#### Machine Learning for Cosmology

#### **Michelle Ntampaka**



# Astronomy is the ideal sandbox for machine learning.

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- Minimal privacy concerns.
- Culture of sharing data.
- Cosmological surveys are rich with information.
- Well-posed questions.
- Public interest and support.
- Data are non-monetizable.
- This does not exempt us from ethical concerns!

# Is machine learning the right tool for astronomy?

# Is machine learning the right tool for astronomy?

- 1. Can ML be trusted?
- 2. Can ML be used to make physical discoveries?
- 3. What role will machine learning play in the future of astronomy?

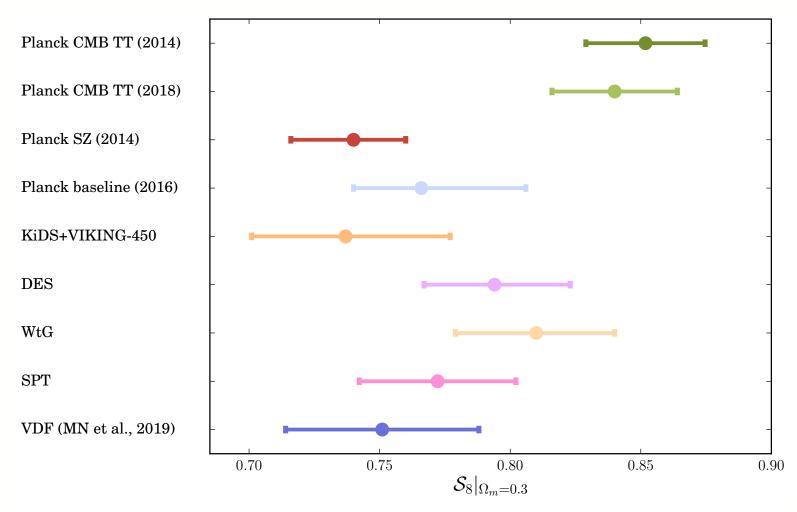
A Case Study in Interpretability: Galaxy Clusters as a Cosmological Probe





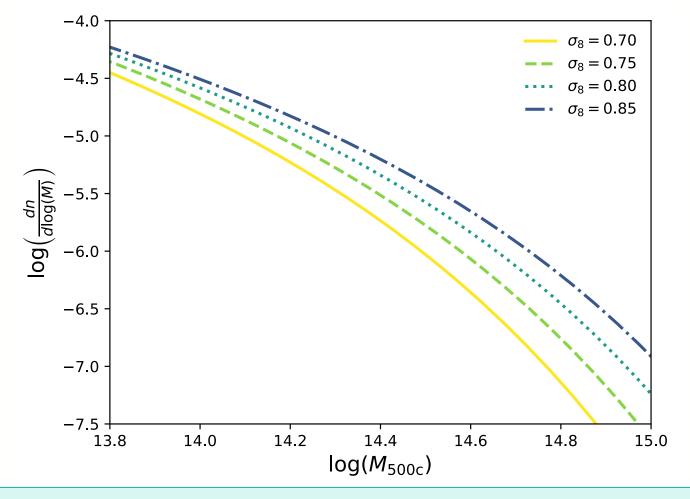
X-ray image credit: NASA/CXC/MIT/E.-H Peng et al; Optical image credit: NASA/STScl

## Tensions in the current cosmological model: $S_8$ (CMB vs. LSS)



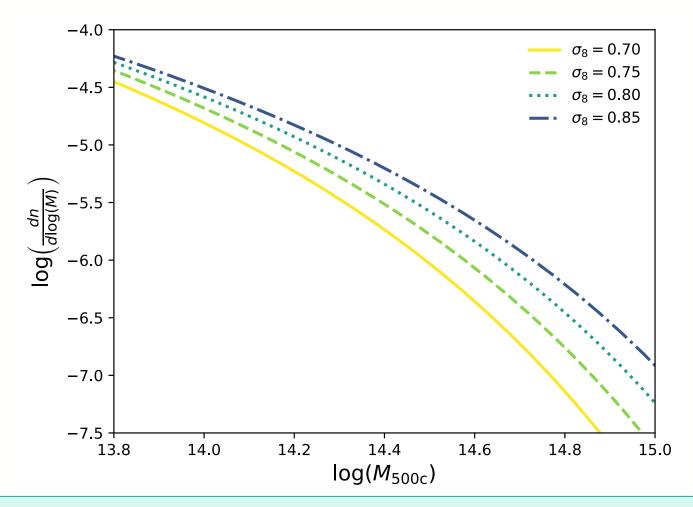
#### Ntampaka+ 2019b

#### **Clusters as a Cosmological Probe**



Cluster abundance is sensitive to the underlying cosmology (especially to  $\sigma_8!$ )

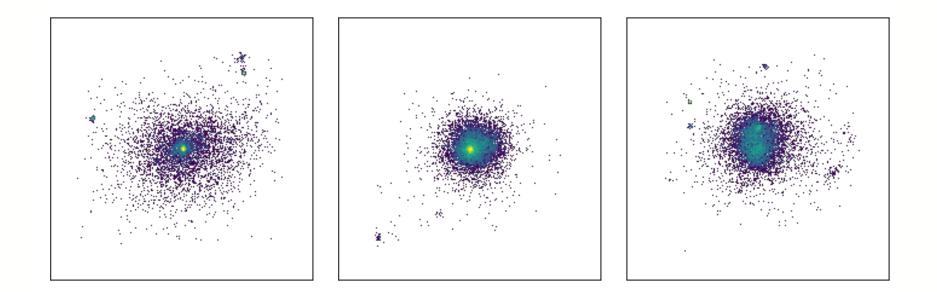
#### Mass Bias vs. Cosmology



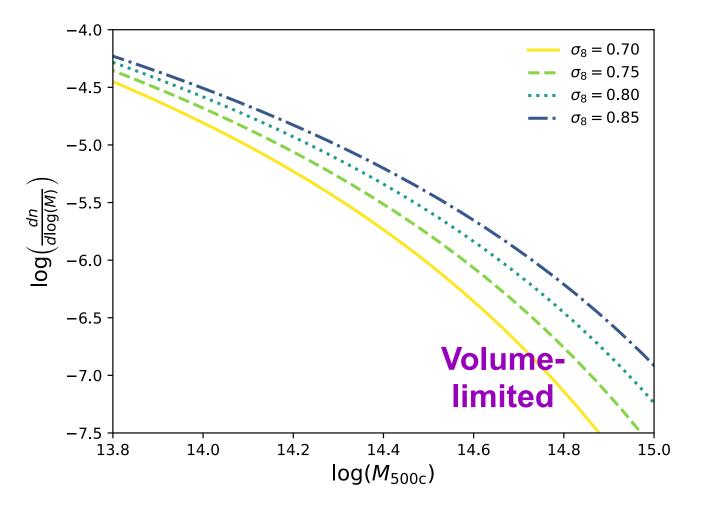
It is difficult to discern between biased cluster masses and a change in cosmology (WL masses can help us with mass calibration).

### Simulated Galaxy Clusters

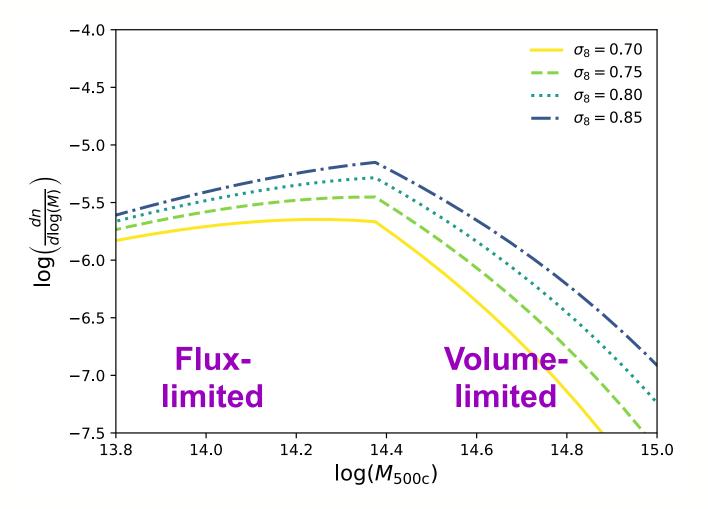
- Magneticum Cosmological Hydrodynamical Simulation
- WMAP7 Cosmology
- Box length =  $352h^{-1}$  Mpc
- Chandra-like mock observation



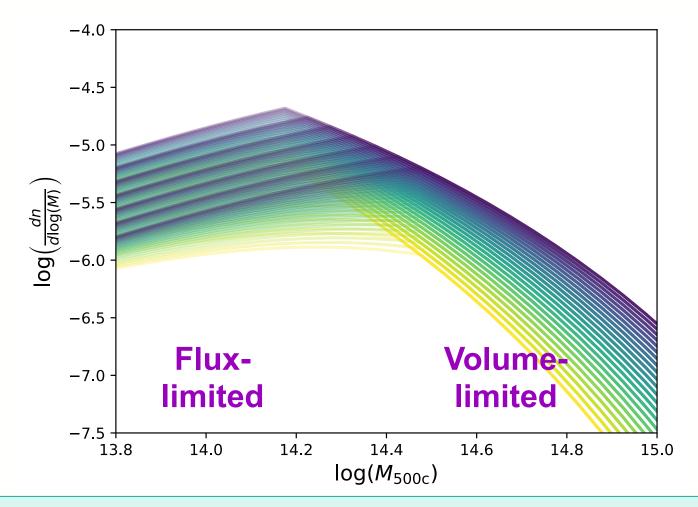
#### Many Cosmologies from One Simulation



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#### Many Cosmologies from One Simulation



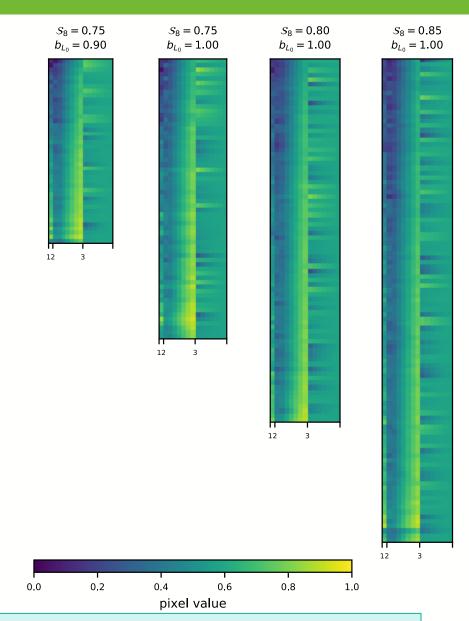
Typical X-ray surveys include both flux- and volume-limited regimes, and also uncertainty in the mass at which these two join.

### Simulated Cluster Samples

- Pixel = cluster observable
- Row = cluster
- Image = simulated cluster sample

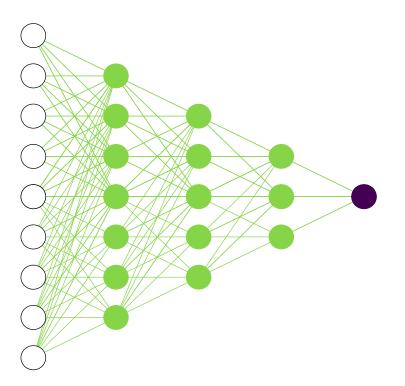
#### Pixels are cluster observables:

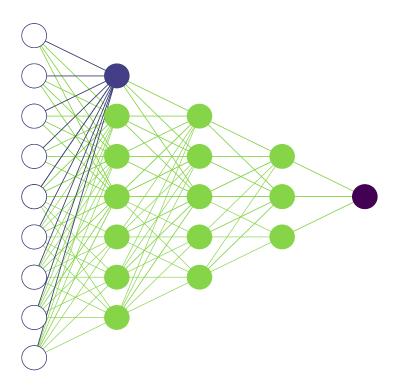
- 1. Temperature
- 2. Gas mass profile
- 3. Gas density slope profile

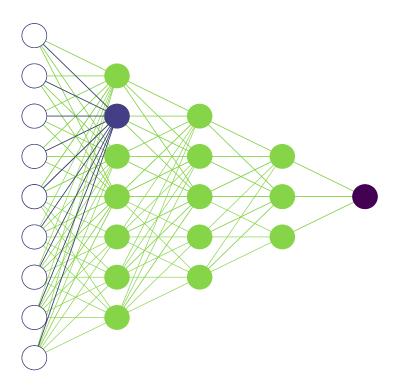


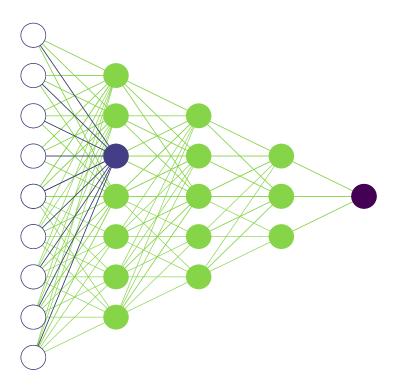
We have simulated realistic cluster samples for many cosmologies & have cast each into a 2D array.

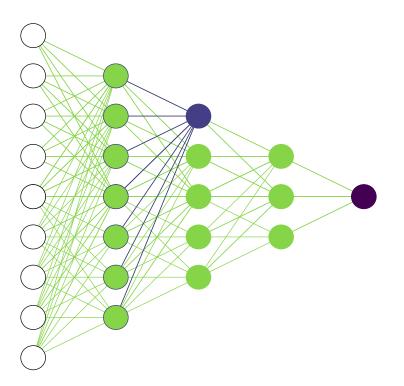
## ML tools: Neural Networks & Encoders

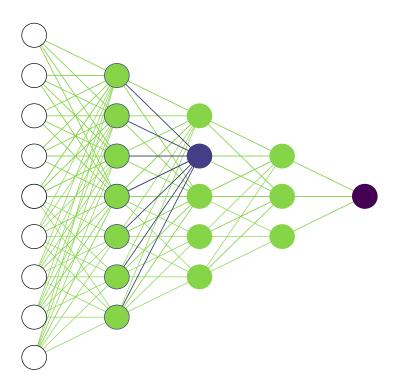


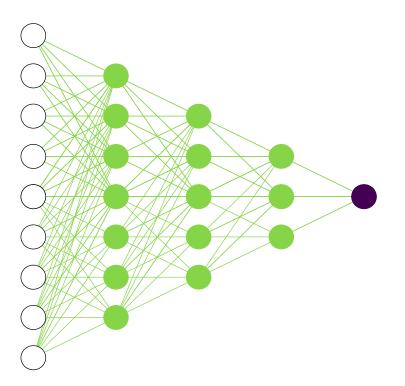




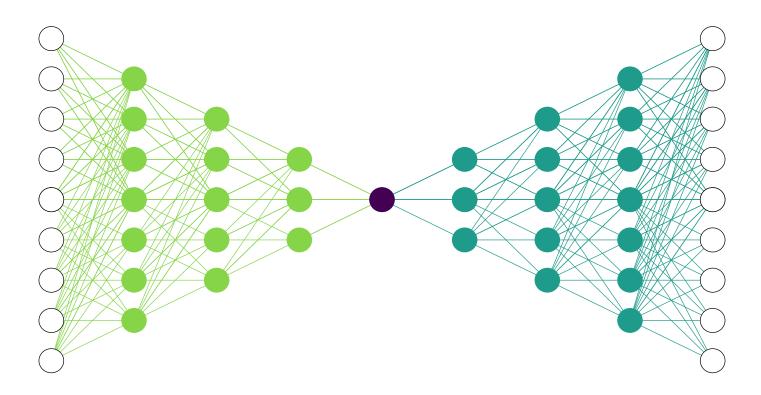




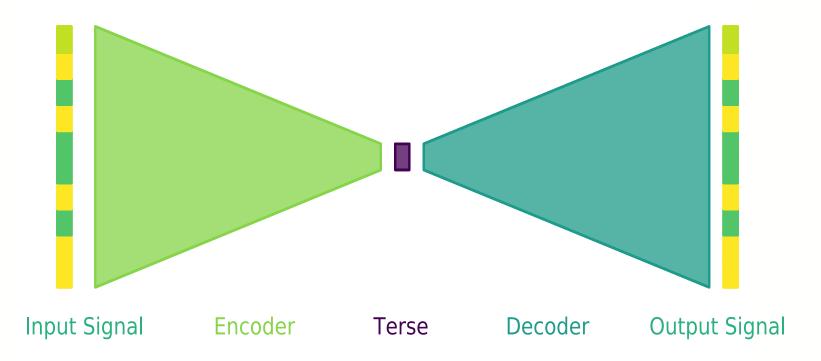




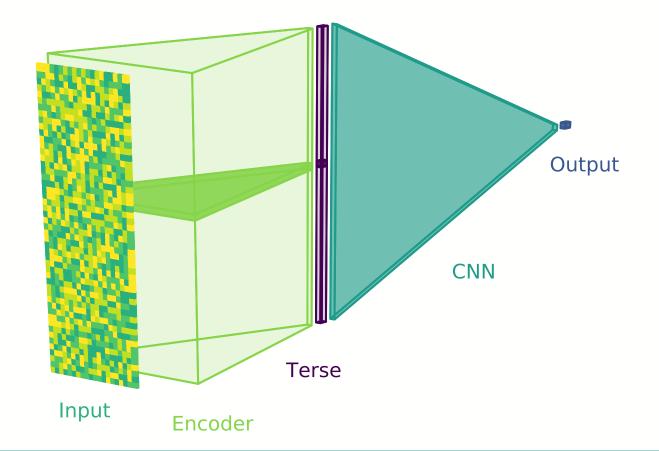
#### Autoencoder



#### Autoencoder

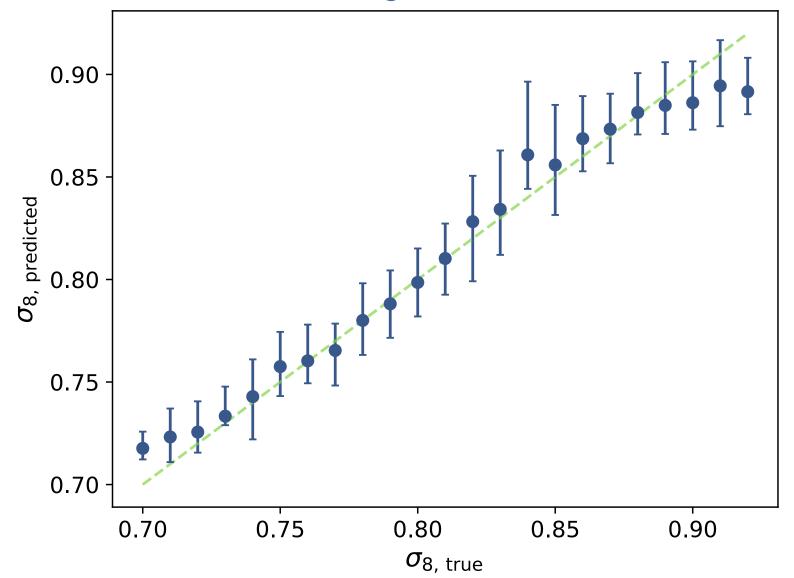


#### Supervised Encoder



This encoder is engineered to mimic a human approach to the task: cluster observables → cluster mass proxy list of mass estimates → cosmology

#### Results: $\sigma_8$ Predictions



## Interpretation

### Can ML be trusted? Can it drive physical understanding?

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1. Terse Value Correlations: to assess whether the model will generalize. These can help us to <u>trust</u> in our ML models.

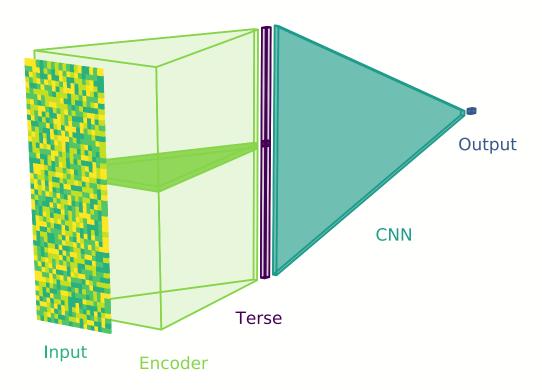
2. Saliency maps: identify what part of the cluster carries the most cosmological information. These can lead to ML-driven <u>discoveries</u>.

## 1. Correlations -> Trust

#### **Terse Value Correlations**

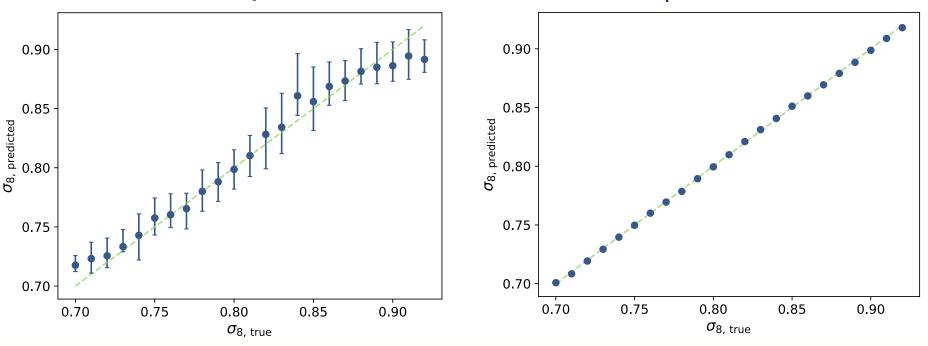
Big idea: visualize how the terse layer correlates with cluster parameters.

**Build trust** in the model by asking "Is the terse layer summarizing sensible features?"



Trustworthy Results:

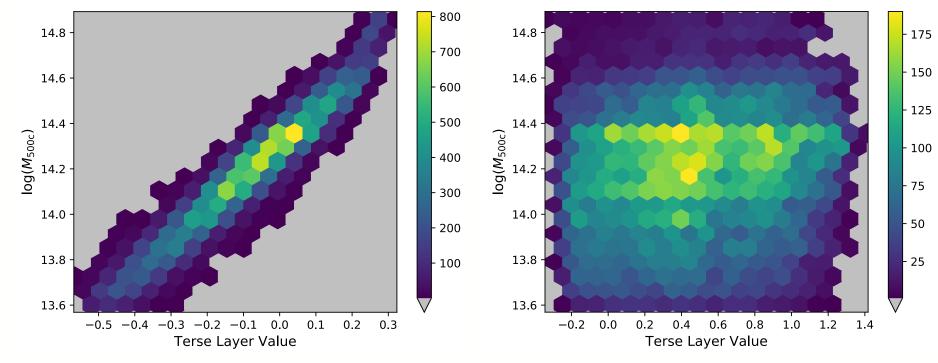
Suspicious Results:



Here, we find a factor-of-10 better results. But is it trustworthy? Will it generalize to real data?

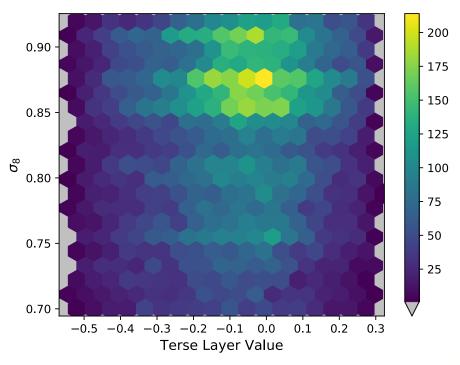
Trustworthy Results:



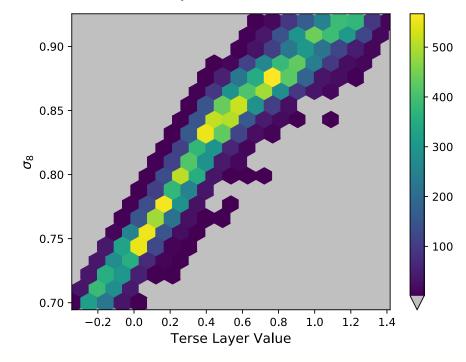


Why is the terse value uncorrelated with mass? This is a red flag that something isn't right.

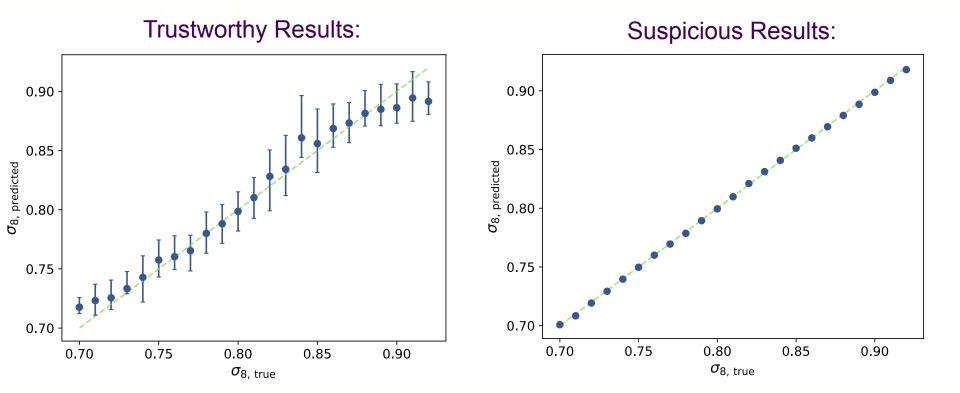
Trustworthy Results:



#### Suspicious Results:



Here's the problem. The model has picked up on a cosmology-dependent simulation artifact and is able to infer  $\sigma_8$  from just one cluster.

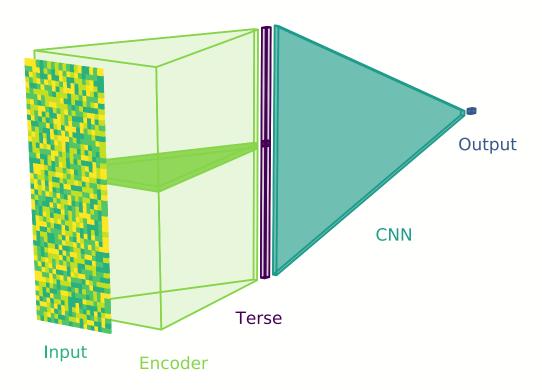


The verdict? For the "Suspicious Results," ML cheated. The "Suspicious Results" are not robust. The model will not generalize to real observations because it depends on a simulation artifact.

#### **Terse Value Correlations**

Big idea: visualize how the terse layer correlates with cluster parameters.

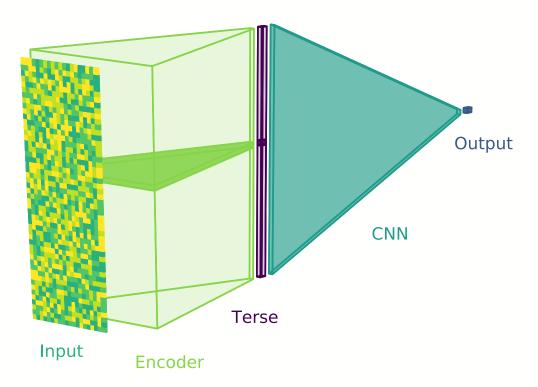
**Build trust** in the model by asking "Is the terse layer summarizing sensible features?"



# 2. Saliency -> Discovery

Big idea: assess the importance of each cluster feature by looking at gradients in the encoder network. "How does changing the input change the terse value?"

Develop a physical framework for understanding surprising results.



Simonyan+ 2014



How can you tell that this is a horse? How can ML tell that this is a horse?



Horse-picture from Pascal VOC data set

#### Source tag present Classified as horse No source tag present the present the present the present the present the present

Lapuschkin et al., 2019

Artificial picture of a car

## Input

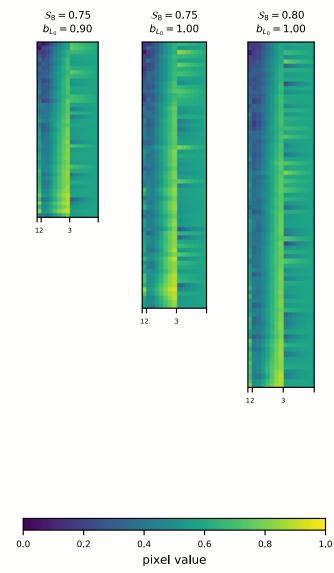
 $S_8 = 0.85$ 

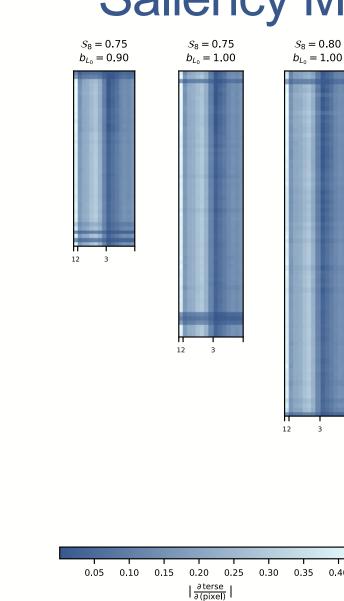
 $b_{L_0} = 1.00$ 

Π

12

3





## **Saliency Maps**

3

0.40

Π

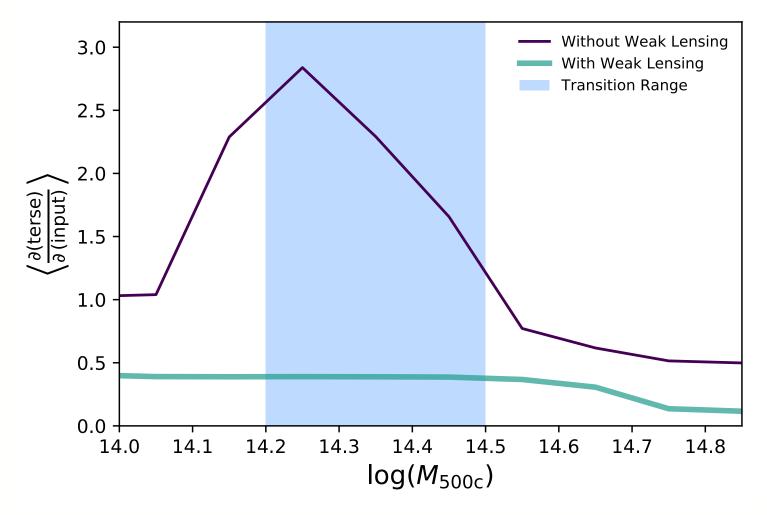
12

3

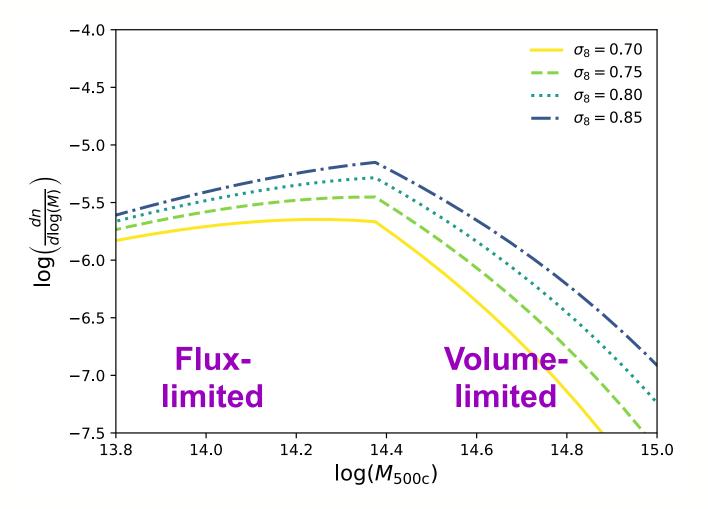
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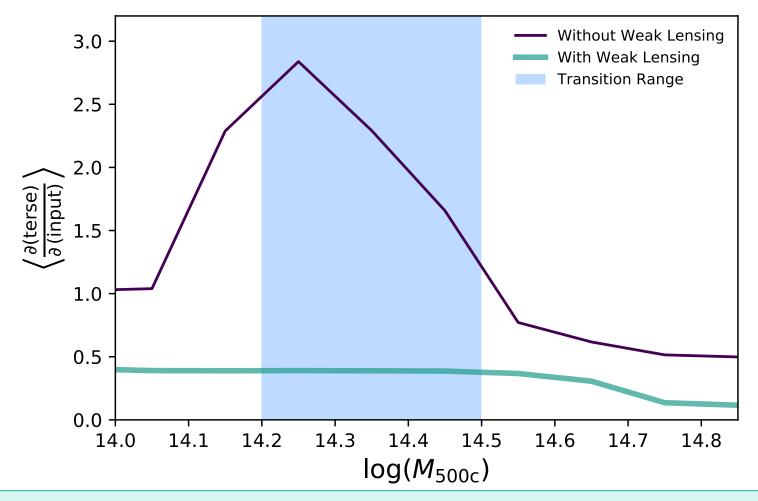
#### Saliency Trend: A Smoking Gun



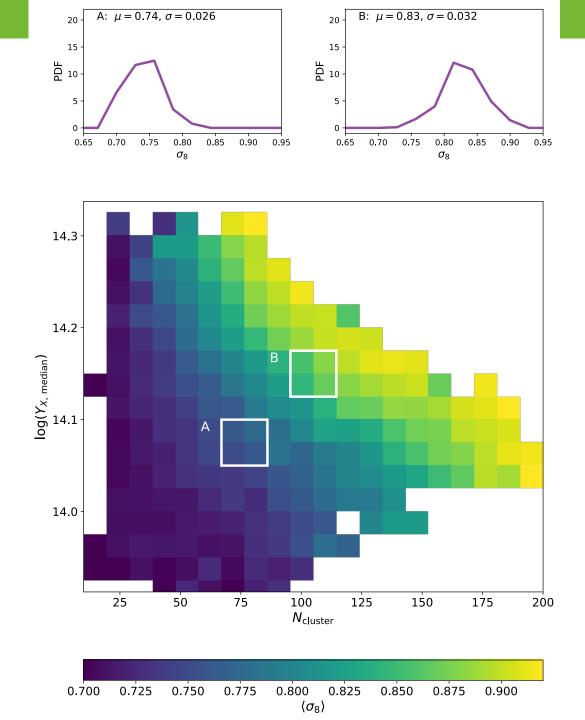
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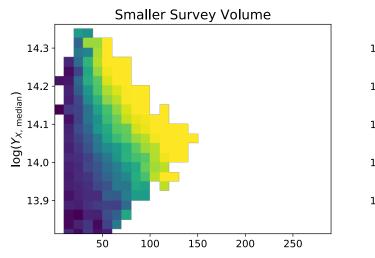
## Saliency Trend: A Smoking Gun

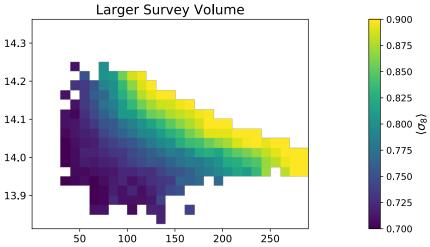


Is the model finding the transition point between the flux- and volumelimited samples to calibrate cluster masses without weak lensing?



#### Survey-Dependent





0.06

0.05

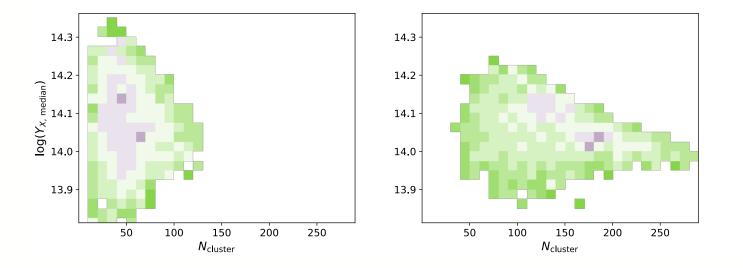
0.04

0.02

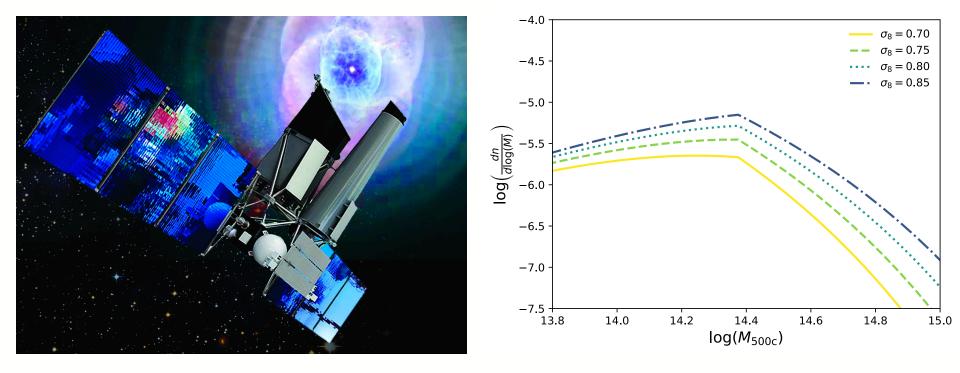
0.01

0.00

 $\operatorname{std}_{\operatorname{c0.0}}^{\operatorname{c0.0}}$ 

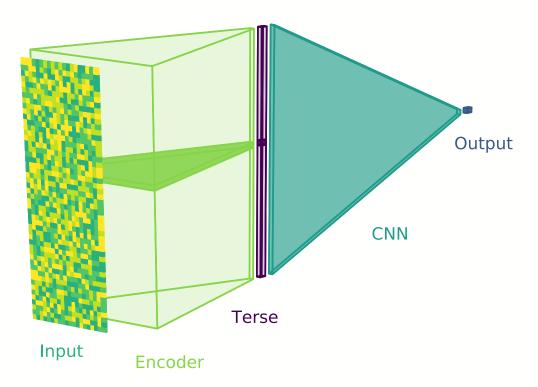


# ML-aided discovery of a self-calibration mode for eROSITA



Big idea: assess the importance of each cluster feature by looking at gradients in the encoder network. "How does changing the input change the terse value?"

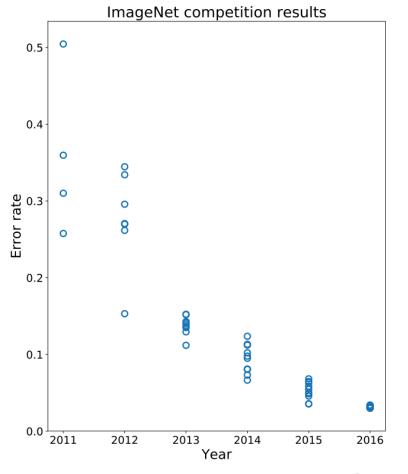
Develop a physical framework for understanding surprising results.



Simonyan+ 2014

# What role will ML play in the future of astronomy?

# ML & Astronomy can – and should! – move forward together.



Plot credit: Gkrusze, public domain

# ML & Astronomy can – and should! – move forward together.



archive.stsci.edu/hello-universe/



Home

#### A framework for testing and benchmarking machine learning methods on astronomical data

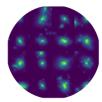
*Hello Universe* is a new project at MAST designed to help astronomers develop machine learning (ML) methods for astronomical discovery. ML will be an essential tool for analyzing the rich data sets of the upcoming decade, and *Hello Universe* provides a framework for testing ML algorithms and new techniques. Each entry in the *Hello Universe* collection includes:

- Data: a high-level science product (HLSP) data set for testing and benchmarking ML algorithms
- Code: a tutorial Jupyter notebook that provides step-by-step examples of how to apply an ML technique to the data

Though these data sets are motivated by the needs of a novice data science learner, they are sufficient for a wide range of tasks. *Hello Universe* entries include examples of:

- analyzing 2D (image) and 1D (vector or light curve) data sets.
- applying techniques for regression and for classification.
- developing supervised and unsupervised learning models.
- using best practices for training and optimizing models.
- selecting metrics for assessing model performance.





#### Classifying JWST/HST galaxy mergers with CNNs neural networks | 2d data classification | overfitting | confusion matrix



#### Classifying TESS stellar flares with CNNs neural networks | 1d data |

neural networks | 1d data | classification | prediction



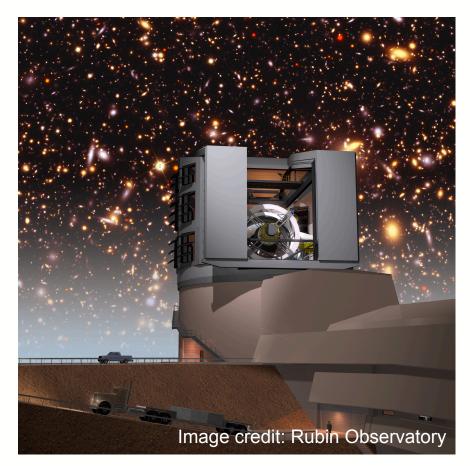
Predicting 3D-HST redshift with decision trees decision trees | 1d data | regression | cross-validation



Classifying Pan-STARRS with (un)supervised learning classification | 1d data | PCA | tSNE | k-means | SGD |

unsupervised supervised

#### **Rich Upcoming Data Sets**







Roman image credit: NASA (public domain)

#### Style Transfer









Image credit: Google AI blog, adapted from Gatys+ 2015

## **Pixel Recursive Super Resolution**

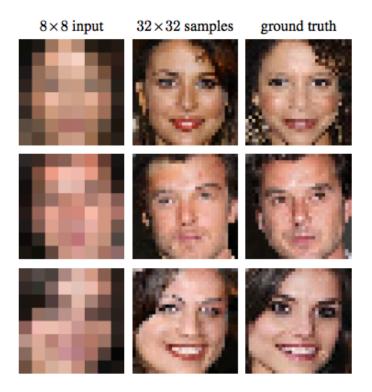


Figure 1: Illustration of our probabilistic pixel recursive super resolution model trained end-to-end on a dataset of celebrity faces. The left column shows  $8 \times 8$  low resolution inputs from the test set. The middle and last columns show  $32 \times 32$  images as predicted by our model *vs*. the ground truth. Our model incorporates strong face priors to synthesize realistic hair and skin details.

#### Dahl+ 2017

#### **Bias in Training Data**

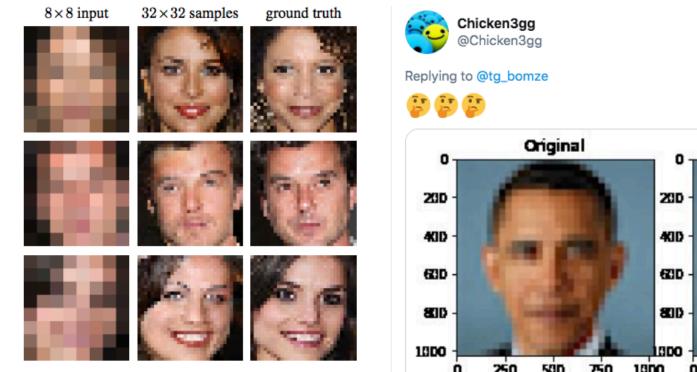
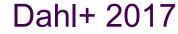


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Result 250 50D 250 50D 750 1000 750 1000 0 0

## **Bias in Language Translation**

Х

#### O bir doktor. O bir hemşire.

Translate from: Turkish

#### He is a doctor. She is a nurse.

Open in Google Translate · Feedback

Caliskan+ 2017

## **Automating Human Bias**

RETAIL OCTOBER 10, 2018 / 7:04 PM / UPDATED 4 YEARS AGO

# Amazon scraps secret AI recruiting tool that showed bias against women

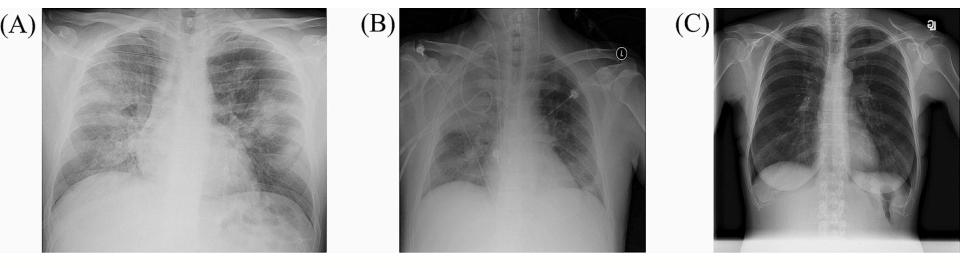
By Jeffrey Dastin

8 MIN READ

f ¥

SAN FRANCISCO (Reuters) - Amazon.com Inc's <u>AMZN.O</u> machine-learning specialists uncovered a big problem: their new recruiting engine did not like women.

#### **Undesirable Learning Behaviors**



Covid

#### Pneumonia

Healthy

image credit: Nishia et al., 2020

ARTIFICIAL	INTELLIGENCE
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# Hundreds of AI tools have been built to catch covid. None of them helped.

Some have been used in hospitals, despite not being properly tested. But the pandemic could help make medical AI better.

**By Will Douglas Heaven** 

July 30, 2021



# Is ML the right tool for astronomy?

# Machine Learning *can be* the right tool for astronomy:

- Engineer models that follow a human approach with checkpoints to make sure that the model is learning something sensible.
- Interrogate and interpret models.
- Approach high-accuracy results with scrutiny.
- Treat ML as a tool to be used in partnership with traditional statistical methods & human exploration.

#### 1. Can ML be trusted?

Terse value correlations: to assess whether the model will generalize. Building Trustworthy ML Models for Astronomy Ntampaka, Ho, & Nord 2021, 2111.14566

#### 2. Can it drive physical understanding?

Saliency maps: to identify what part of the cluster carries the most cosmological information.

The Importance of Being Interpretable Ntampaka & Vikhlinin 2022, 2112.05768

#### 3. What role will ML play in the future of astronomy?

ML in partnership – not in competition! – with traditional methods. **The Role of ML in the Next Decade of Cosmology Ntampaka+ 2019, 1902.10159** 

Hello Universe archive.stsci.edu/hello-universe

mntampaka@stsci.edu