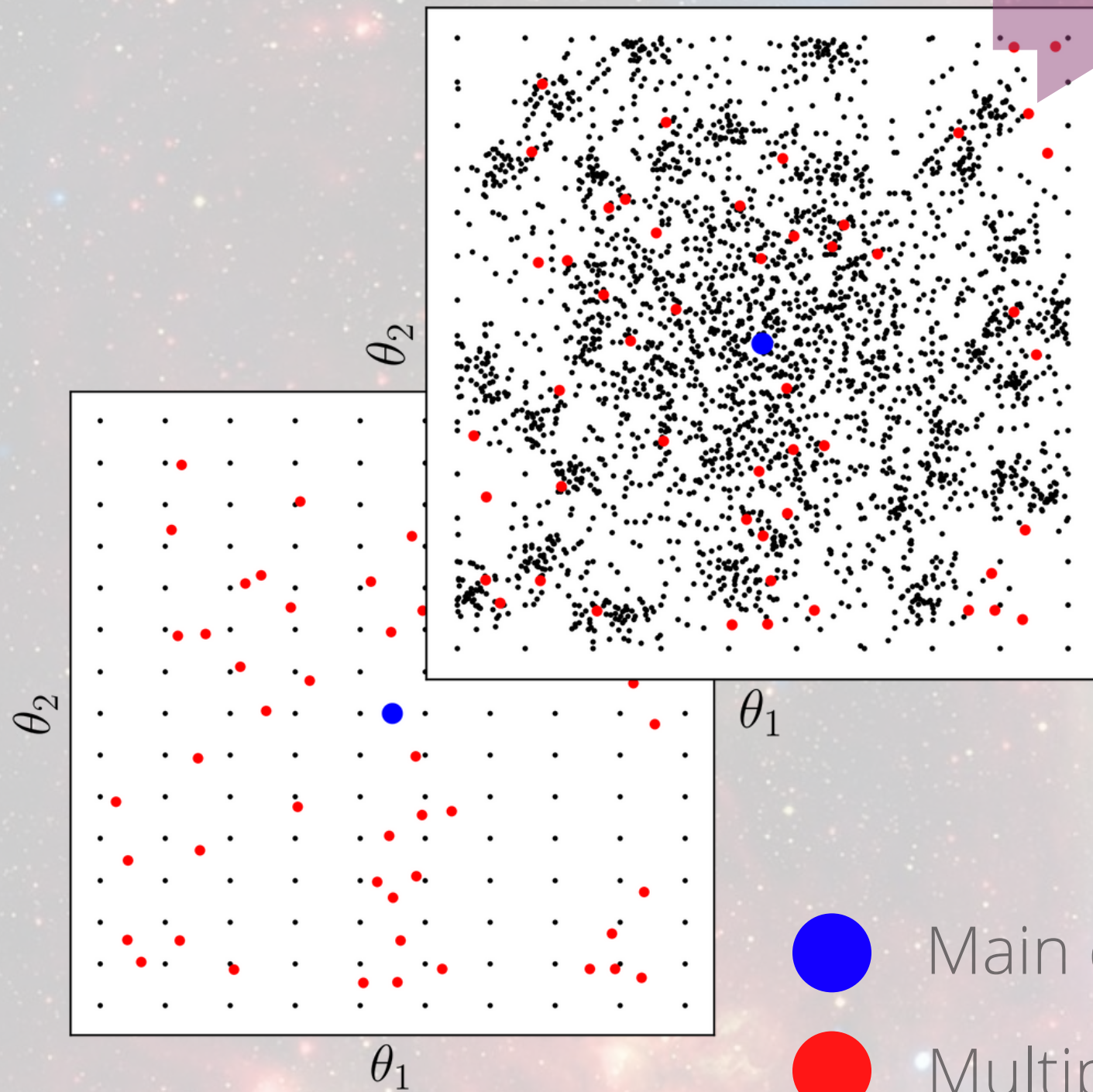
$$\gamma_{n,i} = \sum_{j=1}^N \mathcal{G}_{n,ij} \psi_j; \quad \alpha_{n,i} = \sum_{j=1}^N \mathcal{D}_{n,ij} \psi_j \quad (n = 1, 2)$$

Generalized
Finite
Differences



GRID

N nodes in total



- Main deflector
- Multiple images

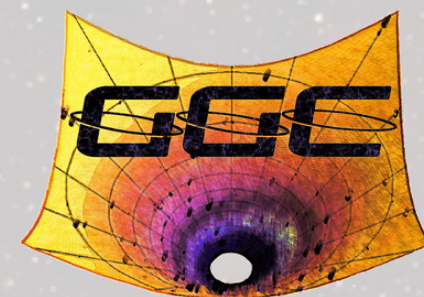
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$$\psi_i = \sum_{j=1}^N \mathcal{P}_{ij} \psi_j, \quad \kappa_i = \sum_{j=1}^N \kappa_{ij} \psi_j,$$

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Free-Form reconstruction!

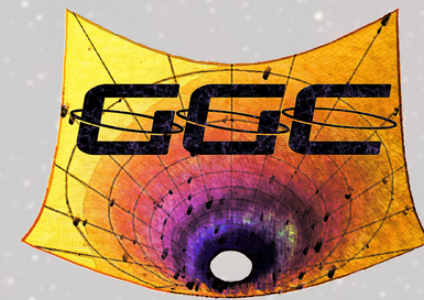
Generalized Finite Differences



ALGORITHM



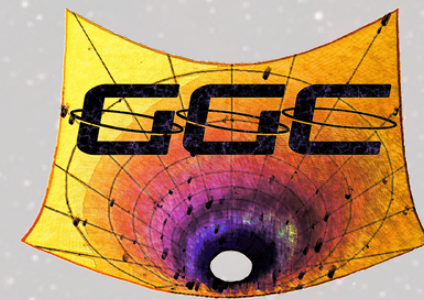
- 1: **Input:** Initial conditions and the initial guess $\psi^{(0)}$.
- 2: Compute $\kappa^{(0)}$, $\gamma_1^{(0)}$ and $\gamma_2^{(0)}$ from $\psi^{(0)}$.
- 3: **while** outer_end = False **do**
- 4: Compute the nodes weight.
- 5: **while** inner_end = False **do**
- 6: Compute the system of linear equations needed and solve it $\rightarrow \psi^{(n)}$.
- 7: Smoothing of $\psi^{(n)}$ \rightarrow New $\psi^{(n)}$ (optional).
- 8: Compute $\kappa^{(n)}$ from $\psi^{(n)}$.
- 9: **if** $|\kappa_j^{(n)} - \kappa_j^{(n-1)}| \leq \text{tolerance}$ **then**
- 10: inner_end = True.
- 11: Compute χ_s^2 .
- 12: **if** $\chi_s^2 \leq q(2N_{\text{img}})$ **then**
- 13: outer_end = True.
- 14: **if** outer_end = False **then**
- 15: Apply grid refinement.
- 16: Compute $\psi^{(m)}$, $\kappa^{(m)}$, $\gamma_1^{(m)}$ and $\gamma_2^{(m)}$ on the new grid.
- 17: **return** ψ



ALGORITHM

Outer level

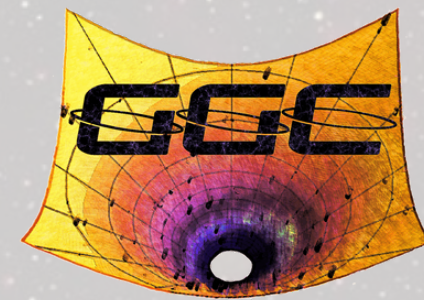
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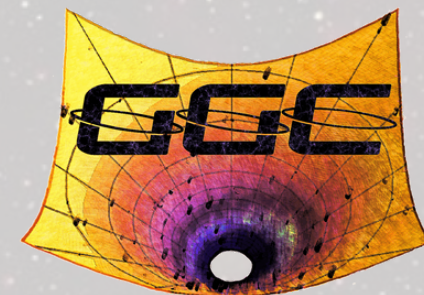


ALGORITHM

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16:     Compute  $\psi^{(m)}$ ,  $\kappa^{(m)}$ ,  $\gamma_1^{(m)}$  and  $\gamma_2^{(m)}$  on the new grid.
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```

Outer level

Inner level

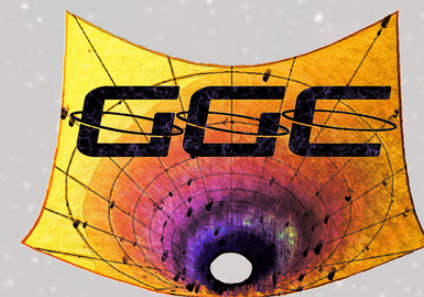


ALGORITHM

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Outer level

Inner level



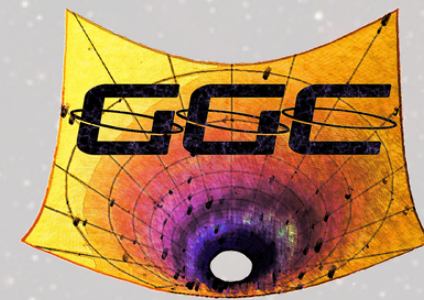
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Outer level

Inner level

$$\chi^2(\psi_j) := \underbrace{\chi_s^2(\psi_j)}_{\text{Strong lensing}} + \underbrace{\chi_w^2(\psi_j)}_{\text{Weak lensing}} + \underbrace{\chi_{\kappa(R)}^2(\psi_j) + \chi_{\gamma(R)}^2(\psi_j)}_{\text{Regularization}}$$



ALGORITHM

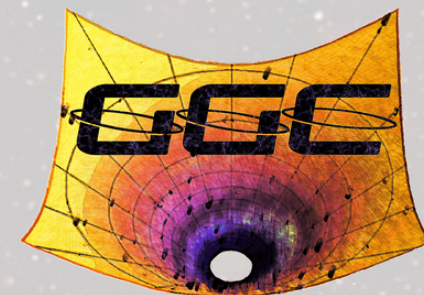
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$$\frac{\partial \chi^2(\psi_j)}{\partial \psi_k} = 0$$



ALGORITHM

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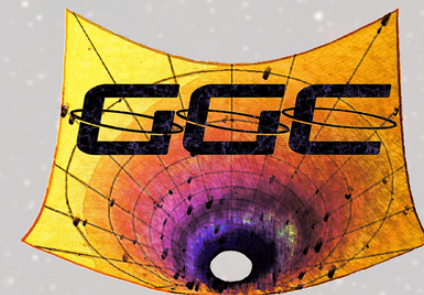
Outer level

Inner level

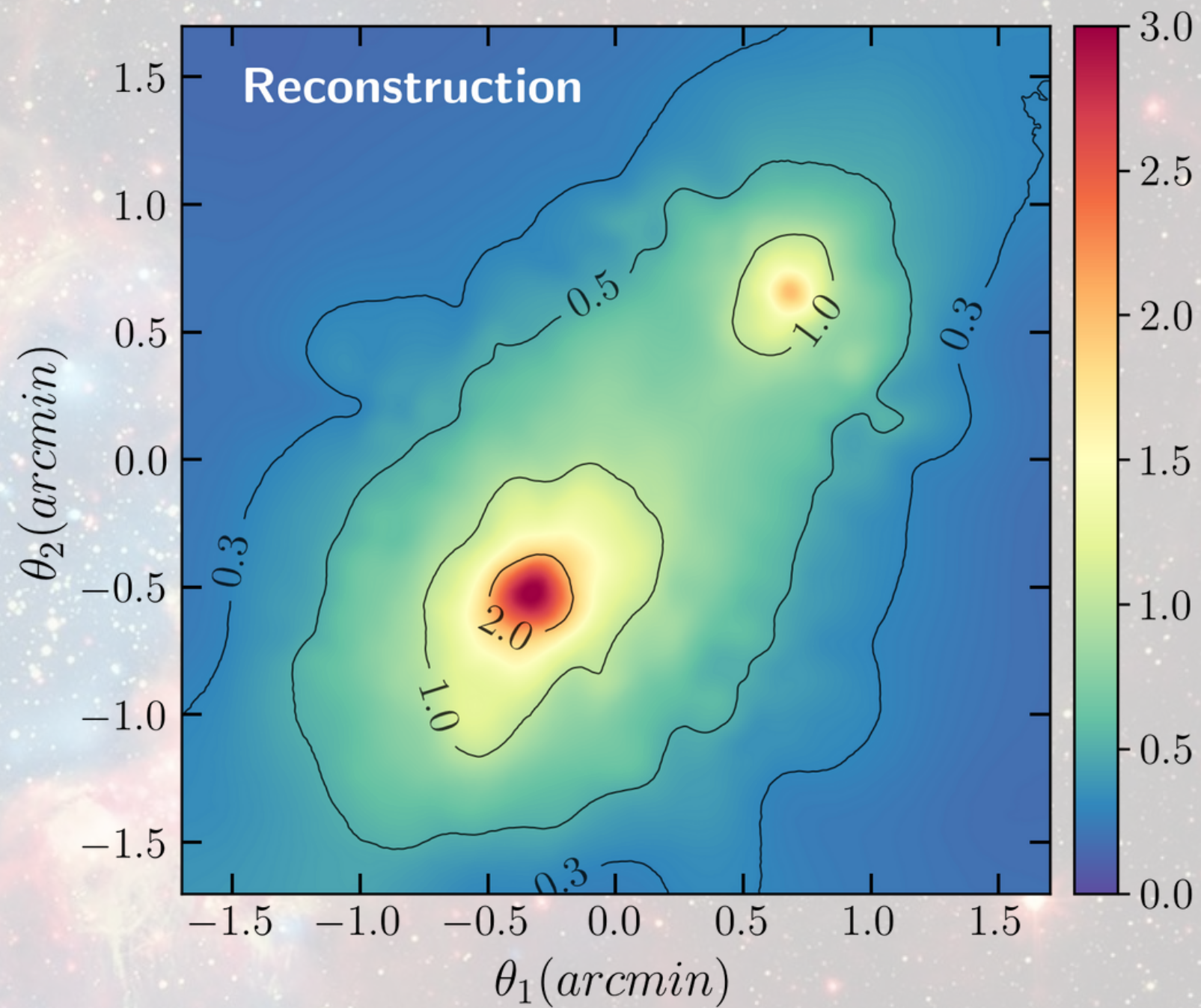
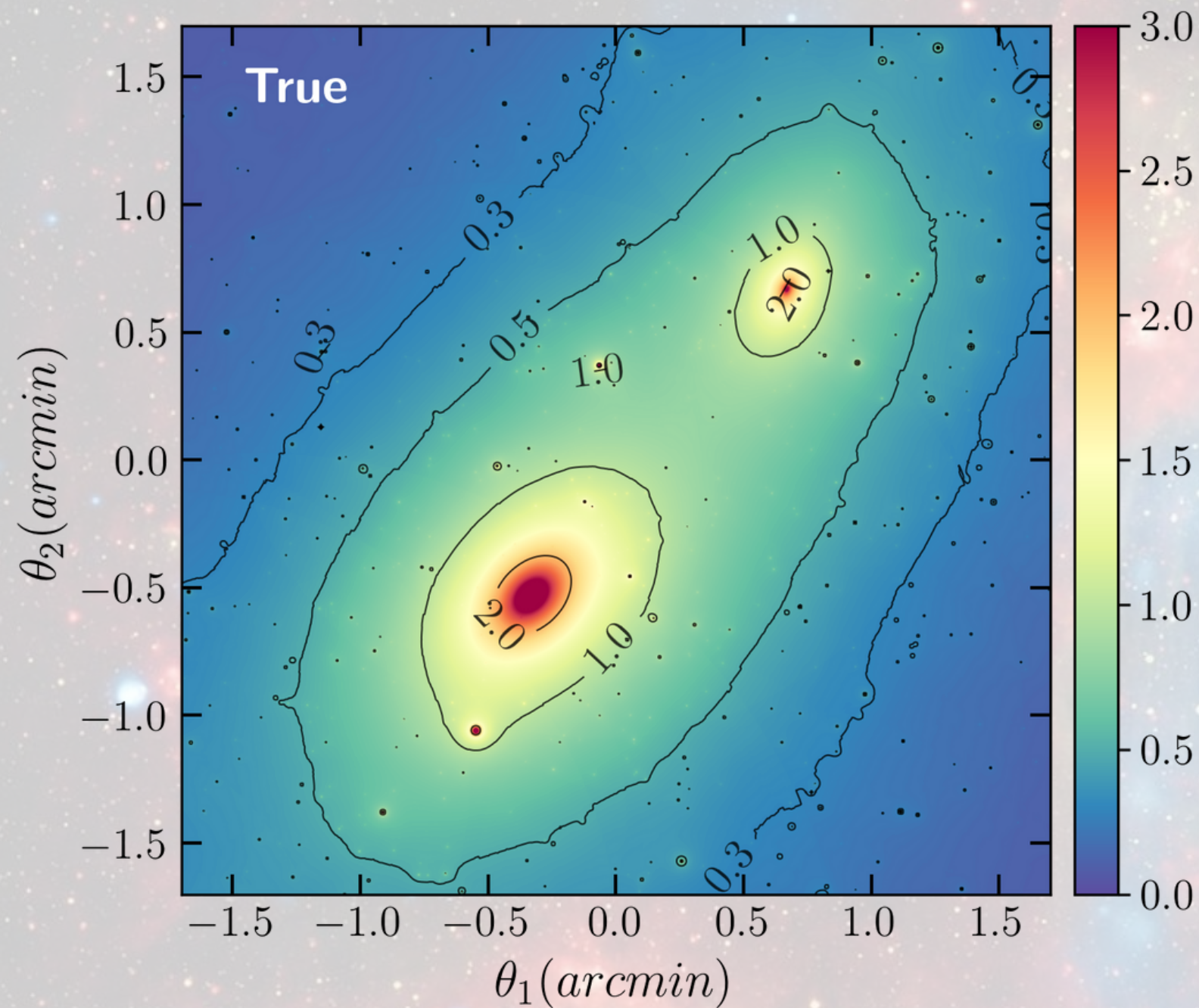
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$$\frac{\partial \chi^2(\psi_j)}{\partial \psi_k} = 0 \quad \rightarrow \quad \sum_{j=1}^N \mathcal{W}_{kj} \psi_j = \mathcal{V}_k$$

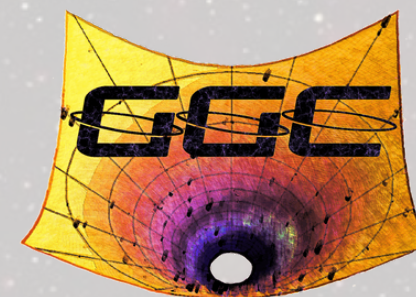
Leads to a system of linear equations!



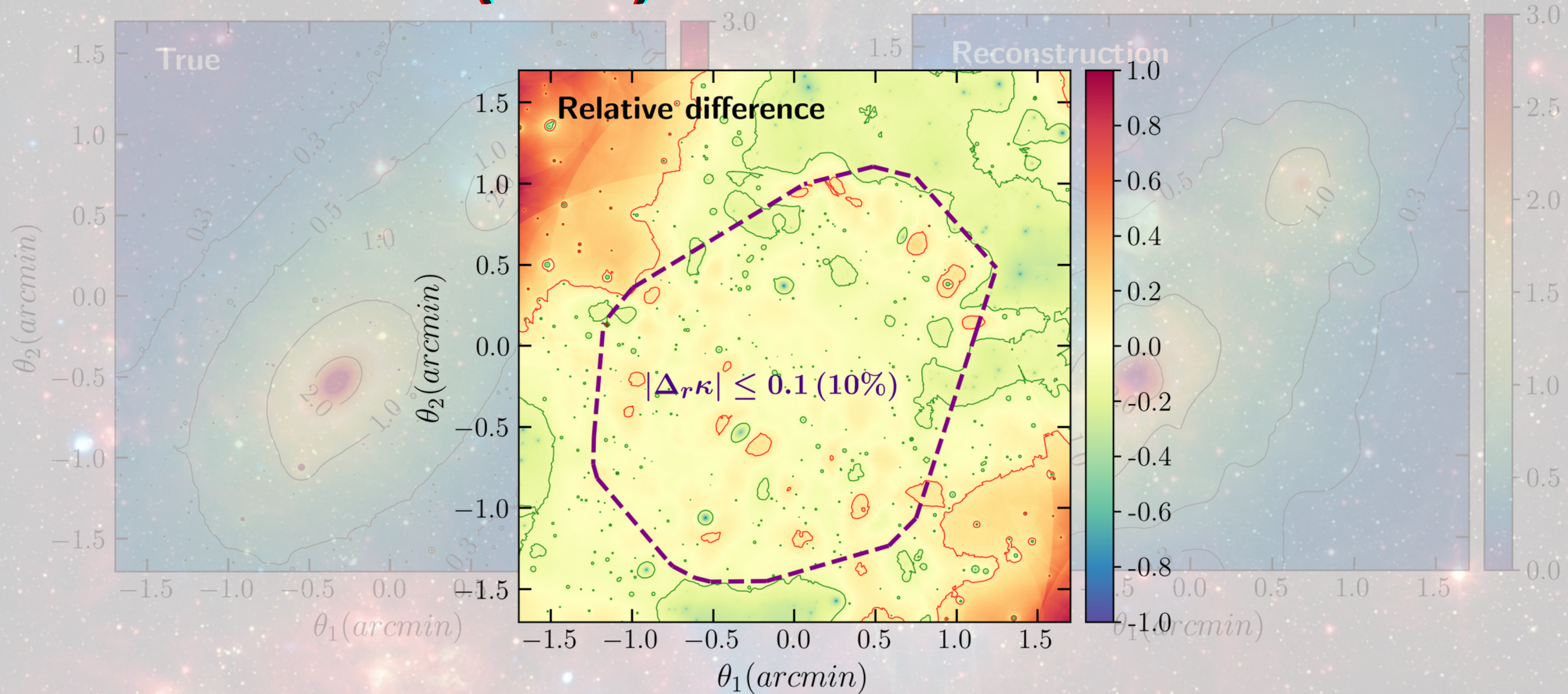
RECONSTRUCTION (ARES)



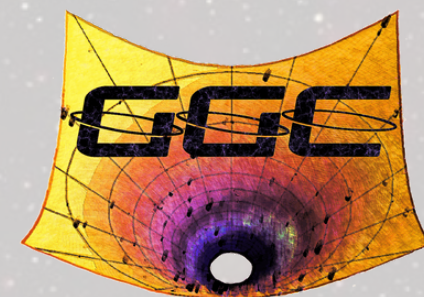
Convergence maps (κ) for a source with $z = 9$



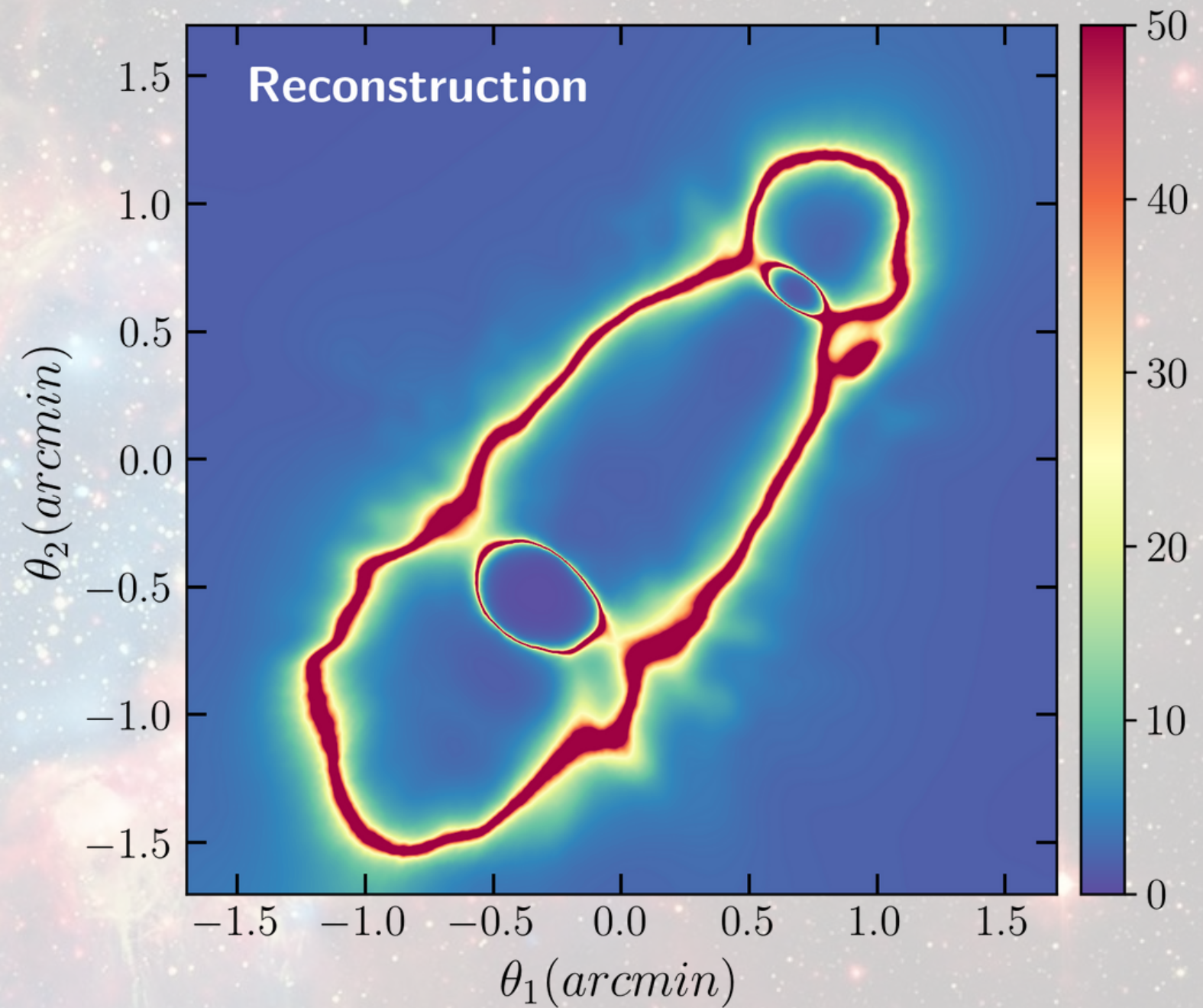
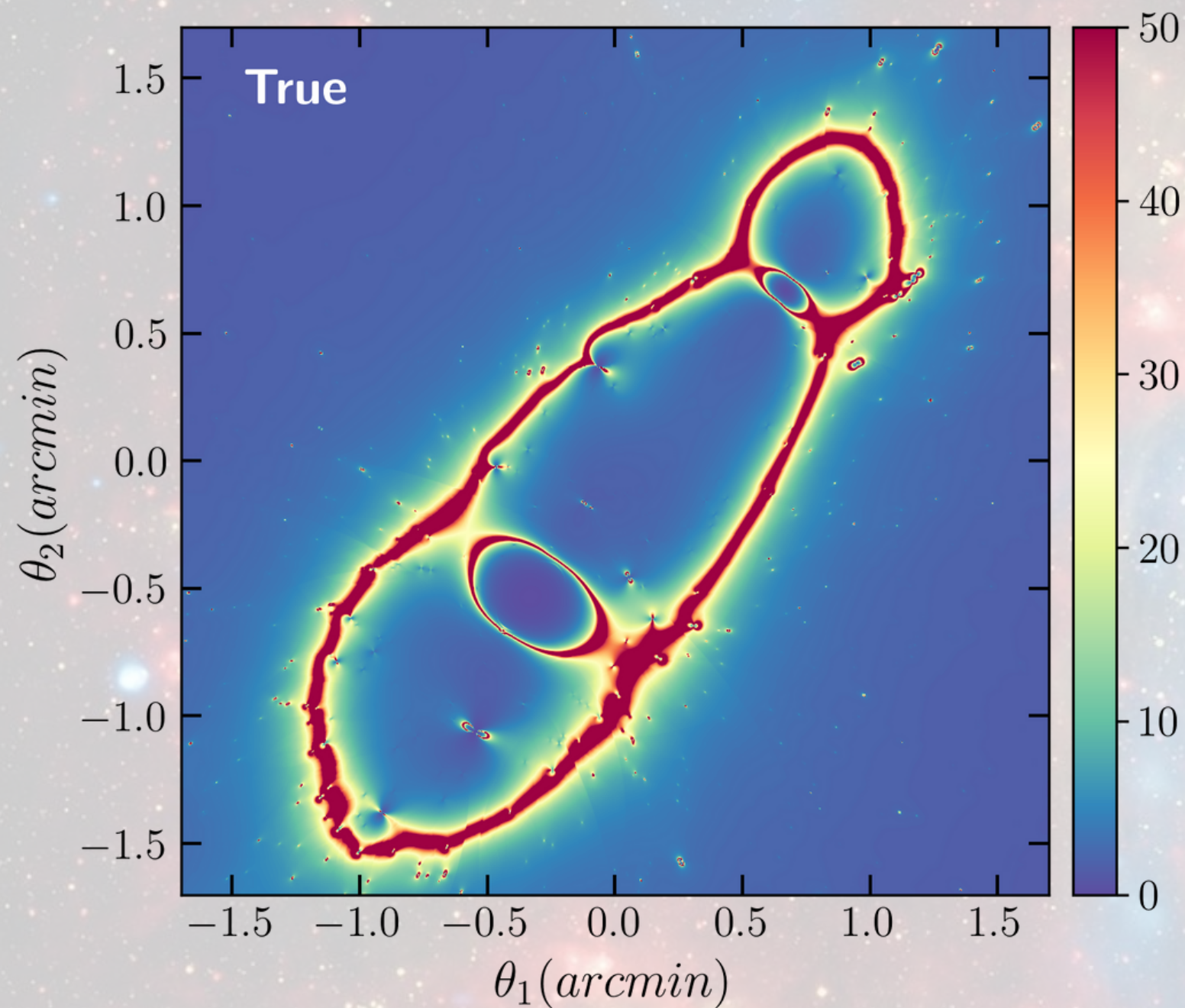
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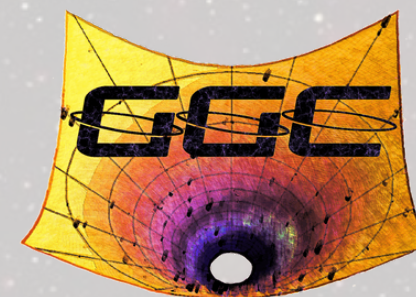
Convergence maps (κ) for a source with $z = 9$



RECONSTRUCTION (ARES)



Magnification maps ($|\mu|$) for a source with $z = 9$



RECONSTRUCTION (ARES)

