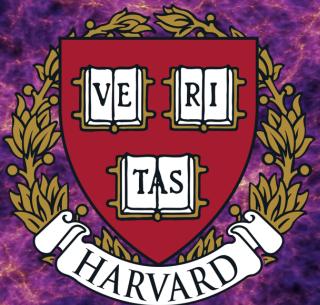


Going Beyond the Galaxy Power Spectrum: an Analysis of BOSS data with Wavelet Scattering Transforms

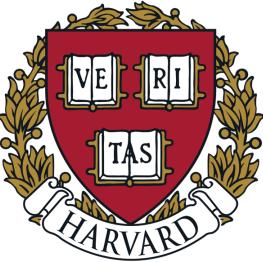


Georgios Valogiannis
Harvard University

Cosmology from Home
Parallel Talk
July 2022

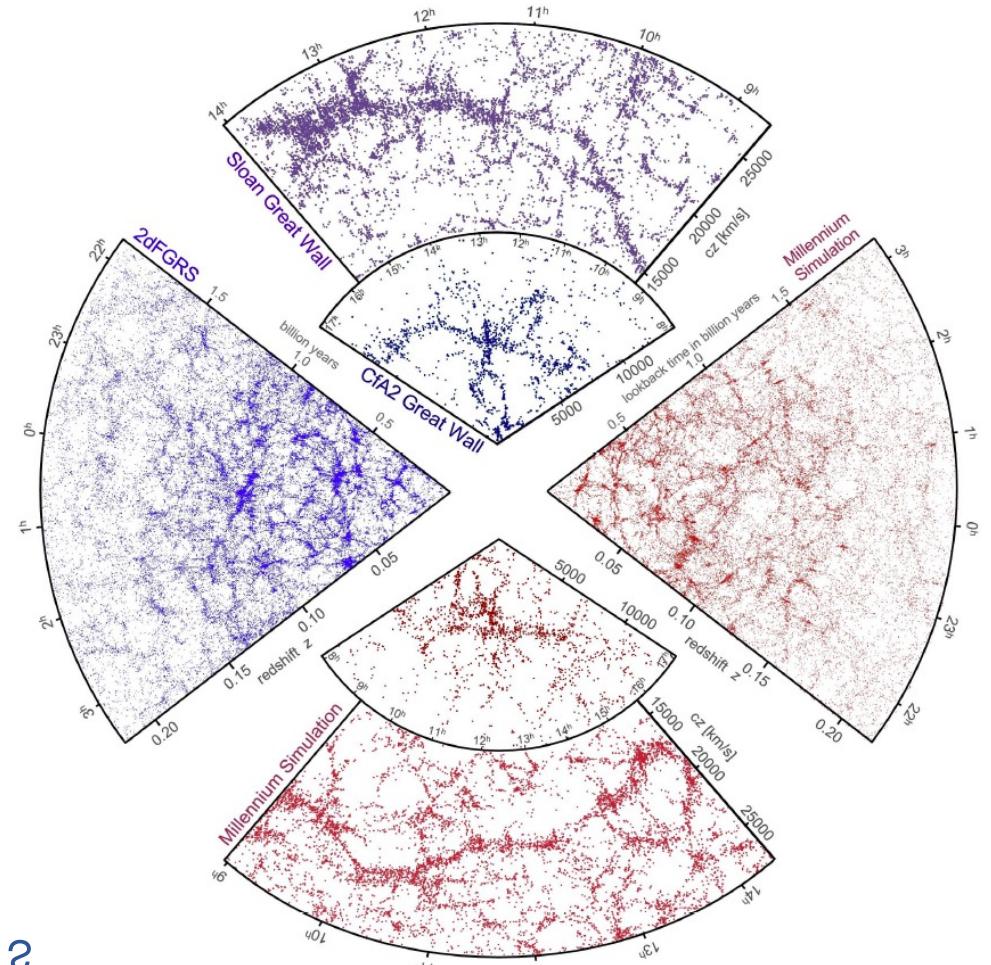
Background from
Millennium Simulation, 2005

Based on
arXiv: 2204.13717 & 2108.07821
in collaboration with Cora Dvorkin

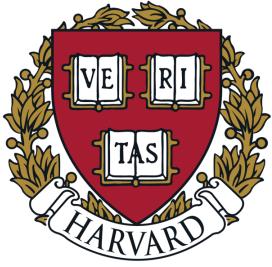


Challenges in the era of precision cosmology

- Large-Scale Structure (LSS) of the universe a powerful probe of *fundamental physics*
 - Dark energy
 - Dark matter
 - Massive neutrinos
 - Gravity
- Will soon be mapped precisely by:
 - Dark Energy Scientific Instrument (DESI)
 - V. Rubin Observatory LSST
 - Euclid
 - Nancy Grace Roman Space Telescope
 - SPHEREx
 - + Synergies with CMB
- How do we optimally extract information from the LSS??

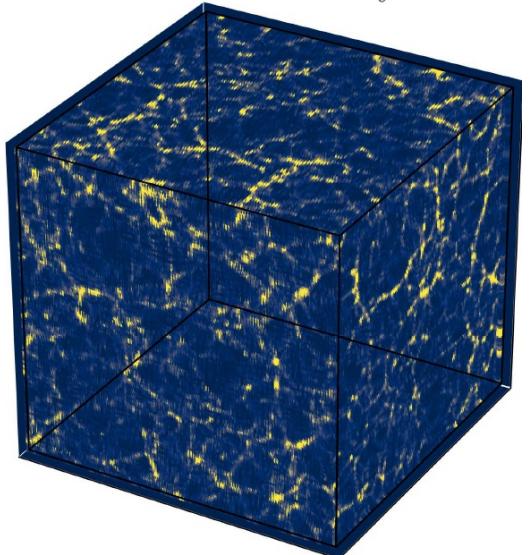


V. Springel et al. (2006)

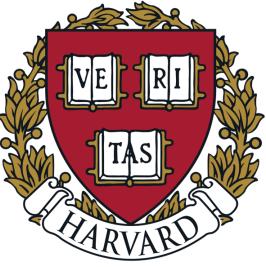


The quest for an ideal estimator

- Attempts to describe the information encoded in the 3D cosmic density field

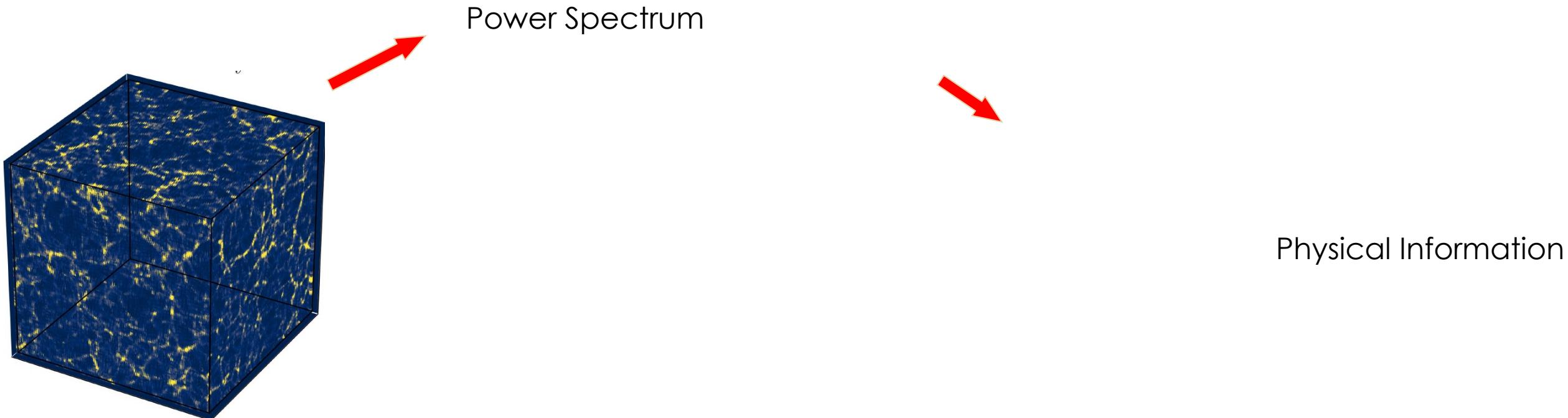


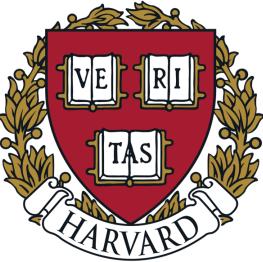
F. Villaescusa-
Navaro et al. (2019)



The quest for an ideal estimator

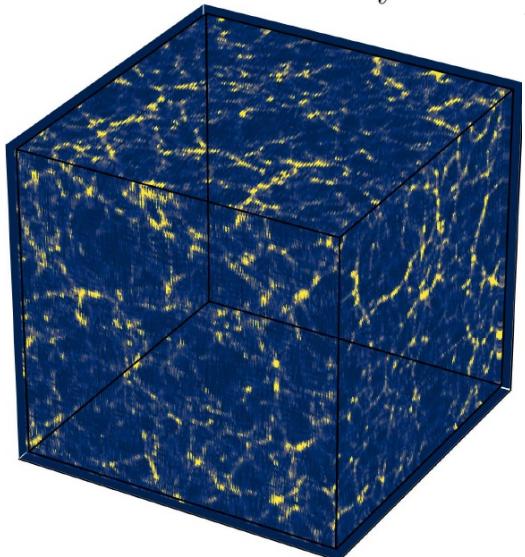
- Attempts to describe the information encoded in the 3D cosmic density field



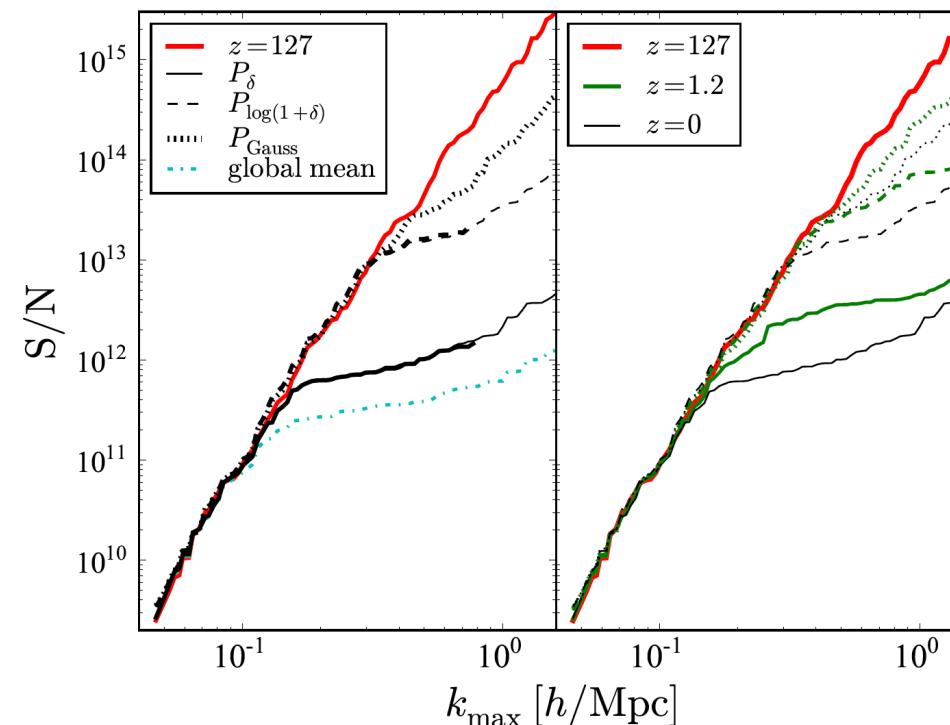


The quest for an ideal estimator

- Attempts to describe the information encoded in the 3D cosmic density field



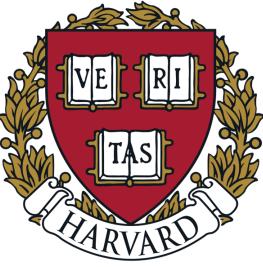
Power Spectrum (Incomplete)



M. Neyrinck et al. (2009)

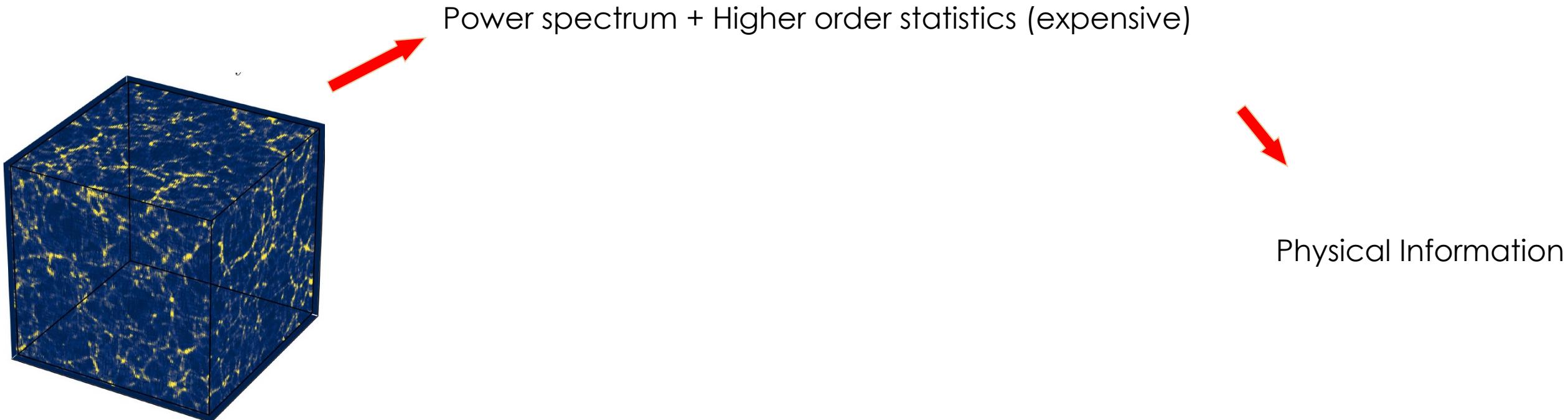
Power Spectrum information
saturates in nonlinear regime.
Inadequate!
(Carron 2011,2012)

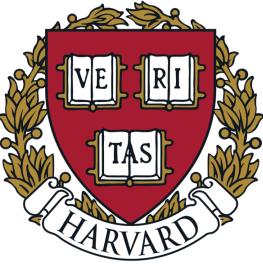
F. Villaescusa-
Navaro et al. (2019)



The quest for an ideal estimator

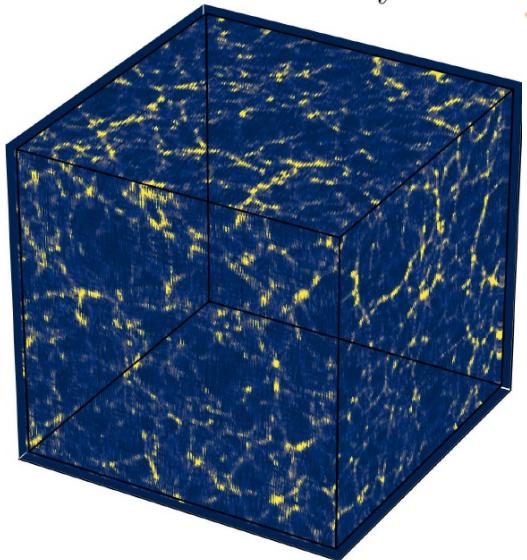
- Attempts to describe the information encoded in the 3D cosmic density field





The quest for an ideal estimator

- Attempts to describe the information encoded in the 3D cosmic density field

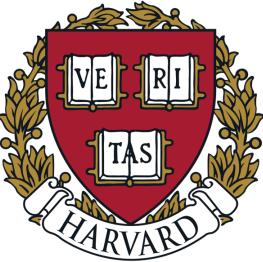


Power spectrum + Higher order statistics

Physical Information

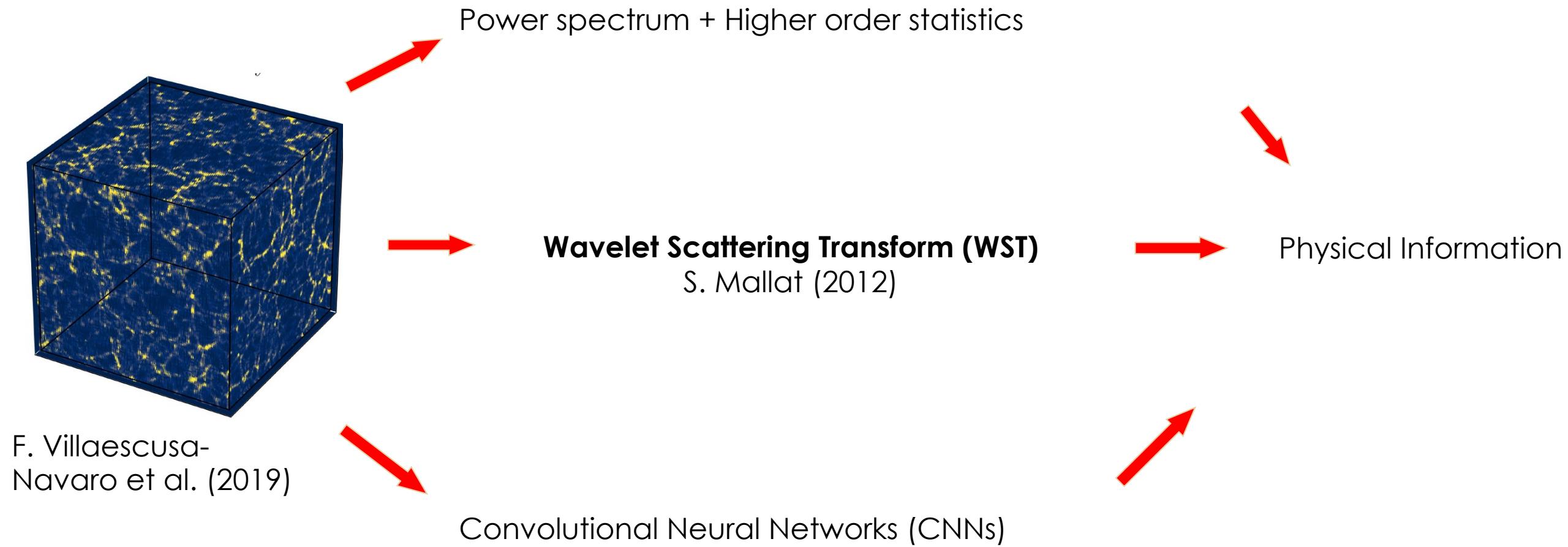
F. Villaescusa-
Navaro et al. (2019)

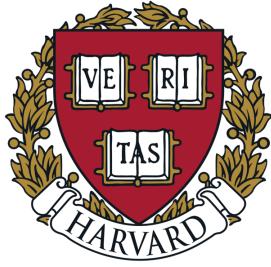
Convolutional Neural Networks (CNNs)
(Training, interpretability)



The quest for an ideal estimator

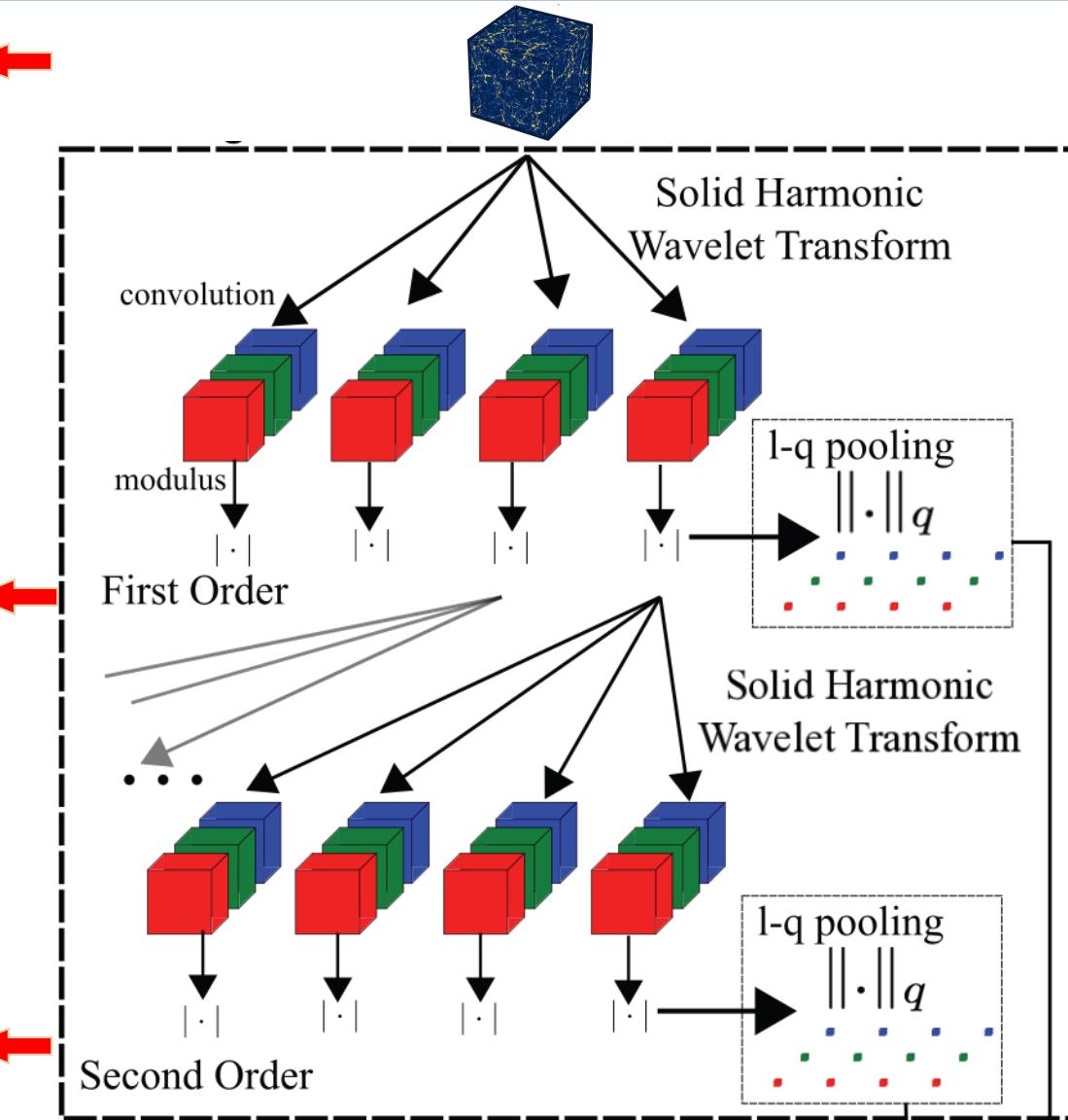
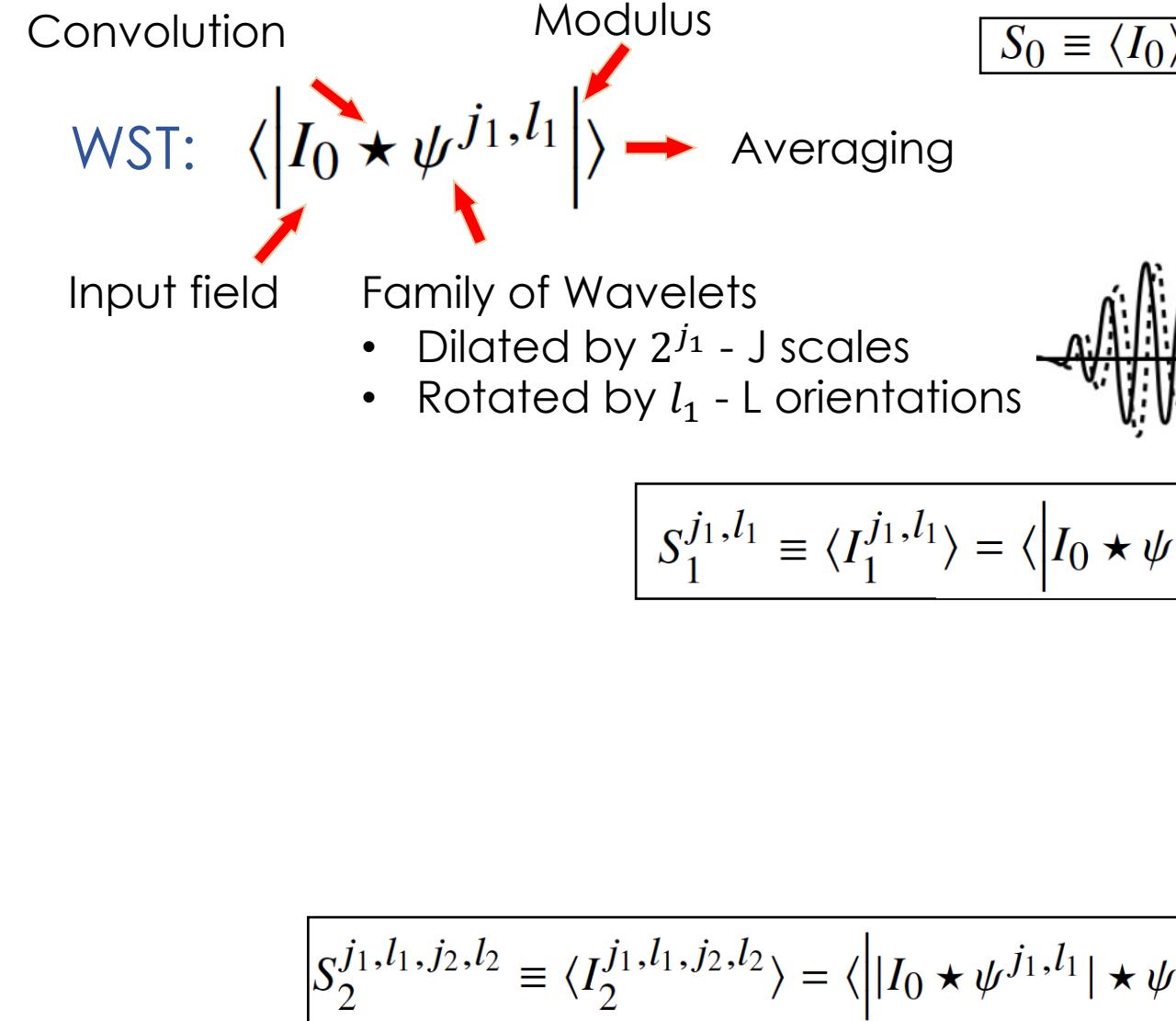
- Attempts to describe the information encoded in the 3D cosmic density field

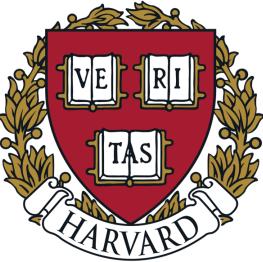




The Wavelet Scattering Transform (WST)

"Scattering Network" image by G. Exarchakis (2018)

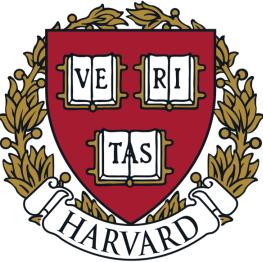




The Wavelet Scattering Transform (WST)

Physical interpretation of WST coefficients

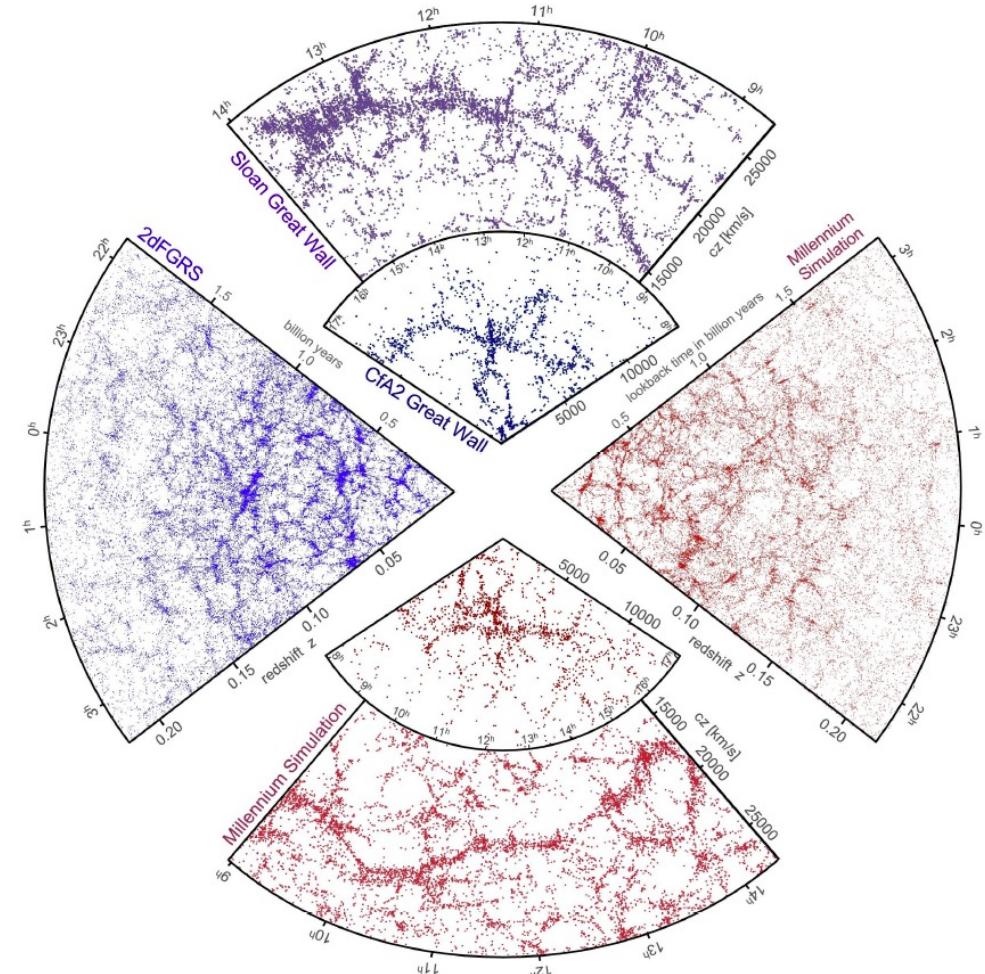
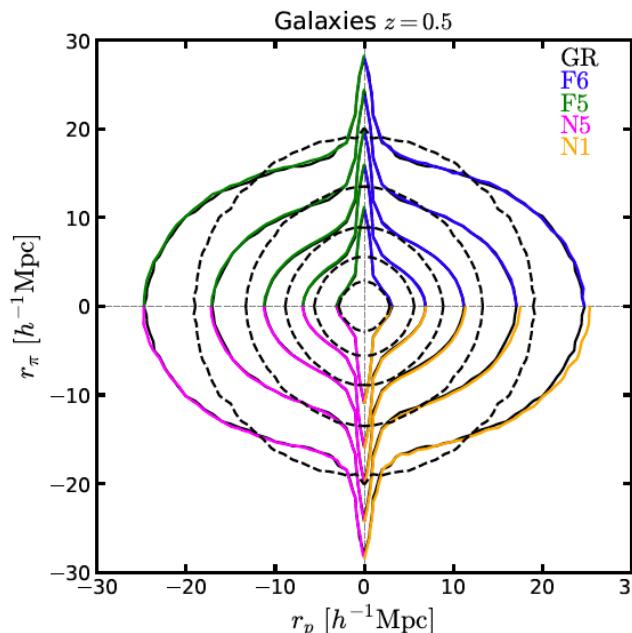
- $S_0 \equiv \langle I_0 \rangle$: Mean field
- $S_1^{j_1, l_1} = \langle |I_0 \star \psi^{j_1, l_1}| \rangle$: $\sim P(k)$. In fact, $P(k) \rightarrow \langle |I \star e^{-ikx}|^2 \rangle$
- $S_2^{j_1, l_1, j_2, l_2} = \langle |I_0 \star \psi^{j_1, l_1} \star \psi^{j_2, l_2}| \rangle$: Non-Gaussian information (up to $2^2 = 4$ pcf , for n=2)
- Basis $S_0 + S_1 + S_2$ reflects clustering properties of target field $I_0(x)$
- Retaining all *desirable* properties of regular $P(k)$ ✓ Mallat (2012)
 - +
- Compactness ✓ (Anden & Mallat, 2011, 2014, Bruna & Mallat, 2013)
- Robustness/Stability ✓ (Carron 2011, 2012, Cheng & Menard 2021b)
- A CNN with fixed weights, but interpretable! (Bruna & Mallat 2013)
 - Performance on par with a CNN in WL applications! (Cheng et al. 2020b, Cheng & Menard 2021a)
- WST exceeds performance of regular & marked $P(k)$ in 3D LSS studies (**Valogiannis & Dvorkin 2021**)

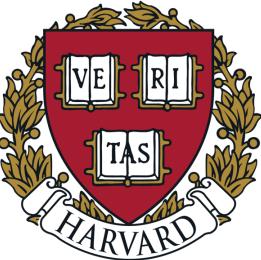


Realistic galaxy survey data

However

- LSS surveys observe galaxies:
 - Biased tracers of dark matter field
 - Redshift-Space Distortions (RSD)
 - Systematics (Geometry, fiber collisions, etc..)
 - Lightcone
 - etc





First WST application on BOSS

- First WST application on 3D redshift-space galaxy density field! (Valogiannis & Dvorkin 2022)
 - Working with BOSS CMASS DR12 sample at $0.46 < z < 0.60$
 - Northern + Southern Galactic Cap
- For survey data, fundamental quantity of interest is the *FKP field* (Feldman, Kaiser, Peacock et al., 1994) :

$$F(\mathbf{r}) = \frac{w_{\text{FKP}}(\mathbf{r})}{I_2^{1/2}} [w_c(\mathbf{r})n_g(\mathbf{r}) - \alpha_r n_s(\mathbf{r})]$$

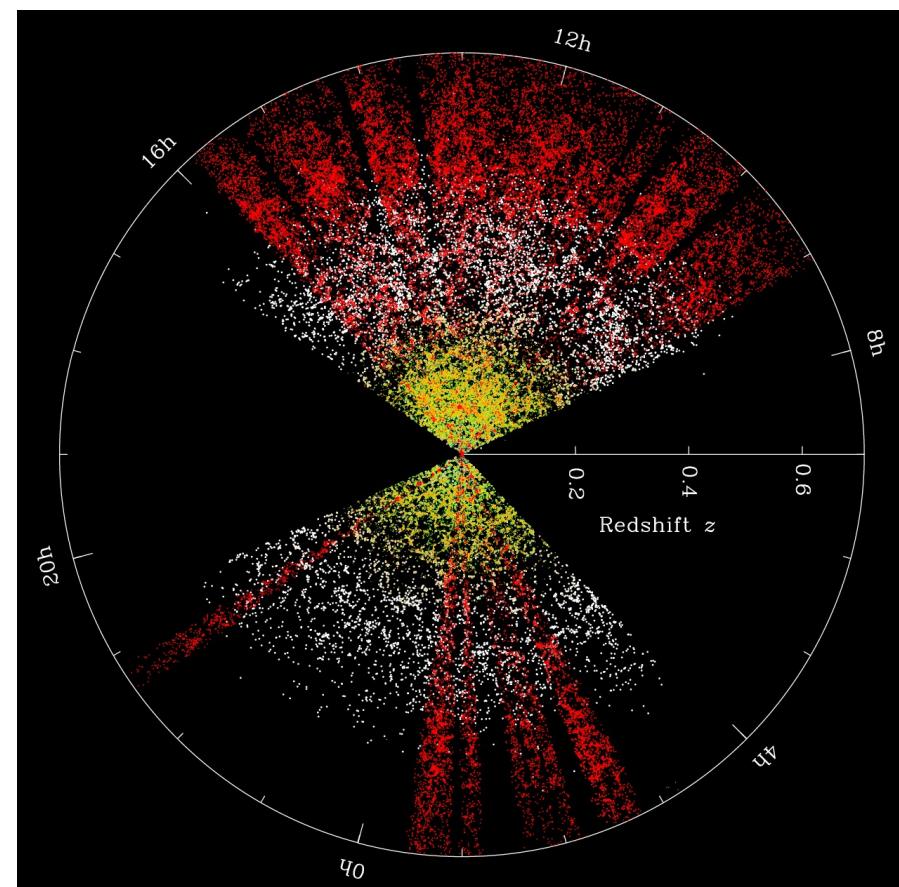
Galaxies Randoms

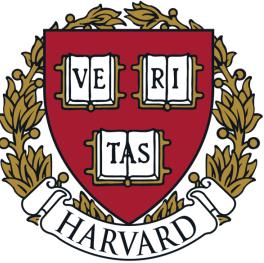
- Systematic + FKP weights

$$w_c(\mathbf{r}) = (w_{\text{rf}}(\mathbf{r}) + w_{\text{fc}}(\mathbf{r}) - 1.0) w_{\text{sys}}(\mathbf{r})$$

$$w_{\text{FKP}}(\mathbf{r}) = [1 + \bar{n}_g(\mathbf{r})P_0]^{-1}$$

- Serves as input into WST network
 - With $N_{\text{grid}} = 282^3$ and $L_{\text{Box}} = 2820 \text{ Mpc}/h$



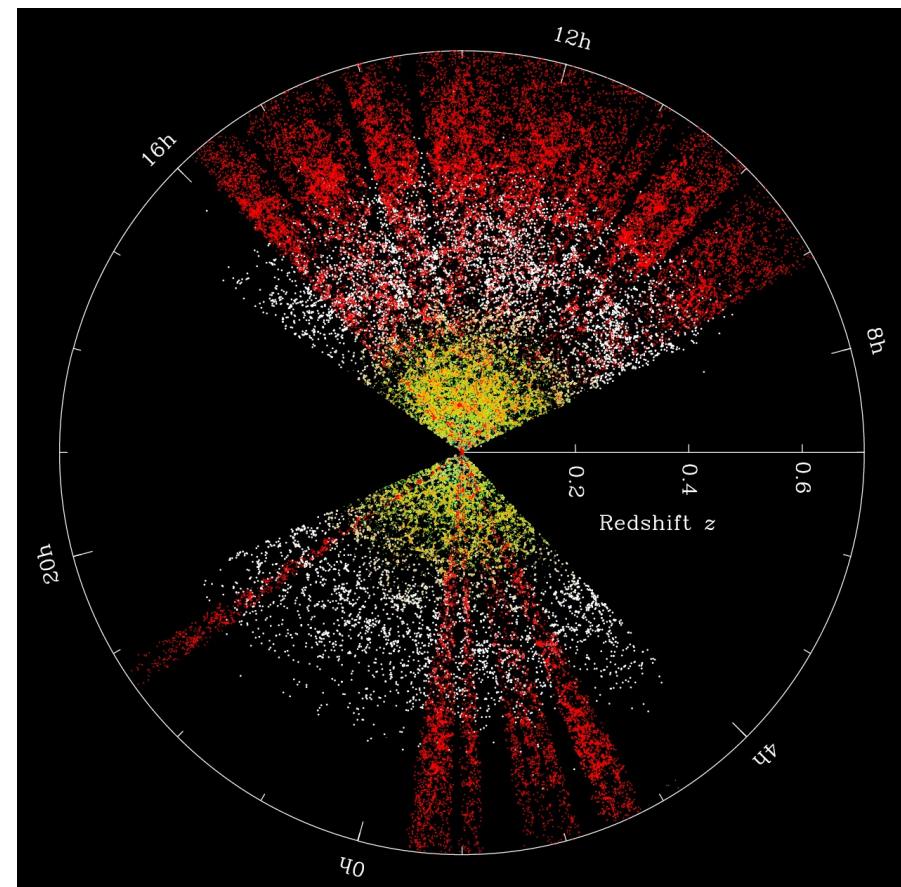


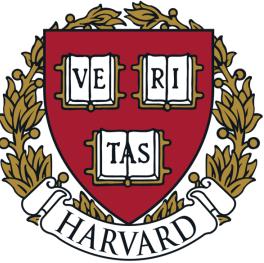
Likelihood analysis

- Data

$$\log \mathcal{L}(\theta | \mathbf{d}) \propto -\frac{1}{2} [\mathbf{X}_d - \mathbf{X}_t(\theta)]^T C^{-1} [\mathbf{X}_d - \mathbf{X}_t(\theta)]$$

- Use vector of WST coefficients as observable
- Extracted from BOSS CMASS FKP field, using J=4 scales and L=4 orientations
- $S_0 + S_1 + S_2 = 76$ WST coefficients
- Also, use galaxy power spectrum multipoles $P_{l=0,2}(k)$ ($k_{max} = 0.25 \text{ Mpc/h}$) as benchmark





Likelihood analysis

- Theory model

$$\log \mathcal{L}(\theta | \mathbf{d}) \propto -\frac{1}{2} [\mathbf{X}_d - \mathbf{X}_t(\theta)]^T C^{-1} [\mathbf{X}_d - \mathbf{X}_t(\theta)]$$

- Capture cosmological dependence using

Abacus Summit simulations (Maksimova et al. 2021, Garrison et al. 2019&2021)

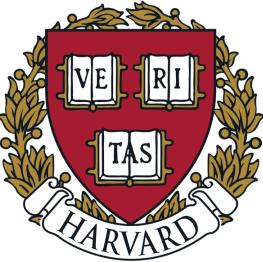
HOD tuned to BOSS CMASS at $0.46 < z < 0.60$ with AbacusHOD (**Yuan et al. 2021**)

Box $L=2000$ Mpc/h, $N_{grid} = 200^3$

- Fiducial cosmology from Planck 2018 $\{\omega_b, \omega_c, n_s, \sigma_8\} = \{0.02237, 0.120, 0.9649, 0.8114\}$
- + Fixed angular size of sound horizon at last scattering. $100\theta_\star = 1.041533$
- + 7 HOD model parameters (vanilla HOD + velocity bias)

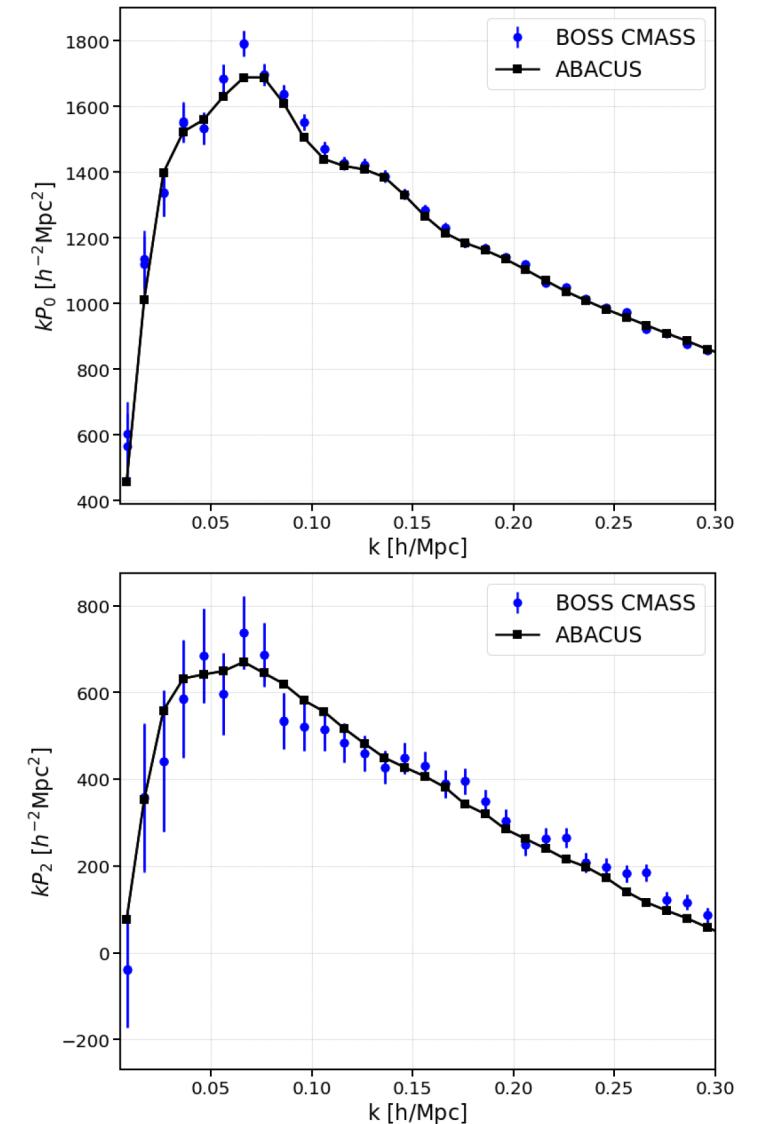
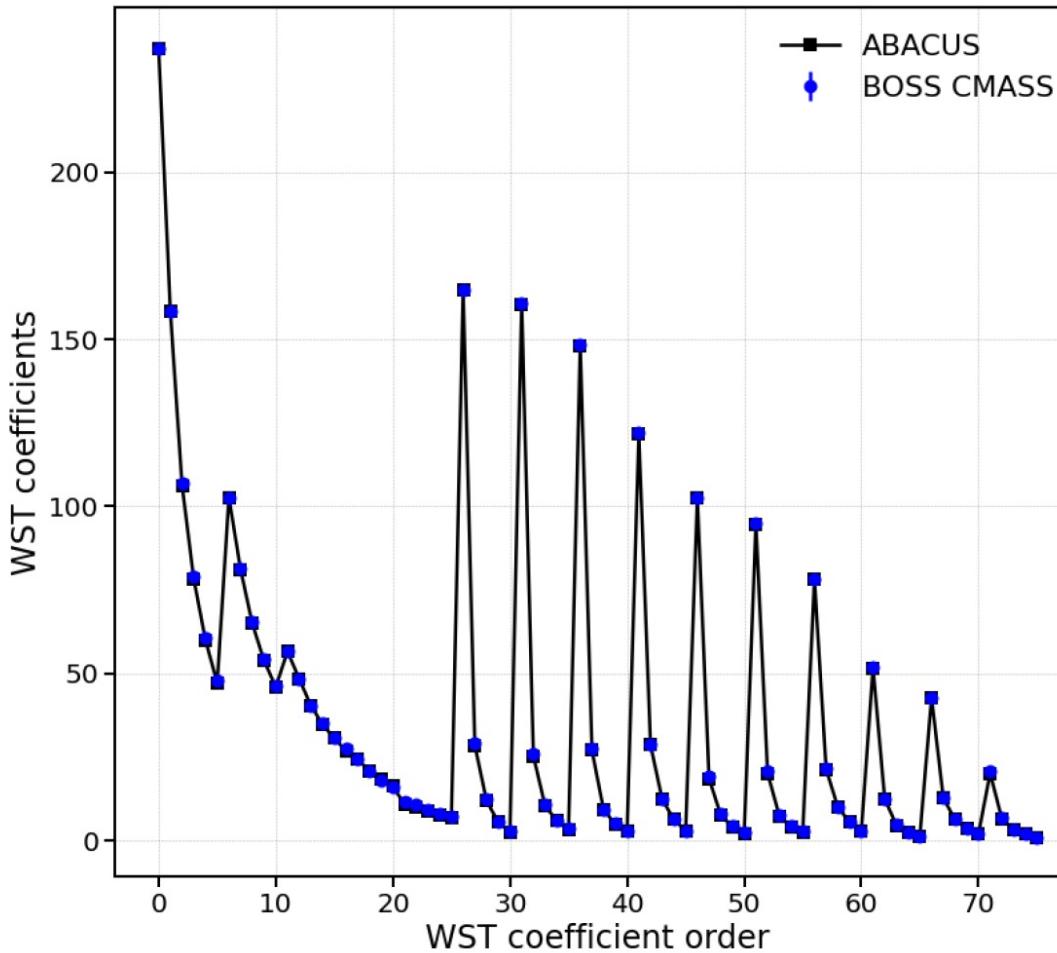
$\{\alpha, \alpha_c, \alpha_s, \kappa, \log M_1, \log M_{cut}, \sigma\} = \{0.9022, 0.2499, 1.1807, 0.3288, 14.313, 12.8881, 0.02084\}$

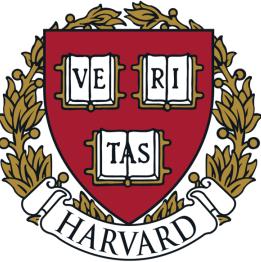
- We cut Abacus cubic boxes into actual CMASS geometry
 - Using ‘make survey’ (White et al., 2013)



Likelihood analysis

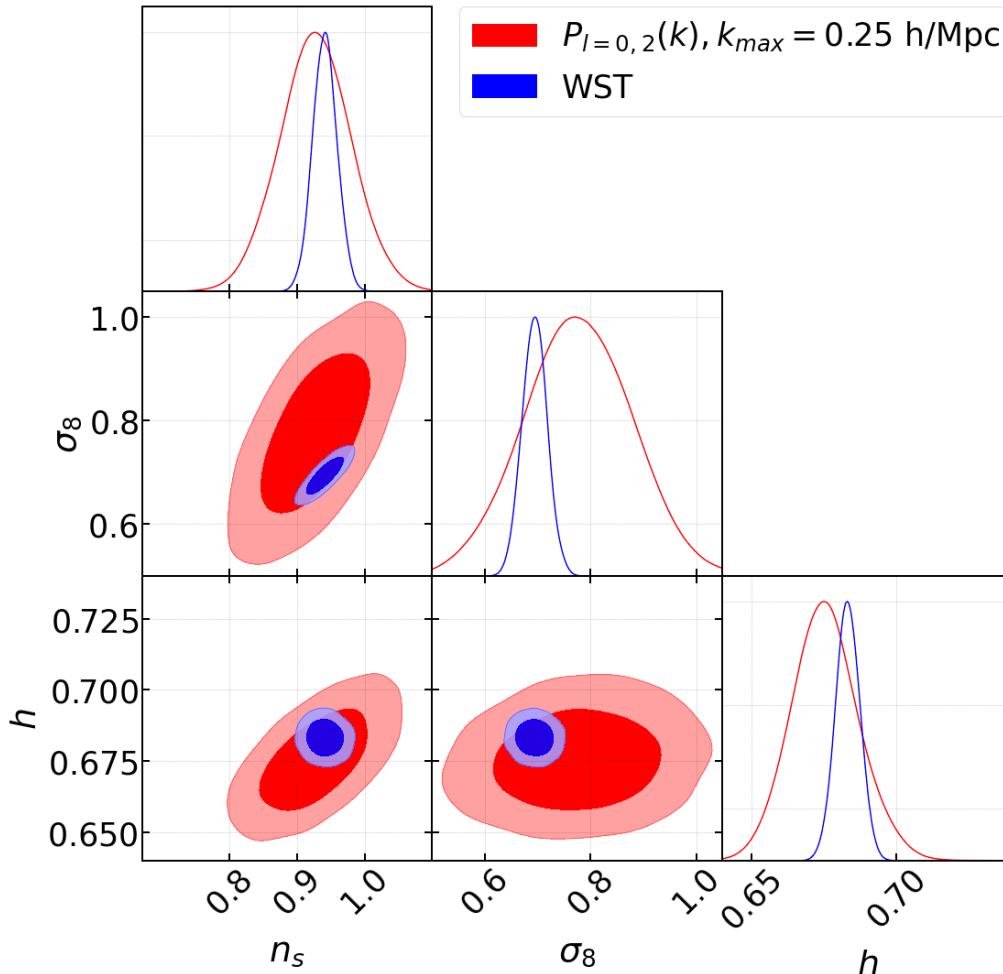
- Fiducial cosmology predictions





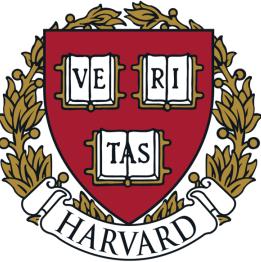
Likelihood analysis

- Likelihood analysis using a BBN prior on ω_b $\omega_b = 0.02268 \pm 0.00038$



	BBN prior on ω_b		unrestricted priors	
	P(k)	WST	P(k)	WST
ω_b	$0.02267^{+0.00045}_{-0.00045}$	$0.02268^{+0.00036}_{-0.00036}$	$0.0217^{+0.0043}_{-0.0043}$	$0.01946^{+0.0008}_{-0.0008}$
ω_c	$0.1223^{+0.0031}_{-0.0028}$	$0.1202^{+0.00013}_{-0.00013}$	$0.1217^{+0.0058}_{-0.0058}$	$0.11672^{+0.001}_{-0.001}$
n_s	$0.928^{+0.075}_{-0.075}$	$0.942^{+0.018}_{-0.018}$	$0.921^{+0.057}_{-0.049}$	$0.959^{+0.019}_{-0.019}$
σ_8	$0.77^{+0.14}_{-0.14}$	$0.695^{+0.024}_{-0.024}$	$0.762^{+0.11}_{-0.094}$	$0.716^{+0.025}_{-0.025}$
h	$0.676^{+0.010}_{-0.012}$	$0.6831^{+0.0042}_{-0.0042}$	$0.668^{+0.024}_{-0.024}$	$0.66^{+0.0055}_{-0.0055}$

- Parameter mean values from WST & P(k) always consistent with each other within 1σ (of the P(k))
- Much **tighter** errors from WST compared to P(k)!
- H_0 determined from WST with 0.6% accuracy!

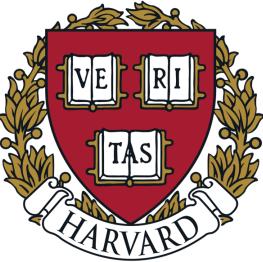


Conclusions

- Wavelet Scattering Transform: a novel statistic that efficiently extracts non-Gaussian information from physical fields. *Ideal middle ground between CNN and traditional estimators*
- *First WST application on actual spectroscopic data (Valogiannis & Dvorkin , arXiv: 2204.13717)*
 - Worked with BOSS CMASS galaxy sample at $0.46 < z < 0.60$
 - Great improvement in the 1σ errors over traditional galaxy $P(k)$ multipoles
 - Much **tighter** errors both using BBN prior on ω_b , and using flat unrestricted priors!
 - 0.6% determination of the Hubble constant!
- *Future improvements (in progress)*
 - Construct full emulator for WST coefficients (Eg. Yuan et al. 2022)
 - Include lightcone effects in galaxy mocks
- *Future applications*
 - Constrain neutrino mass (Eg. as in Valogiannis & Dvorkin, arXiv: 2108.07821, Phys. Rev. D 105, 103534, 2022)
 - Analysis explicitly varying H_0 (rather than derived parameter)
 - Comparison with higher-point function analyses (Eg. Philcox et al., 2021)
 - Application on DESI data



Thank you!



Likelihood analysis

- Theory model

$$\log \mathcal{L}(\theta|\mathbf{d}) \propto -\frac{1}{2} [\mathbf{X}_d - \mathbf{X}_t(\theta)]^T C^{-1} [\mathbf{X}_d - \mathbf{X}_t(\theta)]$$

- To model WST (and P(k)) cosmological dependence, we use the approximation:

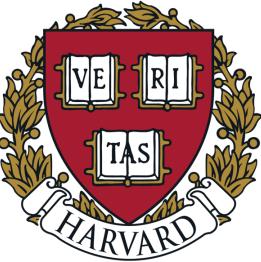
$$\mathbf{X}_t(\theta) = \mathbf{X}_t(\theta_{\text{fid}}) + (\theta - \theta_{\text{fid}}) \nabla_{\theta} \mathbf{X}$$

Prediction for fiducial cosmology
↓

Constructed from
'Linear derivative grid'
of cosmologies
↓

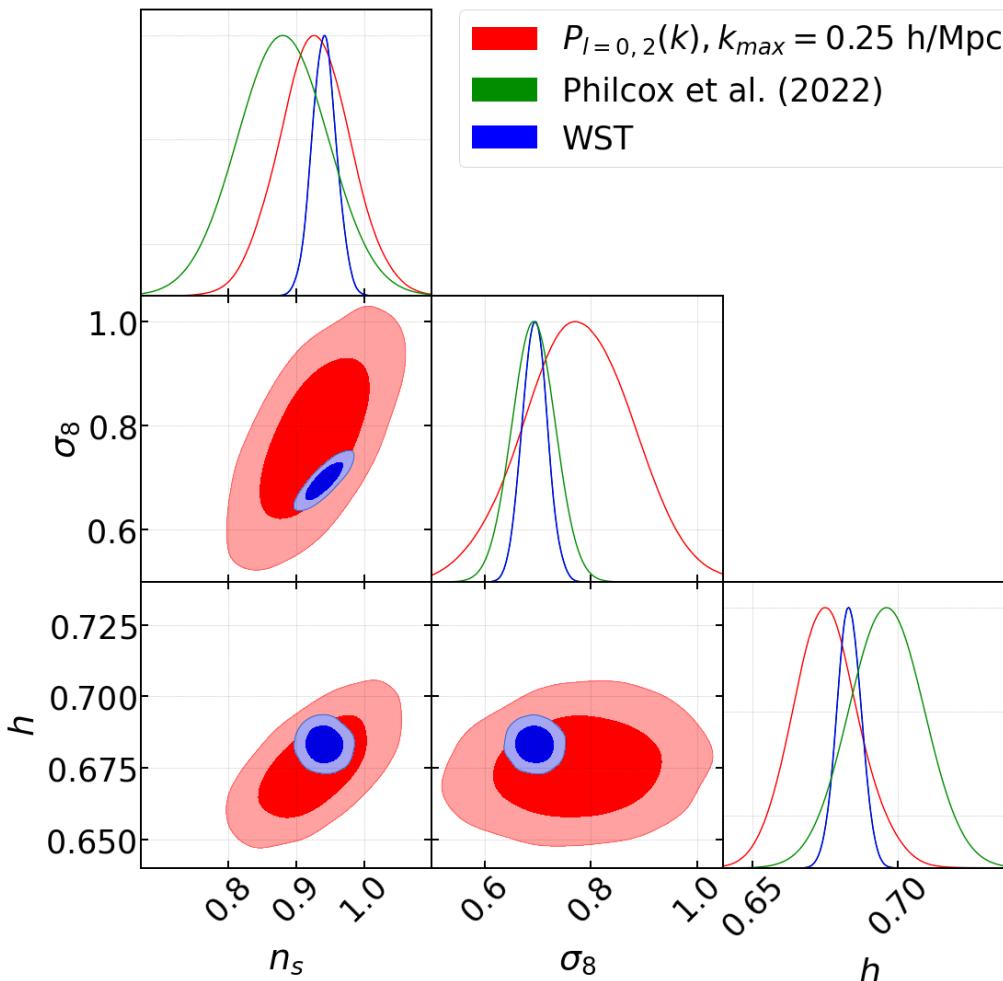
- + Additional derivative steps in the 7 HOD parameters

ω_b	ω_c	n_s	σ_8
0.02237	0.1200	0.9649	0.8114
0.02282	0.1200	0.9649	0.8114
0.02193	0.1200	0.9649	0.8114
0.02237	0.1240	0.9649	0.8114
0.02237	0.1161	0.9649	0.8114
0.02237	0.1200	1.0249	0.8114
0.02237	0.1200	0.9049	0.8114
0.02237	0.1200	0.9649	0.8698
0.02237	0.1200	0.9649	0.7532



Likelihood analysis

- Likelihood analysis using a BBN prior on ω_b $\omega_b = 0.02268 \pm 0.00038$



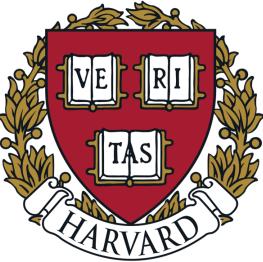
- BOSS with WST:

$$H_0 = 68.31^{+0.42}_{-0.42} \text{ km/s/Mpc} \quad \sigma_8 = 0.695^{+0.024}_{-0.024}$$

- Planck results:

$$H_0 = 67.4 \pm 0.5 \text{ km/s/Mpc} \quad \sigma_8 = 0.811 \pm 0.006$$

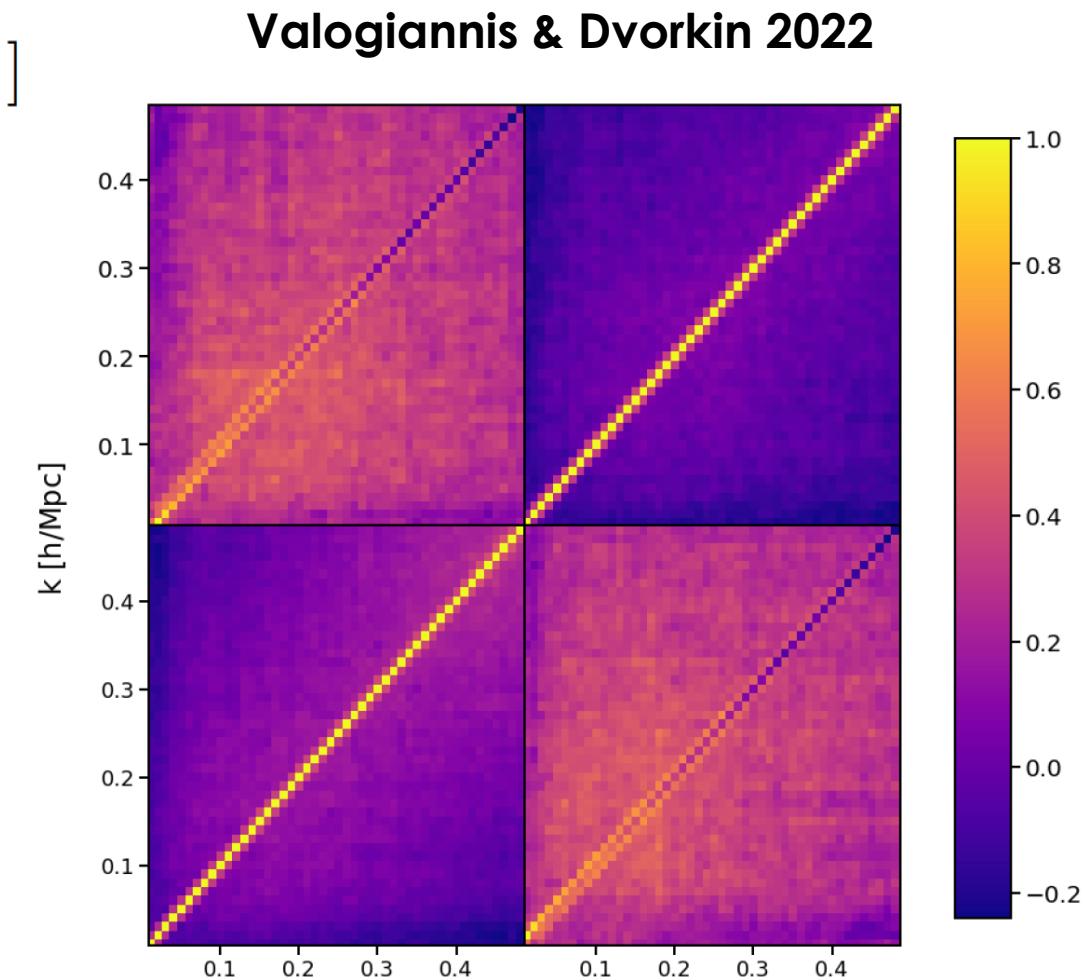
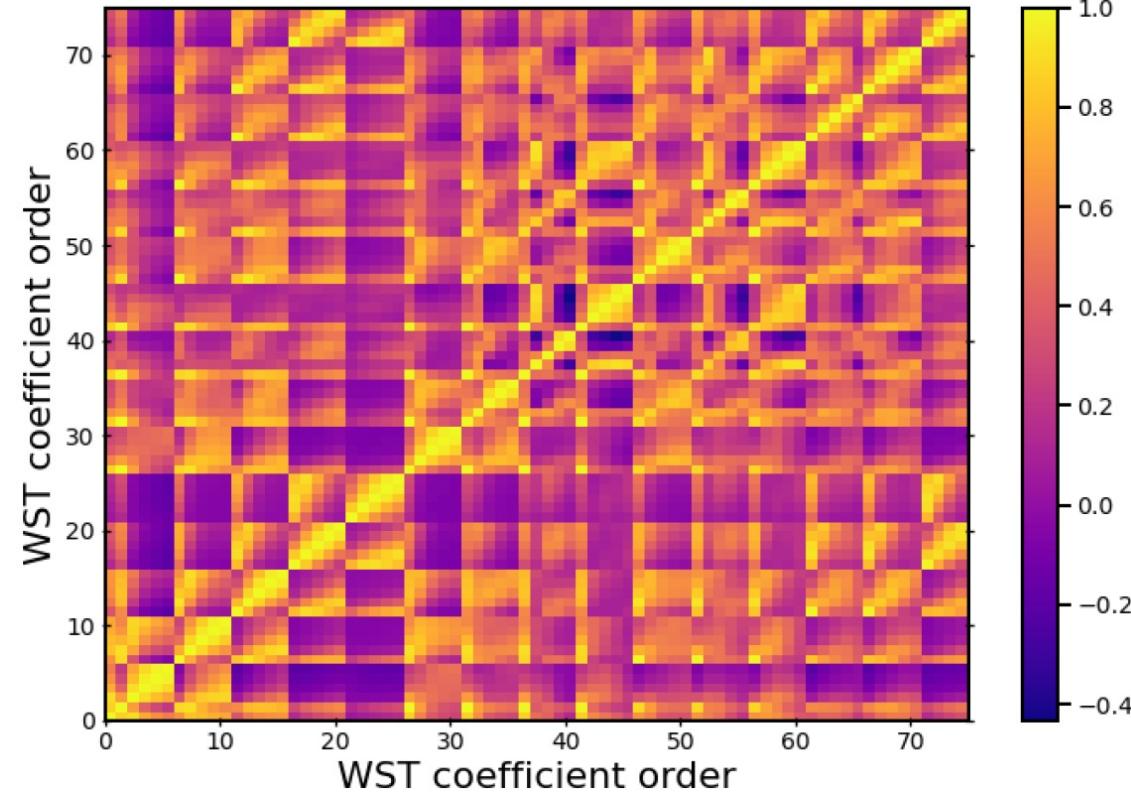
- Parameter mean values from WST & $P(k)$ always consistent with each other within 1σ (of the $P(k)$)
- Much **tighter** errors from WST compared to $P(k)$!
- H_0 determined from WST with 0.6% accuracy!
- $\sigma_8 = 0.695^{+0.024}_{-0.024}$ in tension with Planck result
In agreement with recent BOSS analyses
(Philcox & Ivanov, 2022, Chen et al. 2022 a & b)

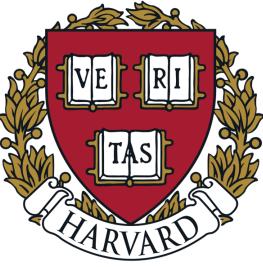


Likelihood analysis

- Covariance matrix obtained from N=2048 PATCHY mocks (S. A. Rodriguez-Torres et al., 2016)

$$\log \mathcal{L}(\theta|\mathbf{d}) \propto -\frac{1}{2} [\mathbf{X}_d - \mathbf{X}_t(\theta)]^T C^{-1} [\mathbf{X}_d - \mathbf{X}_t(\theta)]$$





The Wavelet Scattering Transform (WST)

- 3-dimensional WST implementation with 'solid harmonic' wavelets (Eickenberg et al. (2018))

$$S_0 = \langle |I(\vec{x})|^q \rangle,$$

$$S_1(j_1, l_1) = \left\langle \left(\sum_{m=-l_1}^{m=l_1} |I(\vec{x}) * \psi_{j_1, l_1}^m(\vec{x})|^2 \right)^{\frac{q}{2}} \right\rangle,$$

$$S_2(j_2, j_1, l_1) = \left\langle \left(\sum_{m=1}^{m=l_1} |U_1(j_1, l_1)(\vec{x}) * \psi_{j_2, l_1}^m(\vec{x})|^2 \right)^{\frac{q}{2}} \right\rangle$$

$$U_1(j_1, l_1)(\mathbf{x}) = \left(\sum_{m=-l_1}^{m=l_1} |I(\mathbf{x}) * \psi_{j_1, l_1}^m(\mathbf{x})|^2 \right)^{\frac{1}{2}}$$

$$\psi_l^m(\mathbf{x}) = \underbrace{\frac{1}{(2\pi)^{3/2}} e^{-|\mathbf{x}|^2/2\sigma^2}}_{\text{Gaussian envelope}} |\mathbf{x}|^l Y_l^m \left(\frac{\mathbf{x}}{|\mathbf{x}|} \right) \underbrace{\text{Solid Harmonics}}$$

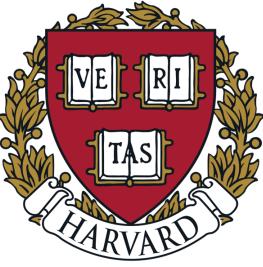
- Dilated by dyadic scales 2^{j_1}

$$\psi_{j_1, l_1}^m(\mathbf{x}) = 2^{-3j_1} \psi_{l_1}^{m_1}(2^{-j_1} \mathbf{x})$$

- Wavelets in the literature

- Bump steerable wavelets (Eickenberg et al, 2022, Allys et al, 2020)
- Morlet wavelets (Cheng et al. 2020b, Cheng & Menard 2021a)

- Implemented in KYMATIO package (Andreux et al. 2019)



The Wavelet Scattering Transform (WST)

- Raising modulus to powers $q < 1$ emphasizes on cosmic voids (while $q > 1$ on density peaks)

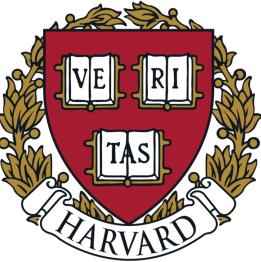
$$S_0 = \langle |I(\vec{x})|^q \rangle,$$

$$S_1(j_1, l_1) = \left\langle \left(\sum_{m=-l_1}^{m=l_1} |I(\vec{x}) * \psi_{j_1, l_1}^m(\vec{x})|^2 \right)^{\frac{q}{2}} \right\rangle,$$

$$S_2(j_2, j_1, l_1) = \left\langle \left(\sum_{m=1}^{m=l_1} |U_1(j_1, l_1)(\vec{x}) * \psi_{j_2, l_1}^m(\vec{x})|^2 \right)^{\frac{q}{2}} \right\rangle$$

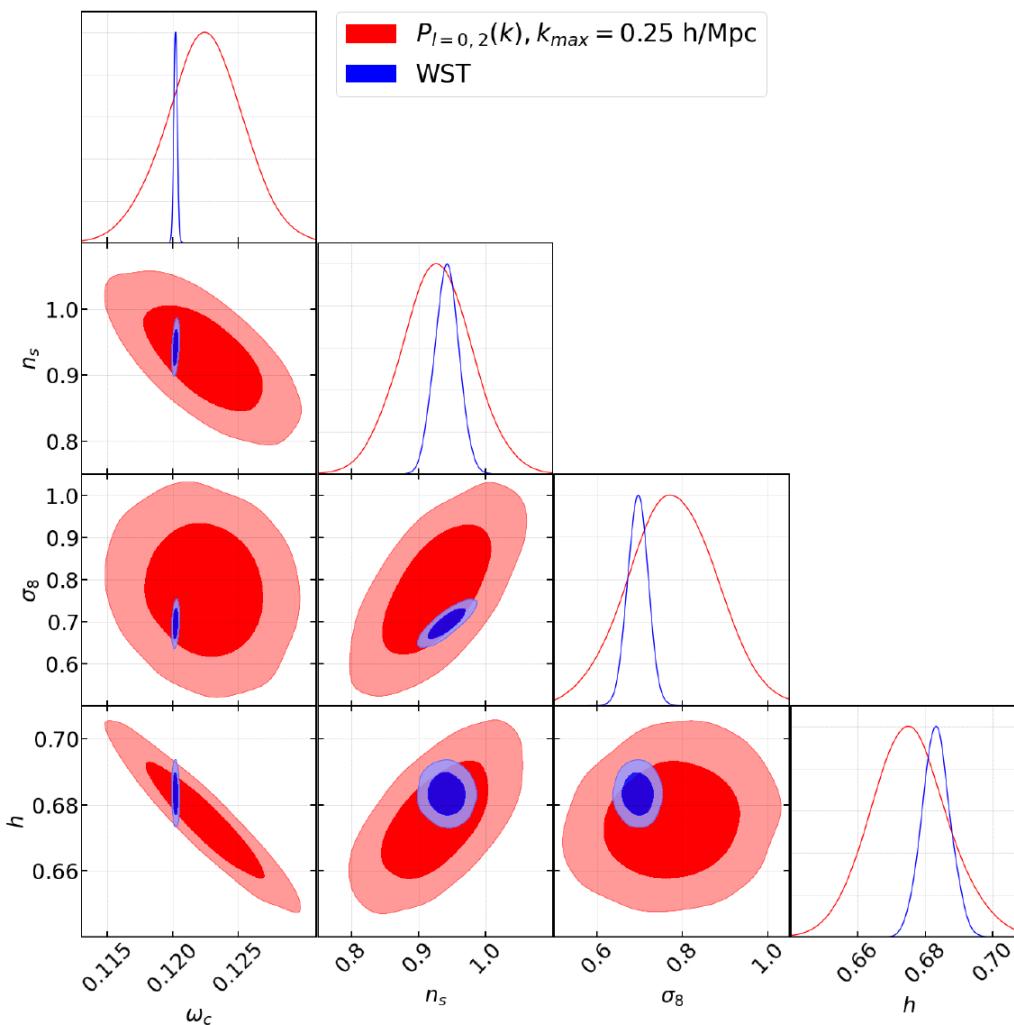
- $q=0.8$ found to be optimal for 3D LSS studies (**Valogiannis & Dvorkin 2021**)

- Applied WST on 3D matter over-density field from *Quijote* simulations (F. Villaescusa-Navarro et al., 2019)
- Particularly sensitive to neutrino mass
- WST matches and exceeds performance of marked $P(k)$ (studied by Massara et al, PRL 2020)



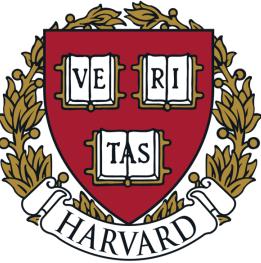
Likelihood analysis

- Likelihood analysis using a BBN prior on ω_b $\omega_b = 0.02268 \pm 0.00038$



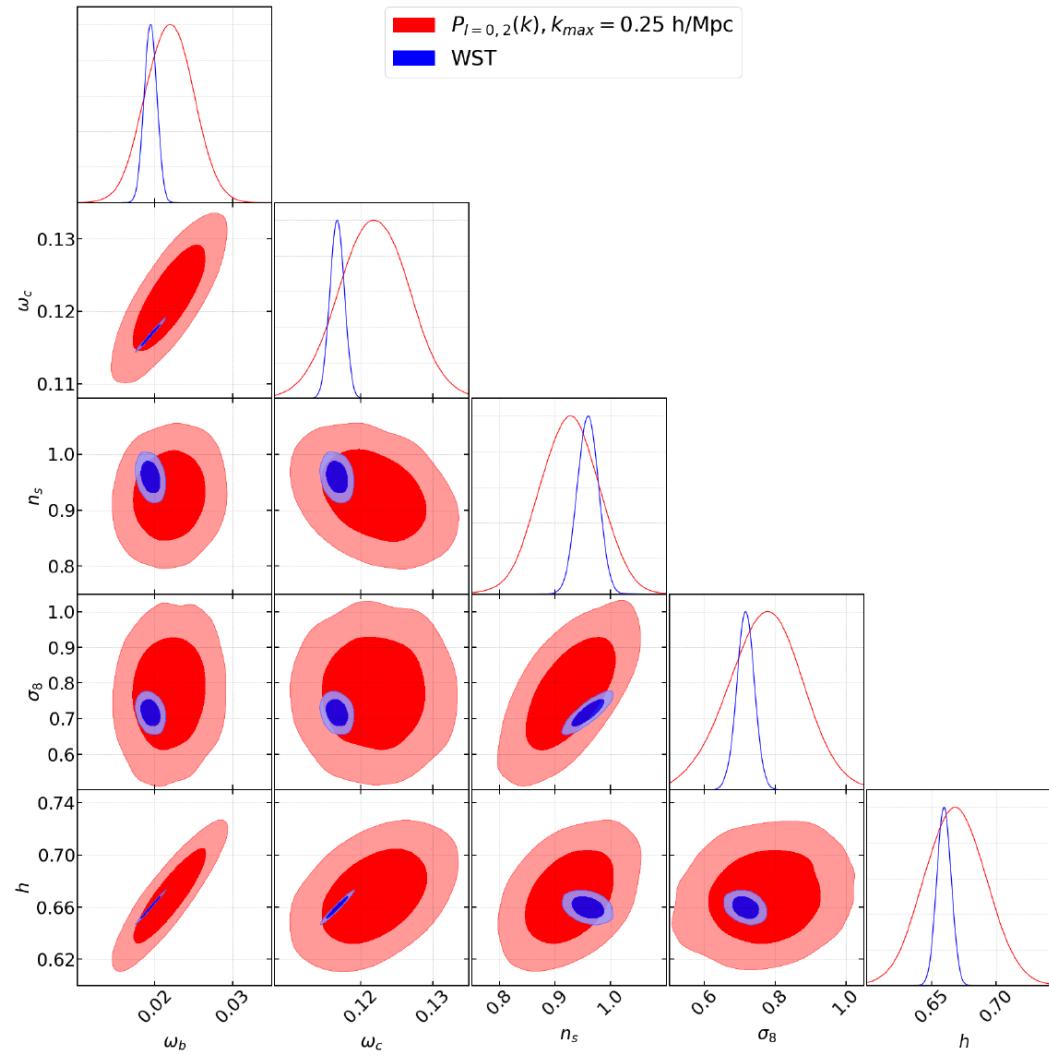
	BBN prior on ω_b		unrestricted priors	
	P(k)	WST	P(k)	WST
ω_b	$0.02267^{+0.00045}_{-0.00045}$	$0.02268^{+0.00036}_{-0.00036}$	$0.0217^{+0.0043}_{-0.0043}$	$0.01946^{+0.0008}_{-0.0008}$
ω_c	$0.1223^{+0.0031}_{-0.0028}$	$0.1202^{+0.00013}_{-0.00013}$	$0.1217^{+0.0058}_{-0.0058}$	$0.11672^{+0.001}_{-0.001}$
n_s	$0.928^{+0.075}_{-0.075}$	$0.942^{+0.018}_{-0.018}$	$0.921^{+0.057}_{-0.049}$	$0.959^{+0.019}_{-0.019}$
σ_8	$0.77^{+0.14}_{-0.14}$	$0.695^{+0.024}_{-0.024}$	$0.762^{+0.11}_{-0.094}$	$0.716^{+0.025}_{-0.025}$
h	$0.676^{+0.010}_{-0.012}$	$0.6831^{+0.0042}_{-0.0042}$	$0.668^{+0.024}_{-0.024}$	$0.66^{+0.0055}_{-0.0055}$

- Parameter mean values from WST & P(k) always consistent with each other within 1σ (of the P(k))
- ~4x28x **tighter** errors from WST compared to P(k)!
- H_0 determined from WST with 0.6% accuracy!
- $\sigma_8 = 0.695^{+0.024}_{-0.024}$ in tension with Planck result
In agreement with recent BOSS analyses
(Philcox & Ivanov, 2022, Chen et al. 2022 a & b)



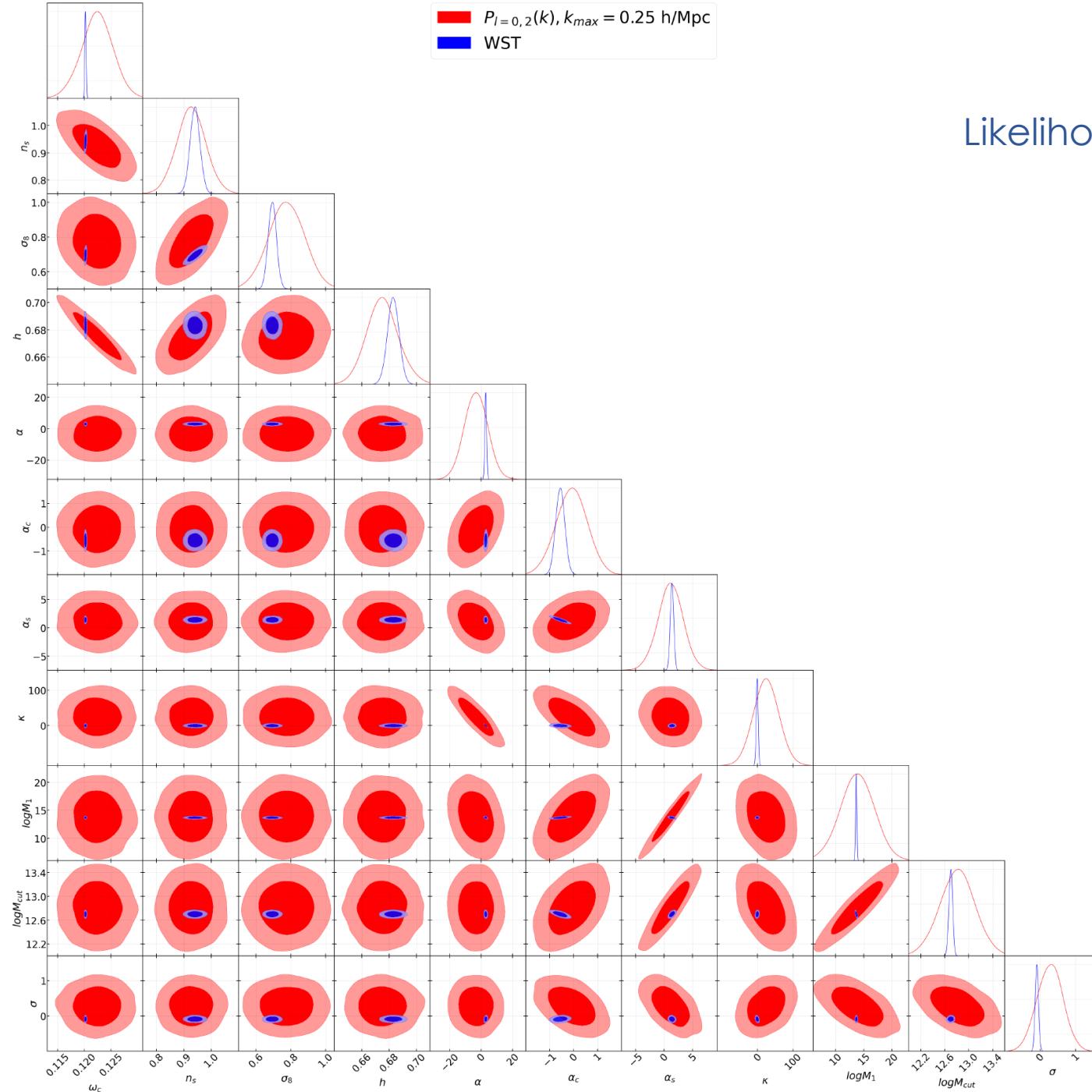
Likelihood analysis

- Likelihood analysis using flat unrestricted priors

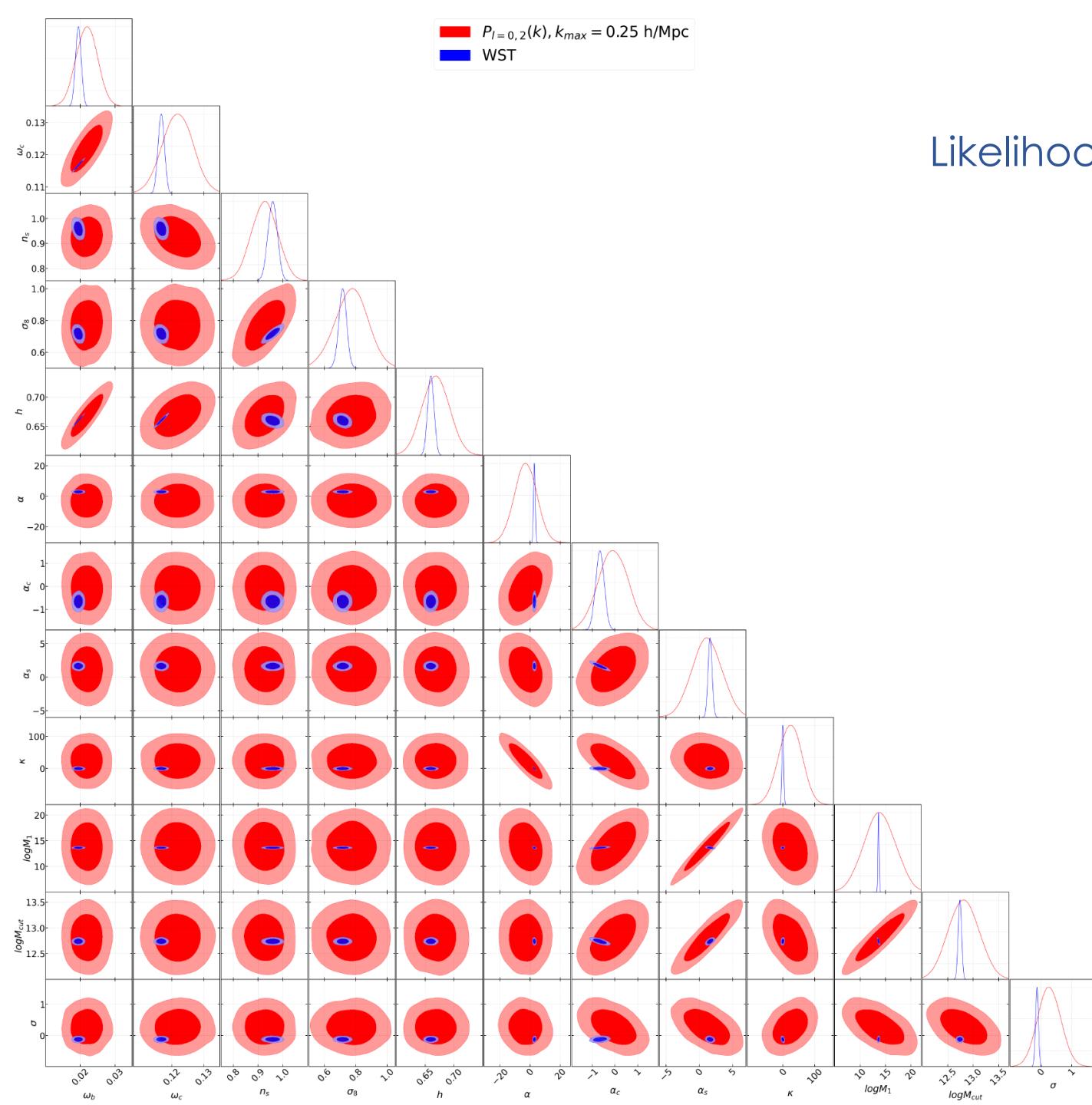


	BBN prior on ω_b		unrestricted priors	
	P(k)	WST	P(k)	WST
ω_b	$0.02267^{+0.00045}_{-0.00045}$	$0.02268^{+0.00036}_{-0.00036}$	$0.0217^{+0.0043}_{-0.0043}$	$0.01946^{+0.0008}_{-0.0008}$
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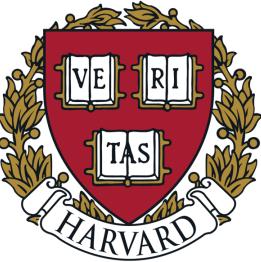
- Parameter mean values from WST & P(k) again always consistent with each other within 1σ (of the P(k))
- ~3x6x **tighter** errors from WST compared to P(k)!



Likelihood analysis using a BBN prior on ω_b



Likelihood analysis using flat unrestricted priors



First WST application on 3D LSS

- First WST application on 3D matter density field! (Valogiannis & Dvorkin, 2021)

$$I(\vec{x}) \equiv \delta_m(\vec{x}) = \frac{\rho_m(\vec{x})}{\bar{\rho}_m} - 1.0 , \text{ resolution } N_{grid} = 256^3$$

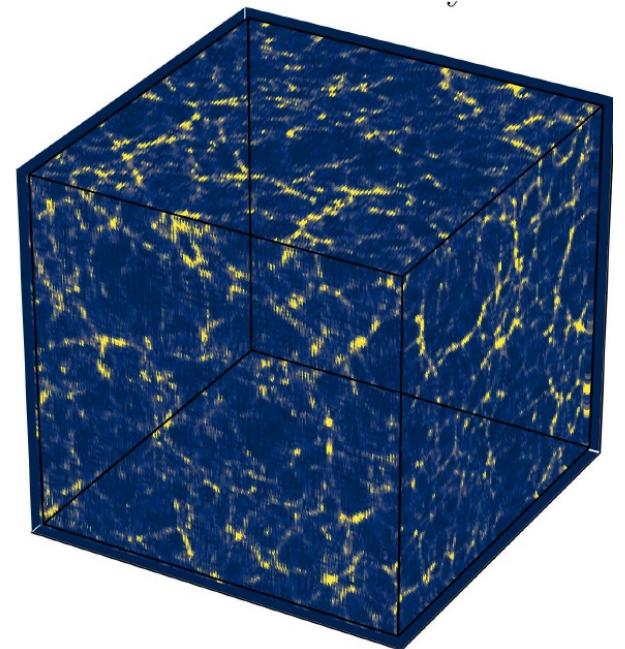
- Evaluated from the *Quijote* simulations (F. Villaescusa-Navarro et al., 2019)
- Fiducial cosmology

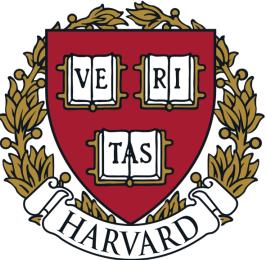
$\Omega_m = 0.3175$, $\Omega_b = 0.049$, $h = 0.6711$

$n_s = 0.9624$, $\sigma_8 = 0.834$, $M_\nu = 0.0$ eV, and $w = -1$

Box L=1.0 Gpc/h

- In presence of massive neutrinos, trace both:
 - $\delta_m = \delta_{CDM} + \delta_b + \delta_\nu$ Total 'm' field
 - $\delta_{cb} = \delta_{CDM} + \delta_b$ 'cb' field





Fisher forecast

Fisher forecasting

- 15,000 realizations for fiducial cosmology
- 7,000 for linear derivatives in parameters

$$F_{\alpha\beta} = \frac{\partial O_i}{\partial \theta_\alpha} C_{ij}^{-1} \frac{\partial O_j^T}{\partial \theta_\beta}$$

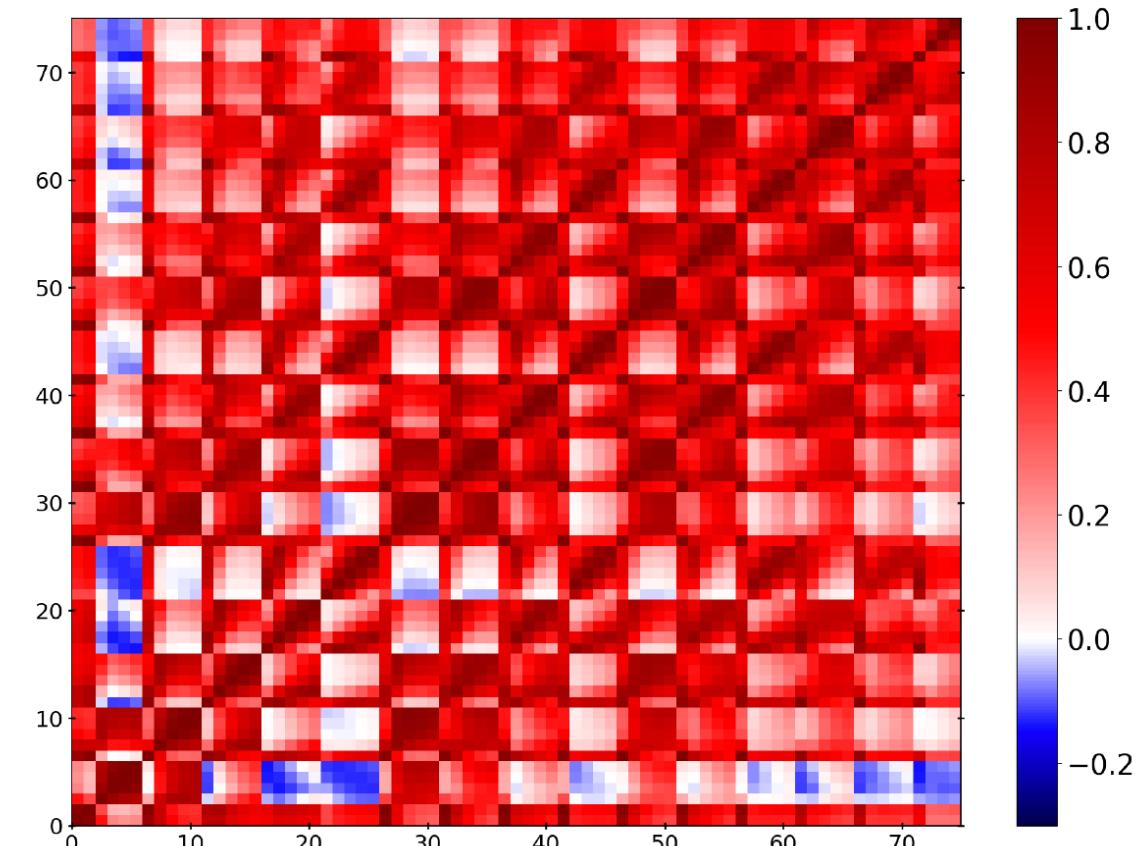
- Marginalized $\sigma_\alpha = \sqrt{(F^{-1})_{\alpha\alpha}}$
- for $\theta_\alpha = \{\Omega_m, \Omega_b, H_0, n_s, \sigma_8, M_\nu\}$, $z=0$

Comparing 3 observables O_i :

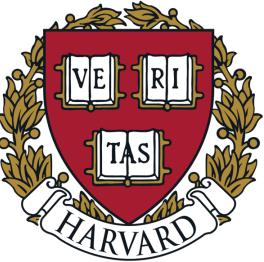
- Power spectrum **P(k)**
- Marked power spectrum **M(k)**
- $S_0 + S_1 + S_2$ **WST** coefficients

Evaluated using *kymatio* package (Andreux et al. 2019)

<https://www.kymat.io/>



WST coefficients - correlation matrix
Valogiannis & Dvorkin 2021



Marked Power Spectrum

- Marked correlation function generalizes 2-point function

$$\mathcal{M}(r) = \frac{1}{n(r)\bar{m}^2} \sum_{ij} \delta_D(|\mathbf{x}_i - \mathbf{x}_j| - r) m_i m_j = \frac{1 + W(r)}{1 + \xi(r)}$$

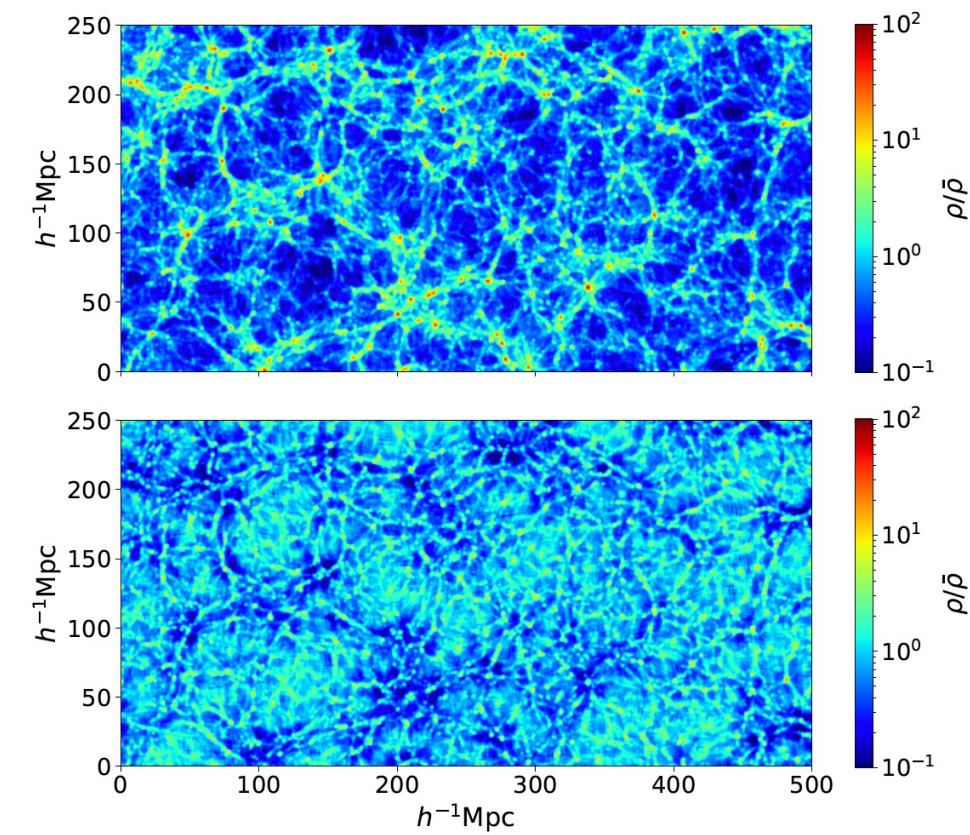


- Each galaxy weighted by mark 'm'
- Inverse density weighted mark (highlights voids)

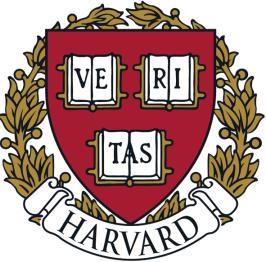
$$m[\mathbf{x}, R, \delta_s, p] = \left(\frac{1 + \delta_s}{1 + \delta_s + \delta_R(\mathbf{x})} \right)^p$$

- Can constrain MG (M. White 2016, **Valogiannis & Bean 2018**, Alam et al., 2021)
- Can constrain neutrino mass (Massara et al 2020)

$\delta(x)$



Massara et al 2020



WST sensitivity to neutrino mass

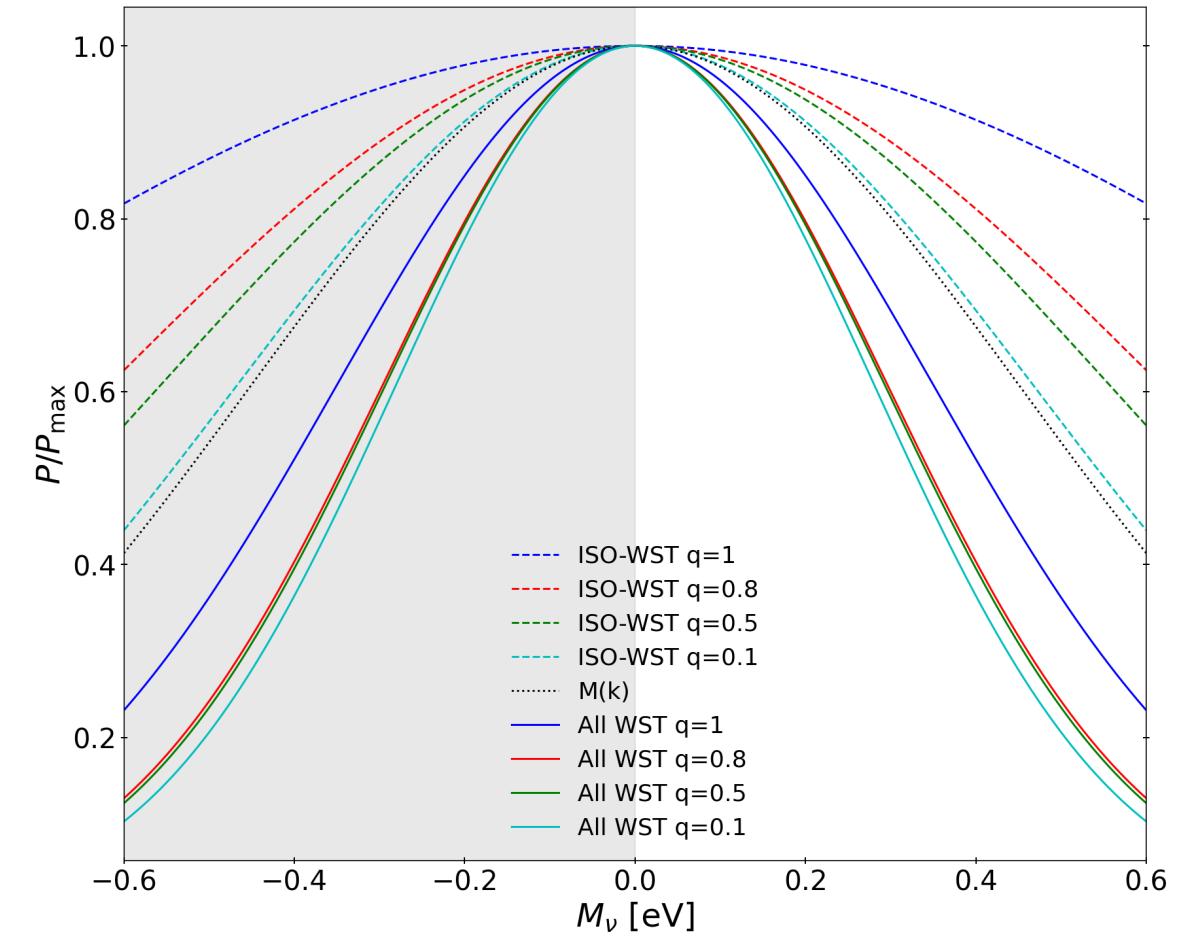
- Raising modulus to powers $q < 1$ emphasizes on cosmic voids
- Very sensitive to neutrino mass!

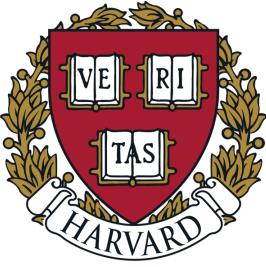
$$S_0 = \langle |I(\vec{x})|^q \rangle,$$

$$S_1(j_1, l_1) = \left\langle \left(\sum_{m=-l_1}^{m=l_1} |I(\vec{x}) * \psi_{j_1, l_1}^m(\vec{x})|^2 \right)^{\frac{q}{2}} \right\rangle,$$

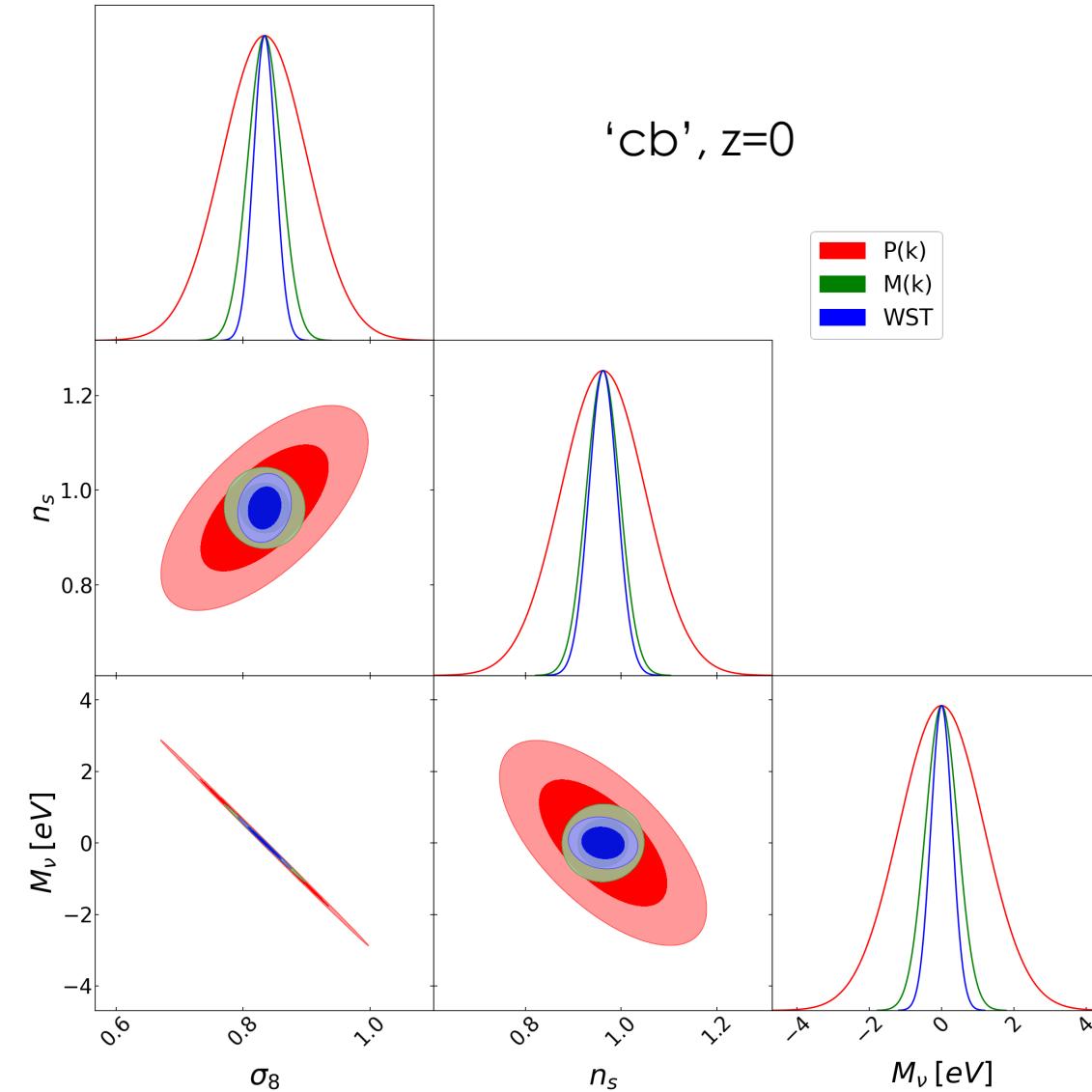
$$S_2(j_2, j_1, l_1) = \left\langle \left(\sum_{m=1}^{m=l_1} |U_1(j_1, l_1)(\vec{x}) * \psi_{j_2, l_1}^m(\vec{x})|^2 \right)^{\frac{q}{2}} \right\rangle$$

- $q=0.8$ found to be optimal



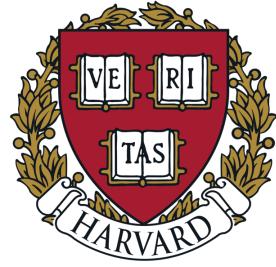


Great improvement over P(k)!

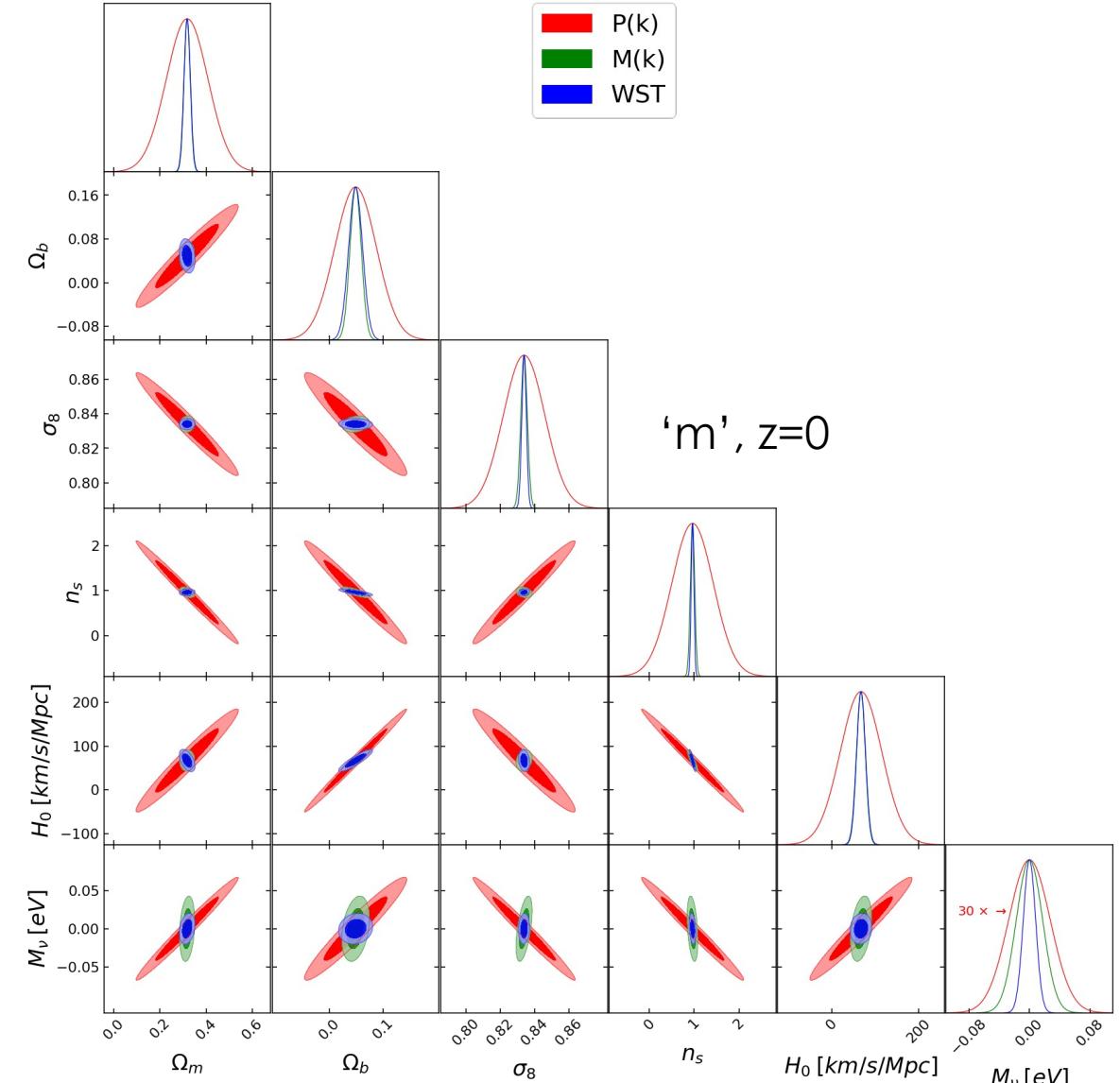
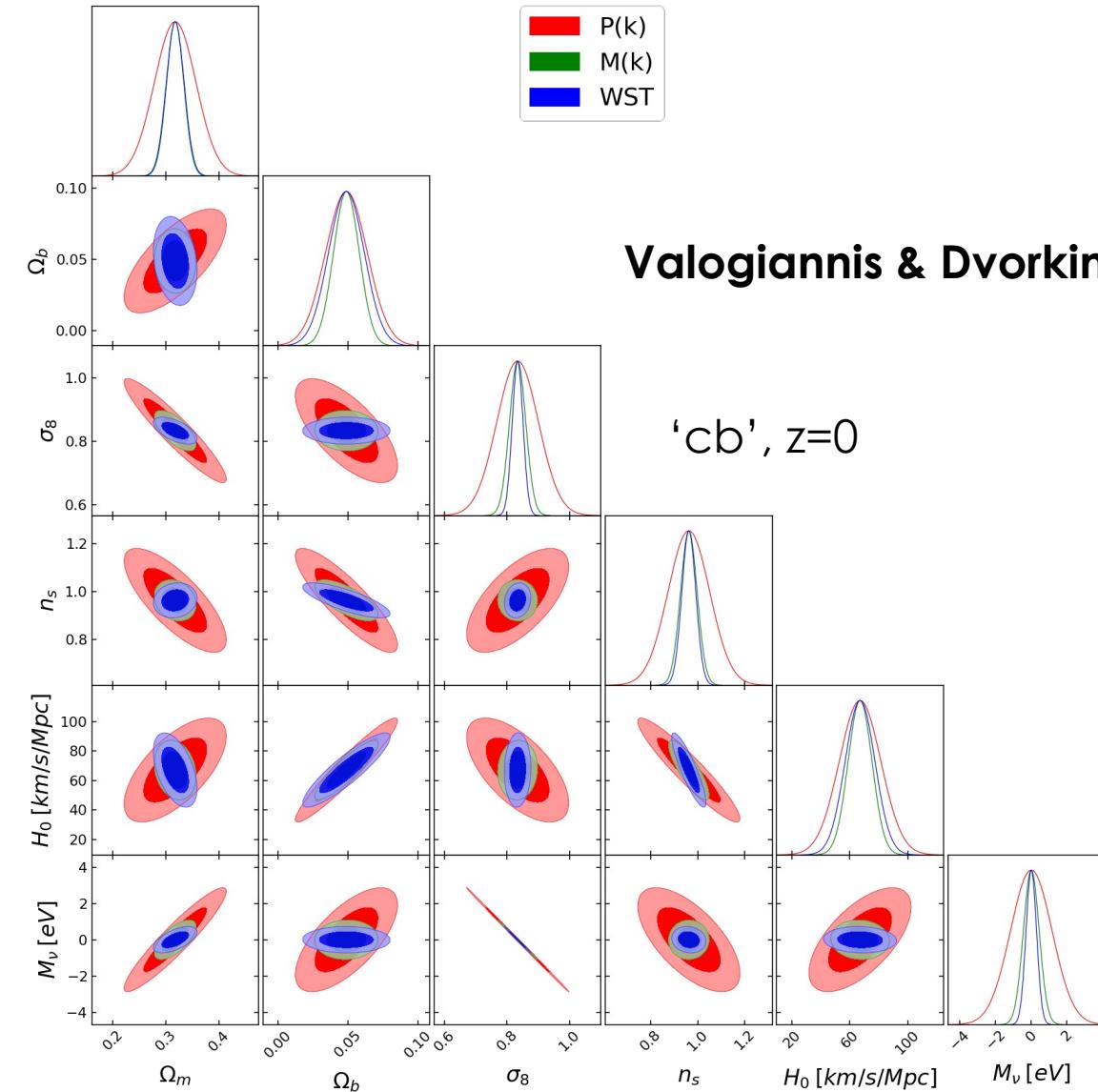


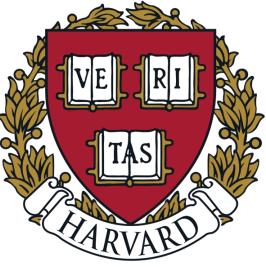
- WST delivers **very large** improvement in the $1-\sigma$ errors for all parameters!
 - $\sim 1.2\text{-}4$ x **tighter** errors than from 'cb' P(k)!
 - Constrains on neutrino mass:
 - ~ 4 x **tighter** than 'cb' P(k)!
 - ~ 1.6 x **tighter** than 'cb' M(k)!
 - $\sim 3\text{x}100$ x **tighter** errors than from 'm' P(k)

Matter type	'm'			'cb'		
	P(k)	M(k)	WST	P(k)	M(k)	WST
$\sigma(\Omega_m)$	0.076	0.013	0.014	0.040	0.016	0.016
$\sigma(\Omega_b)$	0.033	0.010	0.012	0.015	0.009	0.012
$\sigma(\sigma_8)$	0.01	0.002	0.001	0.067	0.026	0.017
$\sigma(n_s)$	0.39	0.044	0.031	0.088	0.035	0.029
$\sigma(H_0)$ [km/s/Mpc]	40.62	9.50	10.34	14.42	8.28	10.32
$\sigma(M_\nu)$ [eV]	0.72	0.016	0.008	1.17	0.45	0.29



Great improvement over P(k)!





Physical explanation of results

Why does the WST work so well??

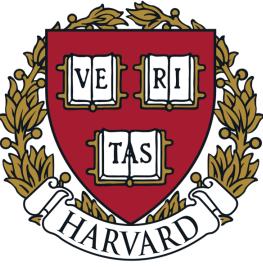
WST key physical properties

- Successive WST layers pick up information >2-point function ✓
 - Known to encode additional information (eg. Hahn et al. 2020 & 2021)
- +- Choice of $q < 1$ highlights cosmic voids (under-densities) ✓
 - Sensitive cosmological probe (eg. Massara et al, 2020)

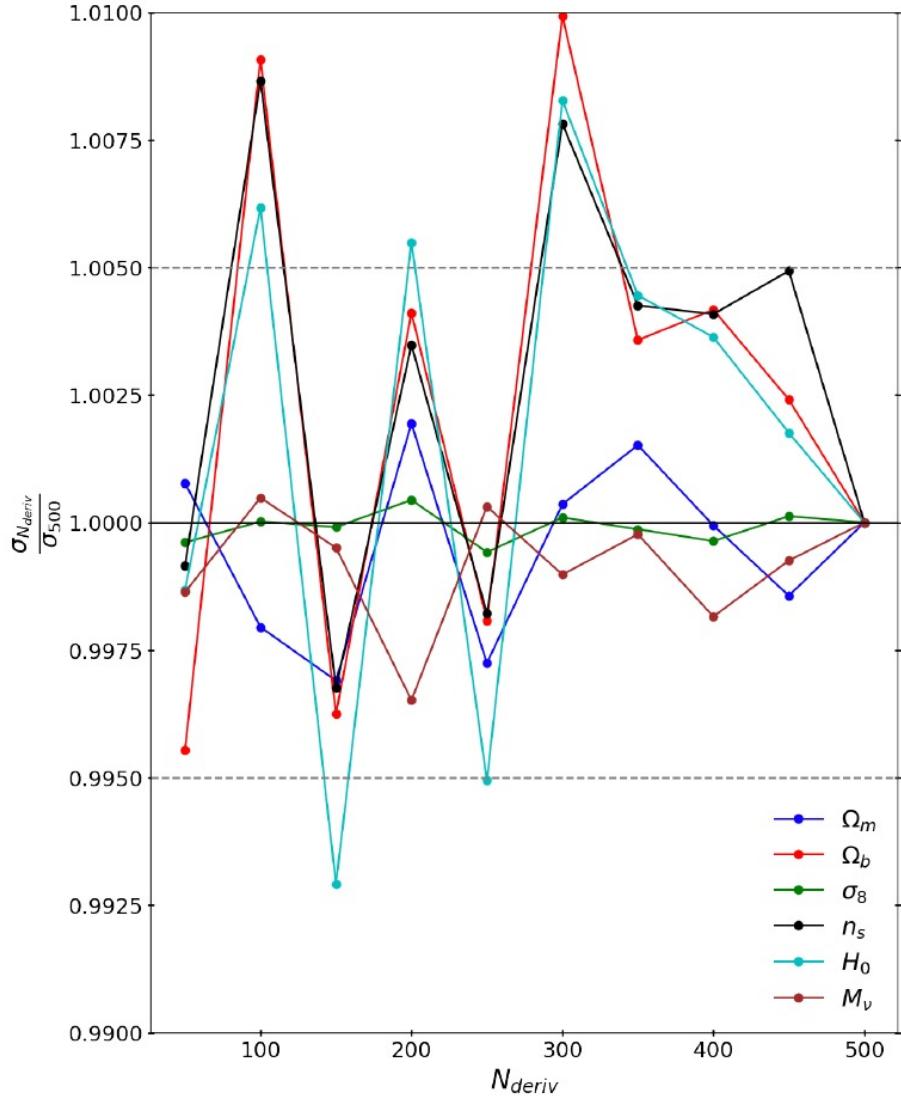


Enhanced cosmological information

- Parallels to marked $M(k)$ (Massara et al, 2020)



Convergence



- WST vector exhibits remarkable numerical convergence w.r.t. Fisher predictions