### Solving the exact background Collisional Boltzmann Equation (CBE) for DM-baryon scattering

Suroor Seher Gandhi

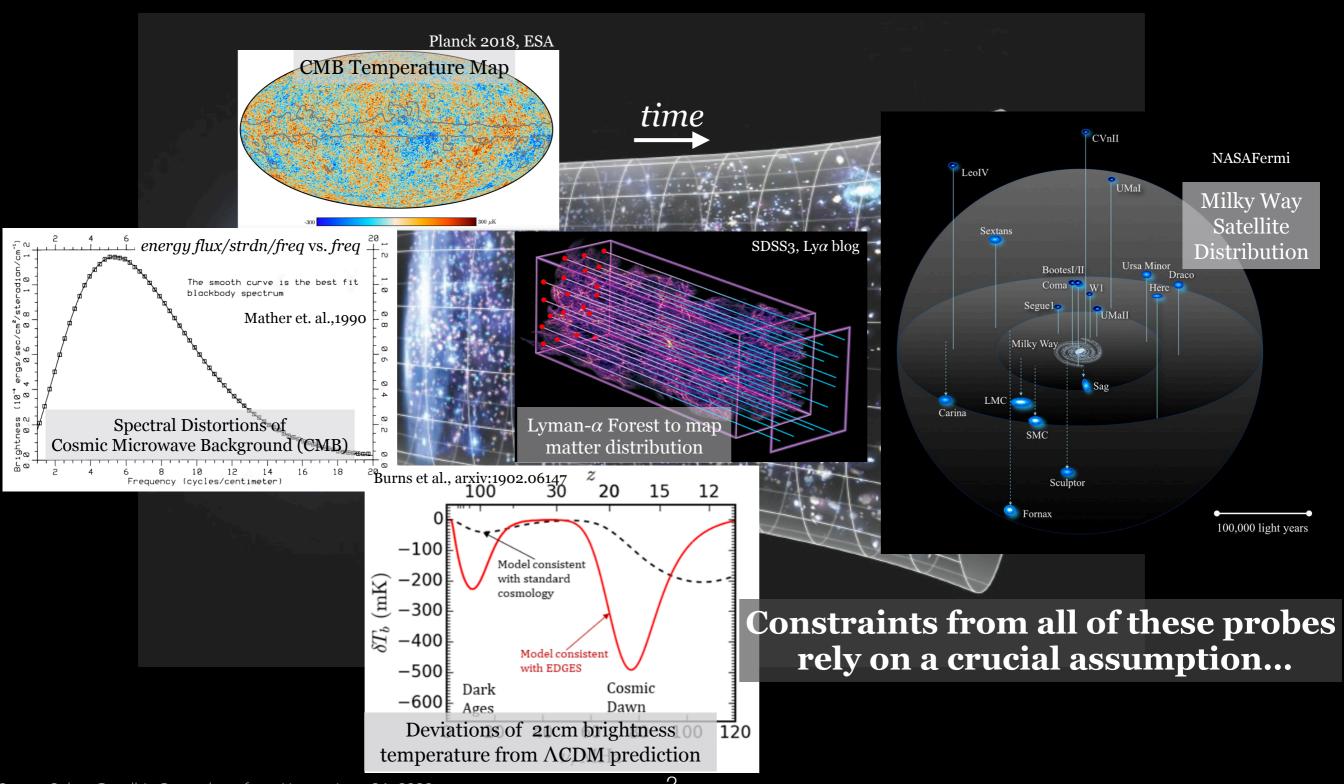
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Parallel Talk Cosmology from Home (2022)

Paper on arXiv: 2205.05536 (submitted to Phys. Rev. D)

# A Few Cosmological Probes of DM-baryon (χ-s) Scattering



$$f_{\chi}(\overrightarrow{v}) = f_{\chi}^{\text{MB}}(\overrightarrow{v})$$

- *Almost all* current constraints on  $\chi$ -s scattering cross-section  $\sigma_{\chi s}$  assume  $f_{\chi}(\overrightarrow{v}) = f_{\chi}^{\text{MB}}(\overrightarrow{v})$ 
  - $\Rightarrow \dot{Q}_{\chi}, \overrightarrow{p}_{\chi}$ : analytical heat & momentum transfer rates
  - $\Rightarrow$  No need to solve the Boltzmann eq.:

$$\mathrm{d}f_{\chi}/\mathrm{d}t = C_{\chi s}[f_{\chi}] + C_{\chi \chi}[f_{\chi}]$$

- But, if  $\chi$ - $\chi$  scattering is ineff.,  $f_{\chi}(\overrightarrow{v}) \neq f_{\chi}^{MB}(\overrightarrow{v})$  after  $\chi$ -s decoupling
- If  $f_{\chi} = f_{\chi}^{\text{MB}} \Rightarrow C_{\chi\chi}[f_{\chi}^{\text{MB}}] = 0$ , eliminates the analysis of DM self-interactions

$$f_{\chi}(\overrightarrow{v}) \neq f_{\chi}^{\text{MB}}(\overrightarrow{v})$$

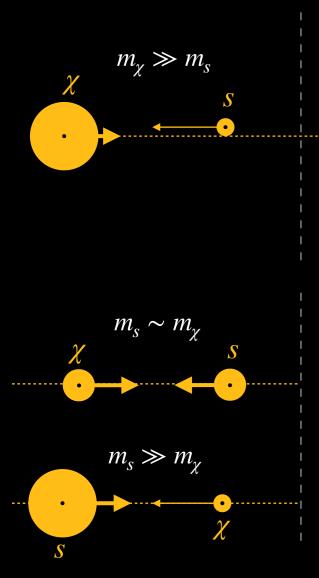
- If  $f_{\chi}(\overrightarrow{v}) \neq f_{\chi}^{MB}(\overrightarrow{v})$ , then  $\dot{Q}_{\chi}$ ,  $\dot{\overrightarrow{p}}_{\chi}$  are no longer analytical
- Need to implement the full collision operator

$$C_{\chi s}[f_{\chi}](\overrightarrow{v}) = \int d^{3}v' \Big( \Gamma_{\chi s}(\overrightarrow{v}' \to \overrightarrow{v}) f_{\chi}(\overrightarrow{v}') - \Gamma_{\chi s}(\overrightarrow{v} \to \overrightarrow{v}') f_{\chi}(\overrightarrow{v}) \Big)$$

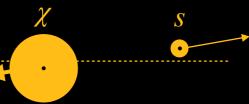
- $\Gamma_{\chi s}(\overrightarrow{v} \to \overrightarrow{v}')$ :  $\chi$ -s scattering rate from DM velocity  $\overrightarrow{v}$  to  $\overrightarrow{v}'$ , is generally a 5D non-analytic integral
- Highly non-trivial calculations, incompatible with standard CMB codes

Motivation: can analyze DM self-interactions, since  $C_{\chi\chi}[f_{\chi}] \neq 0$ 

# Previous Work: Fokker-Planck (FP) Approximation



Diffusive scattering: small change in  $\overrightarrow{v}_{\chi}$ 



Is scattering diffusive?

Which is more accurate, FP approx. or MB assumption?

- The FP (or diffusion)
   approximation to the full CBE
   was formulated in Ali-Haïmoud
   (2019)
- Numerically more tractable
- Scattering is diffusive when  $\Delta \overrightarrow{v}_{\chi}$  is small
- Need an exact method to determine the accuracy of the diffusion formalism for  $m_{\chi} \lesssim m_{s}$

# Finding an *exact* solution to the *background* CBE

Homogeneity & Isotropy:

$$\mathscr{F}_{\chi}(\overrightarrow{x},\overrightarrow{v}) = \overline{n}_{\chi}\overline{f}_{\chi}(v) = \overline{n}_{\chi}f_{\chi}^{1D}(v)/(4\pi v^2)$$

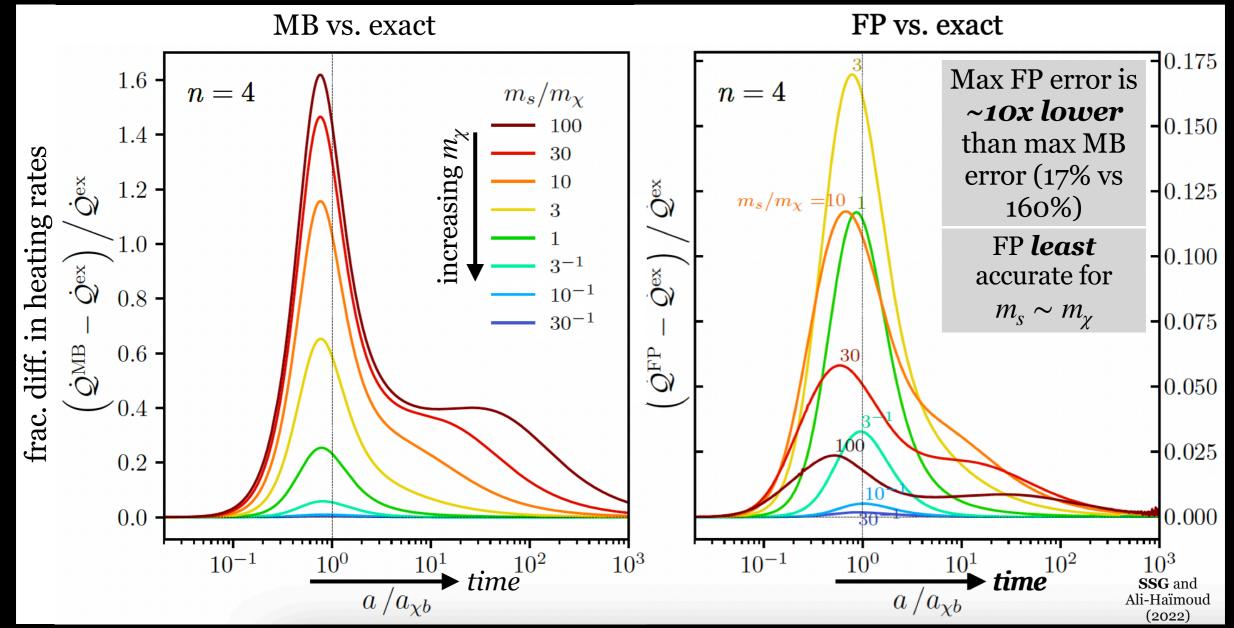
- Do not account for  $\chi$ - $\chi$  scattering (quadratic in  $\bar{f}_{\chi}$ ), and calculate the *maximal* error due to the MB and FP methods
- Need to implement the 1D  $\chi$ -s collision operator:

$$C_{\chi s}^{\mathrm{1D}}[f_{\chi}](v) = \int dv' \left( \Gamma_{\chi s}^{\mathrm{1D}}(v' \to v) f_{\chi}^{\mathrm{1D}}(v') - \Gamma_{\chi s}^{\mathrm{1D}}(v \to v') f_{\chi}^{\mathrm{1D}}(v) \right)$$

- We show that  $\Gamma_{\chi s}^{1D}(v \to v')$ —a 4D integral—is:
  - 1. Reducible to a **1D** integral for isotropic differential cross-sections  $d\sigma_{\gamma s}/d\Omega$ ,
  - 2. Fully analytical if  $d\sigma_{ys}/d\Omega \propto v_{ys}^n$ ,  $n \in \{0, 2, 4, ...\}$
- Solve  $df_{\chi}^{1D}/dt = C_{\chi s}^{1D}[f_{\chi}^{1D}]$  and obtain  $f_{\chi}^{1D}(v,t)$ !

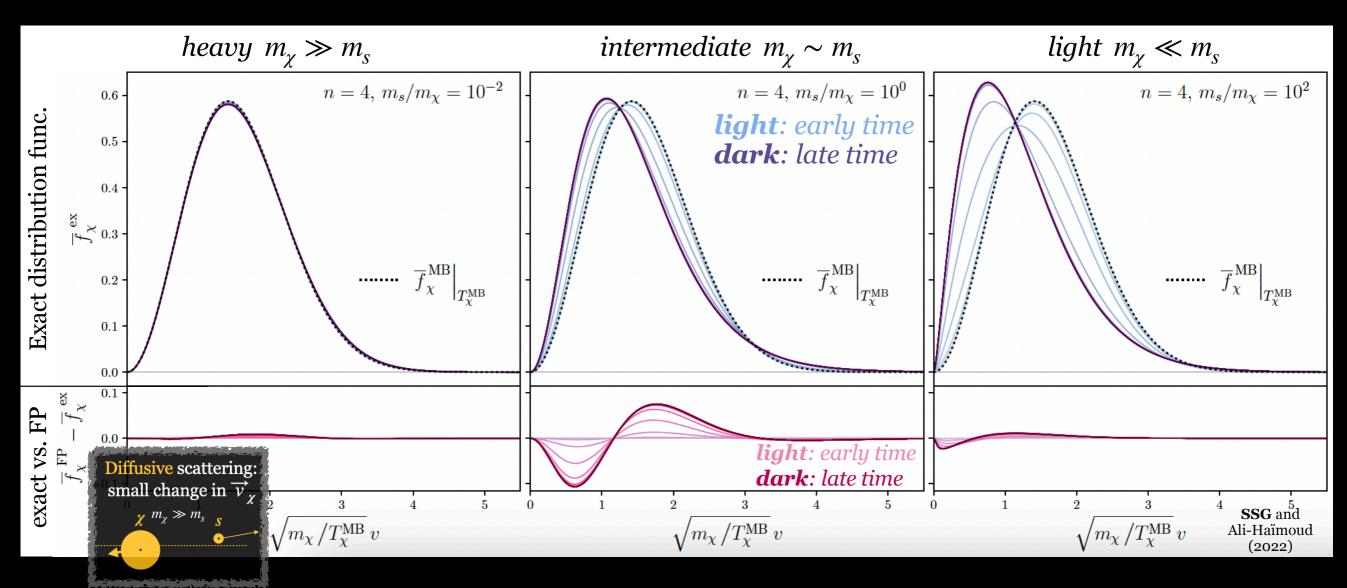
# Results: Comparing $\dot{Q}_{\chi}^{\text{MB}}$ and $\dot{Q}_{\chi}^{\text{FP}}$ to $\dot{Q}_{\chi}^{\text{ex}}$

- Use heating rate  $\dot{Q}_{\chi}$  to compare MB, FP, and exact methods
- Showing fractional differences in  $\dot{Q}_{\chi}$  for a range of mass ratios:



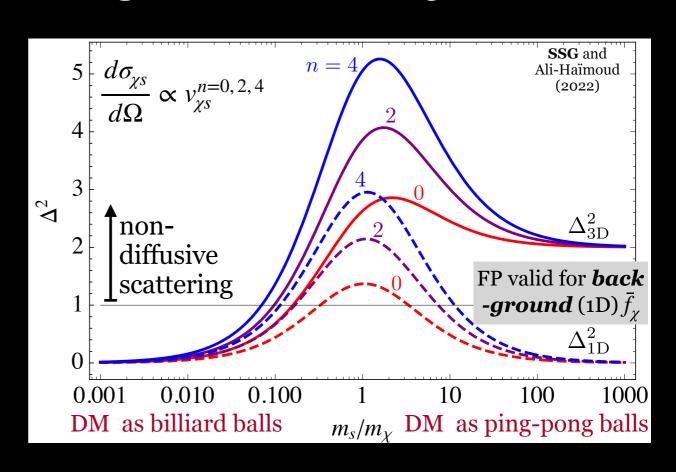
# Results: Comparing $\bar{f}_{\chi}^{\text{MB}}$ and $\bar{f}_{\chi}^{\text{FP}}$ to $\bar{f}_{\chi}^{\text{ex}}$

We can understand the  $\dot{Q}_{\chi}$  results by comparing the different  $\bar{f}_{\chi}$ :



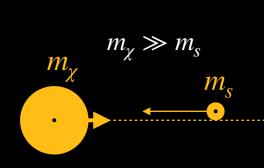
### Is FP accurate for $m_{\chi} \ll m_s$ ? - 3D vs. 1D diffusion

- 3D diffusion requires small change in DM vel. vector  $\overrightarrow{v}$
- 1D diffusion only requires small change in DM vel. magnitude v
- Showing coefficients that quantify diffusivity of 3D & 1D scattering for given  $m_s/m_\chi$
- Scattering is non-diffusive for  $m_s \sim m_\chi$  in **both** 1D and 3D
- FP accuracy  $\lesssim 17\%$  in 1D for  $m_s \sim m_\chi$  bodes well for 3D

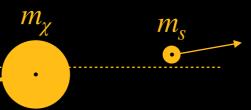


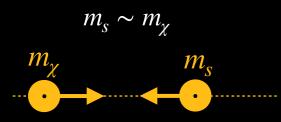
3D diffusion is relevant for CMB analyses in the presence of cosmological perturbations

## Takeaways



Diffusive scattering: small change in  $\overrightarrow{v}_{\chi}$ 





Is scattering diffusive?— No

Which is more accurate, FP approx. or MB assumption?

- FP has higher accuracy

- The FP approximation is valid for heavier DM,
- But even when it is physically invalid, it recovers the correct  $\dot{Q}_{\chi}$  within ~17%—

  10x better than MB in 1D
- Indicative of **better accuracy with** FP, even in 3D—relevant for CMB constraints on  $\sigma_{\gamma s}$
- Integrating FP formalism (as opposed to exact) with CMB codes is a more tractable task
- Importantly, FP approx. would *allow for DM self-scattering* (not possible with MB, and intractable with exact)

# Thank you for watching!

### My research interests include:

- Effects of DM-baryon interactions on cosmological observables
- (Relatively) Late-time ( $z \leq 10^3$ )  $\chi$ -s interaction models
- Reassessing the standard mathematical framework for CMB constraints on  $\sigma_{\gamma s}$  (besides the MB assumption)
- Stellar and galactic dynamics

Comments, questions, and any feedback is very welcome!

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### Fokker-Planck Method

- Computes the correct heating and momentum-exchange rates for a given  $f_{\chi}$
- However, the  $f_{\chi}$  that the FP collision operator determines is only an approximation for the true  $f_{\chi}$ 
  - so the resultant rates might not be accurate
- Turns CBE into PDE
- Discretization leads to sparse, tri-diagonal collision matrix, as opposed to full.

# Mass-dependence of FP validity

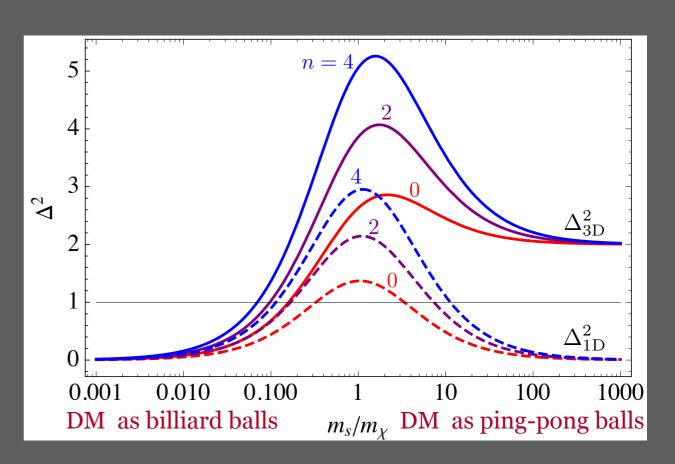
$$\Delta_{\mathrm{3D}}^{2}(v^{2}) \equiv \frac{\left\langle \sigma_{\chi s}(v_{\chi s}) \, v_{\chi s} \, \middle| \, \overrightarrow{v}' - \overrightarrow{v} \, \middle|^{2} \right\rangle}{\left\langle \sigma_{\chi s}(v_{\chi s}) v_{\chi s} \right\rangle v^{2}}$$

$$\langle (\vec{v}' - \vec{v})^2 \rangle_{\hat{n}'} = 2 \left( \frac{m_s}{M} \right)^2 |\vec{v} - \vec{v}_s|^2$$

Assuming fwd-bckwd scattering symmetry, we can avg out the  $\overrightarrow{v}'$  dependence

$$\Delta_{1D}^{2}(v^{2}) \equiv \frac{\left\langle \sigma_{\chi s}(v_{\chi s}) v_{\chi s} \left( |\overrightarrow{v}'|^{2} - |\overrightarrow{v}|^{2} \right) \right\rangle}{\left\langle \sigma_{\chi s}(v_{\chi s}) v_{\chi s} \right\rangle v^{2}}$$

$$\langle v'^2 - v^2 \rangle_{\hat{n}'} = 2 \left( \frac{m_s}{M} \right)^2 |\vec{v} - \vec{v}_s|^2 - 2 \frac{m_s}{M} (\vec{v} - \vec{v}_s) \cdot \vec{v}$$



By triangle inequality,

$$\Delta |\overrightarrow{v}| \le |\Delta \overrightarrow{v}| \Rightarrow \Delta_{1D}^2 \le \Delta_{3D}^2$$

### Simplifying $\Gamma_{\chi s}$

[which of these is the first simplification that allows us to factorize  $\Gamma_{\chi s}$  into (time dep.)  $\times$   $\Gamma$ ?]—the fact that diff cross-sec is a power-law in  $v_{\chi s}$ ! Doesn't have to be isotropic—it can also be a func of  $\hat{n} \cdot \hat{n}'$ . We only need to be able to pull out the factors of  $T_b/m_s$  (or  $T_b/m_{\chi}$ )

• Start with the full 3D expression:

$$\Gamma(\boldsymbol{v} \to \boldsymbol{v}') = \int d^3 v_s \ f(\boldsymbol{v}_s) \ v_{\chi s} \ \int d^2 \hat{n}' \ \frac{d\sigma}{d^2 \hat{n}'} \ \delta^{(3)} \left[ \boldsymbol{v}' - \boldsymbol{v} - R(v_{\chi s} \hat{n}' - \boldsymbol{v}_{\chi s}) \right], \qquad \boldsymbol{v}_{\chi s} \equiv \boldsymbol{v} - \boldsymbol{v}_s, \quad R \equiv \frac{m_s}{M}$$

• Assume (any) isotropic differential cross-section, and split  $\delta_D^{(3)}$  into  $\delta_D^{(2)}$  and  $\delta_D^{(1)}$ . The integral simplifies to 3D, with a 1D  $\delta_D$  function:

$$\Gamma(\boldsymbol{v} \rightarrow \boldsymbol{v}') = \frac{\sigma_n}{2\pi R} \int d^3v_{\chi s} f(\boldsymbol{v} - \boldsymbol{v}_{\chi s}) v_{\chi s}^n \ \delta^{(1)}[v^2 + {v'}^2 - 2R\boldsymbol{v} \cdot \boldsymbol{v}_{\chi s} + 2(R\boldsymbol{v}_{\chi s} - \boldsymbol{v}) \cdot \boldsymbol{v}'].$$

Now if  $f_s(\overrightarrow{v}_s)$  is isotropic in one unique frame,  $\Gamma(\overrightarrow{v} \to \overrightarrow{v}')$  only on  $\left| \overrightarrow{v} - \overrightarrow{V}_b \right|$ ,  $\left| \overrightarrow{v}' - \overrightarrow{V}_b \right|$ , and  $\hat{v} \cdot \hat{v}'$ .

This is useful if we want to find  $\int d^2v' \int \frac{d^2v}{4\pi} \Gamma(\overrightarrow{v} \to \overrightarrow{v}')$ , as it renders one of these integrals redundant.

- $\int d^2v' \int \frac{d^2v}{4\pi} \Gamma(\overrightarrow{v} \to \overrightarrow{v}')$  can then be reduced to a 1D integral over  $v_{\chi s}$ .
- For  $\frac{d\sigma_{\chi s}}{d^2\hat{n}'} = \frac{\sigma_n v_{\chi s}^n}{4\pi}$ , it is reducible to a completely analytical expression for  $n \in \{0, 2, 4, \dots\}$

# Solving the 1D CBE exactly

• For the first time, we find fully analytical expressions for  $G_{\chi s}^{1D}(v \to v')$  with  $G_{\chi s}^{2\hat{n}'} \propto v_{\chi s}^n$ ,  $n \in \{0, 2, 4, ...\}$ 

• Recasting velocities as  $u \equiv \sqrt{m_\chi/T_b} v$ , we can factorize the problem:

(Collision = 
$$\min_{\substack{\text{Operator}}} \tilde{\mathbf{M}}_{ij}$$
 (collision =  $\max_{\substack{\text{Operator}}} \mathbf{X} f(u_i) \times \text{time dep. factor}$  scattering matrix)

• Convert *integro-differential CBE* into a set of *coupled ODEs*—much simpler to solve numerically

### Steps from integro-diffrn. eq to coupled ODEs

#### Integro-differential equation:

$$a^{1/2} \left. \frac{d}{dt} \right|_{\text{free}} \left[ a^{-1/2} \widetilde{f}(u) \right] = R_n(T_b) \ \widetilde{C}[\widetilde{f}](u), \quad \Longrightarrow \quad \frac{\partial \Delta \widetilde{f}}{\partial \ln a} = \frac{R_n(T_b)}{H} \widetilde{C}[\Delta \widetilde{f}] + \frac{1}{2} \frac{\partial}{\partial u} \left[ u(\widetilde{f}^{\text{eq}} + \Delta \widetilde{f}) \right].$$

#### Define rescaled time var., $y = a/a_{\gamma b}$ and get:

$$\frac{\partial \Delta \widetilde{f}}{\partial \ln y} = y^{-\frac{n+3}{2}} \widetilde{C}[\Delta \widetilde{f}] + \frac{1}{2} \frac{\partial}{\partial u} \left[ u(\widetilde{f}^{\mathrm{eq}} + \Delta \widetilde{f}) \right].$$

#### 1D collision operator:

$$\widetilde{C}[\Delta \widetilde{f}](u_i) = \sum_{j=0}^{N-1} M_{ij} \Delta \widetilde{f}_j,$$

$$M_{ij} \equiv d \ln u \left( \widetilde{\Gamma}_{ji} u_j - \delta_{ij} \sum_k \widetilde{\Gamma}_{ik} u_k \right)$$

#### Final eq. with coupled ODEs:

$$\frac{\partial \Delta \widetilde{f}_i}{\partial \ln y} = y^{-\frac{n+3}{2}} \sum_{j=0}^{N-1} M_{ij} \Delta \widetilde{f}_j$$
$$+ \frac{1}{2} \sum_{j=i-1}^{i+1} \alpha_{ij} (\widetilde{f}_j^{eq} + \Delta \widetilde{f}_j)$$

$$a_{\chi b}^{rac{n+3}{2}} \equiv 2 c_n \sigma_n rac{n_{s,0}}{H_0 \Omega_r^{1/2}} \left(rac{M^2}{m_s m_\chi}
ight)^{rac{n-1}{2}} \left(rac{T_{r,0}}{M}
ight)^{rac{n+1}{2}}$$

$$\tilde{f}(u) \equiv \sqrt{T_b/m_{\chi}} f^{\text{1D}}(v)$$

$$\Delta \tilde{f}(u) \equiv \tilde{f}(u) - \tilde{f}^{\text{eq}}(u)$$

#### Numerical gradient operator:

$$\begin{split} \frac{\partial(u\widetilde{f})}{\partial u}\Big|_{i} &= \frac{1}{2u_{i}d\ln u} \\ &\times \begin{cases} (u_{0}\widetilde{f}_{0} + u_{1}\widetilde{f}_{1}) & i = 0 \\ \left(u_{i+1}\widetilde{f}_{i+1} - u_{i-1}\widetilde{f}_{i-1}\right) & 1 \leq i \leq N-2 \\ -(u_{N-2}\widetilde{f}_{N-2} + u_{N-1}\widetilde{f}_{N-1}) & i = N-1 \end{cases} \\ &\equiv \sum_{j=i-1}^{i+1} \alpha_{ij}\widetilde{f}_{j}, \quad \text{where} \quad \alpha_{0,-1} = \alpha_{N-1,N} = 0. \end{split}$$

## Equations for $a_{dec}$ , $Q_{\nu}^{MB}$

$$| \text{If } f = f^{MB} \Big|_{T_{\chi}^{MB}} :$$

$$\dot{\mathcal{Q}}_{\chi} = \frac{\rho_{\chi}\rho_s}{M^2 v_{\rm th}^2} \mathcal{B}(V_{\chi b}; v_{\rm th}^2) \times (T_b - T_{\chi}) + \frac{\rho_{\chi}\rho_s}{M v_{\rm th}^2} \frac{T_{\chi}}{m_{\chi}} \mathcal{A}(V_{\chi b}; v_{\rm th}^2) V_{\chi b}^2.$$

$$\mathcal{A}(w; T/m) = c_n \ \sigma_n \left(\frac{T}{m}\right)^{\frac{n+1}{2}} \alpha_n(\sqrt{m/T} \ w), \quad (36)$$

$$\mathcal{B}(w; T/m) = 3c_n \ \sigma_n \left(\frac{T}{m}\right)^{\frac{n+3}{2}} \beta_n(\sqrt{m/T} \ w), \ (37)$$

$$\alpha_n(x) \equiv {}_{1}F_1\left(-\frac{n+1}{2}, \frac{5}{2}, -\frac{x^2}{2}\right),$$
 (38)

$$\beta_n(x) \equiv {}_{1}F_1\left(-\frac{n+3}{2}, \frac{3}{2}, -\frac{x^2}{2}\right),$$
 (39)

$$c_n \equiv \frac{2^{\frac{5+n}{2}}}{3\sqrt{\pi}}\Gamma(3+n/2). \tag{40}$$

#### Ali-Haïmoud 2019

$$\dot{T}_{\chi} = -2HT_{\chi} + \Gamma_{\chi b} \left( T_b - T_{\chi} \right), \tag{1}$$

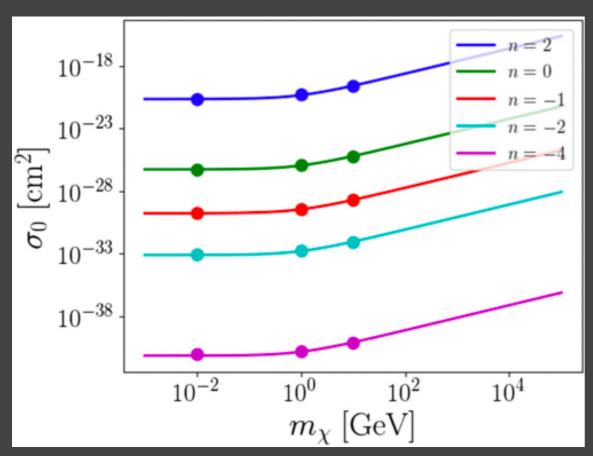
with 
$$\Gamma_{\chi b} \equiv \frac{2c_n N_b \sigma_n m_b m_{\chi}}{(m_b + m_{\chi})^2} \left(\frac{T_b}{m_b} + \frac{T_{\chi}}{m_{\chi}}\right)^{\frac{n+1}{2}}$$
, (2)  $\Gamma_{\chi b}^{eq(2)} \sim \frac{m_{\chi}}{m_{\chi} + m_b} \Gamma_{\chi b, tot} \propto v^{n+1}$ 

$$\Gamma_{\chi b}^{eq(2)} \sim \frac{m_{\chi}}{m_{\chi} + m_b} \Gamma_{\chi b, tot} \propto v^{n+1}$$

$$\frac{\Gamma_{\chi b}}{H} = \left(\frac{a_{\chi b}}{a}\right)^{\frac{n+3}{2}} \left(\frac{m_{\chi}/m_b + T_{\chi}/T_b}{m_{\chi}/m_b + 1}\right)^{\frac{n+1}{2}}, \quad (3)$$

$$(a_{\chi b})^{\frac{n+3}{2}} \equiv \frac{m_b}{m_{\chi}} \left( 1 + \frac{m_b}{m_{\chi}} \right)^{\frac{n-3}{2}} \frac{2c_n \sigma_n N_b^0 \left( \frac{T_{\gamma}^0}{m_b} \right)^{\frac{n+1}{2}}}{H_0(\Omega_r^0)^{1/2}}.(4)$$
 Ali-Haïmoud et al 2015

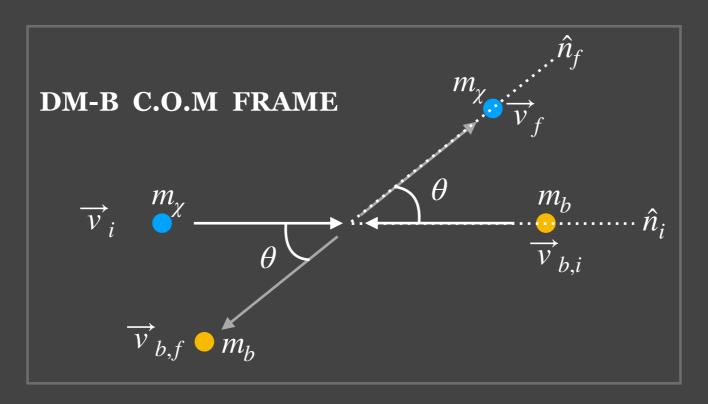
### CMB+Ly $\alpha \sigma_0(m_{\gamma})$ bounds (n = -4, -2, -1, 0, 2)



Xu et. al, PRD 97, 103530, 2018

- Flat profile:  $m_{\chi} \ll m_b$ ,  $R_{\chi} \propto \sigma_0 (v_{th})^{n+1}/M \propto \sim \sigma_0 v_b^{n+1}$  down to  $m_{\chi} \sim 10 MeV$  bec  $T_{\chi}/T_b \ll m_{\chi}/m_b$  still
  - $m_{\chi}$  dep. drops out
  - $\sigma_0 \propto m_{\chi}$  profile: For  $m_{\chi} \gg m_b$ ,  $R_{\chi} \propto \sigma_0 (v_{th})^{n+1}/M \propto \sim \sigma_0 v_b^{n+1}/M$ 
    - $R_{\chi} \propto \sigma_0/m_{\chi}$

# DM-baryon $(\chi-b)$ Scattering



Relevant quantities (for ex.):

- 1. Rate of scattering for DM:  $\Gamma_{\chi b}(f_{\chi}(\overrightarrow{v}), \sigma_{\chi b})$ ,  $v_{\chi b} = |\overrightarrow{v}_{\chi} \overrightarrow{v}_{b}|$
- 2. DM Heat exchange rate:  $\dot{Q}_{\chi}(f_{\chi}(\vec{v}), \sigma_{\chi b}) \propto \dot{T}_{\chi}^{scatt}$
- 3. DM momentum exchange rate:  $\overrightarrow{p}_{\chi}(f_{\chi}(\overrightarrow{v}), \sigma_{\chi b}) = m_{\chi} \overrightarrow{V}_{\chi}$