

Quintessential Inflation in Palatini $f(R)$ Gravity



Samuel Sánchez López
Lancaster University

K. Dimopoulos, S.S.L. Phys.Rev.D 103 (2021) 4, 043533
[arXiv:**2012.06831**]

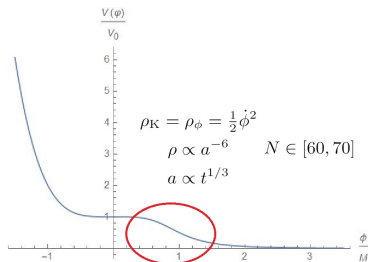
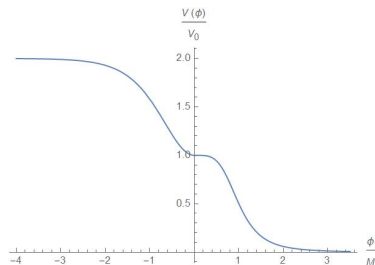
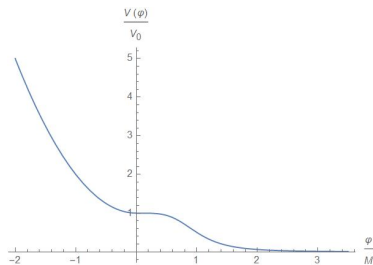


Figure: The potential of the original Peebles-Vilenkin quintessential inflation model [astro-ph/9810509].

Quintessential inflation identifies the inflaton and quintessence fields.

Advantages:

1. Economical approach (one single scalar field explains both the inflationary and dark energy epochs!).
2. Heavily constrained (easily falsifiable).
3. The initial conditions of quintessence are fixed by the inflationary attractor.



$$S_J = \int d^4x \sqrt{-g} \left[\frac{M_{\text{P}}^2}{2} R + \frac{1}{4} \alpha R^2 - \frac{1}{2} g^{\mu\nu} \nabla_\mu \varphi \nabla_\nu \varphi - V(\varphi) \right] + S_m[g_{\mu\nu}, \psi],$$

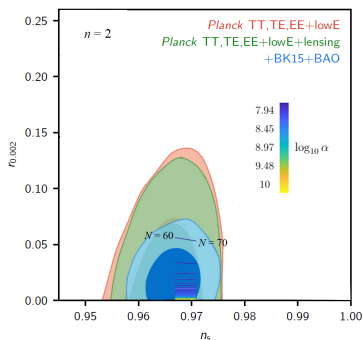


Figure: The predictions of our model superimposed on the Planck data [arXiv:2012.06831].

Planck data at 1σ :

1. $A_s = (2.096 \pm 0.101) \times 10^{-9}$
2. $n_s = (0.9661 \pm 0.0040)$
3. $r < 0.056$

Our data

for $N = 68$ and $\alpha = 8.7 \times 10^7$:

1. $m \sim 10^{13} \text{ GeV}$ ($\frac{1}{2}m^2 = \lambda^2 m_{\text{P}}^2$)
2. $n_s = 0.9708$
3. $r = 0.05$
4. Coincidence fixes $M \sim 10 \text{ GeV}$

$$\begin{aligned}
 V(\varphi) &= \frac{1}{2}m^2(\varphi^2 + M^2) \quad \varphi < 0, \\
 &= \frac{\frac{1}{2}m^2 M^6}{\varphi^4 + M^4}, \quad \varphi \geq 0.
 \end{aligned}$$