

QUANTUM FLUIDS MEETS COSMOLOGY

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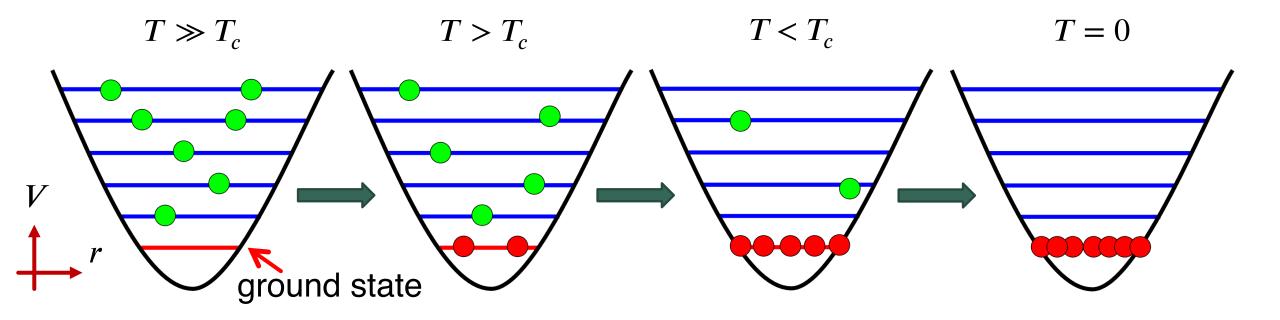
- Bose Einstein Condensation
- Vacuum Decay
- Analogue Quantum-Fluid System
- Numerical Model
- Equilibrium Results
- Bubble Growth
- Future Plans

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BOSE EINSTEIN CONDENSATION

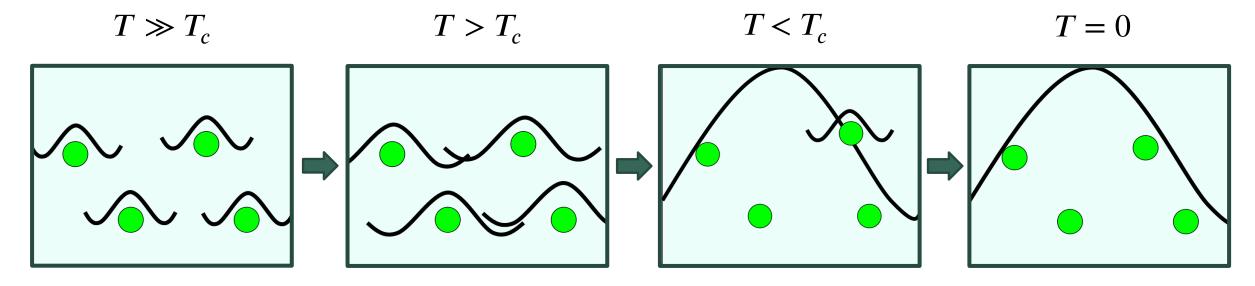
- Weakly interacting gas of identical bosons
- Bosons like to occupy the same state as one another
- Below some critical temperature atoms accumulate in the ground state



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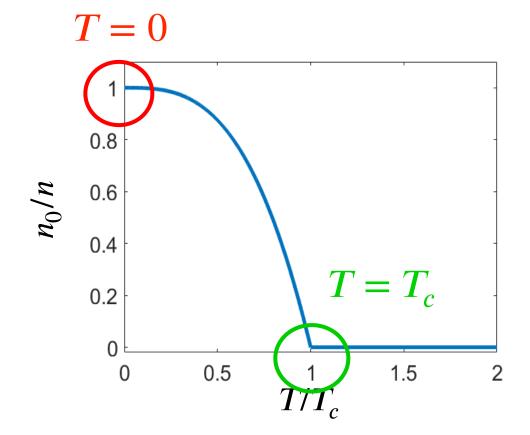
BOSE EINSTEIN CONDENSATION

- The de-Broglie wavelength: the wavelength of matter
- It gets longer as we cool
- Matter waves begin to overlap
- Bose Einstein Condensation: A collection of atoms in the ground state described by one wavefunction



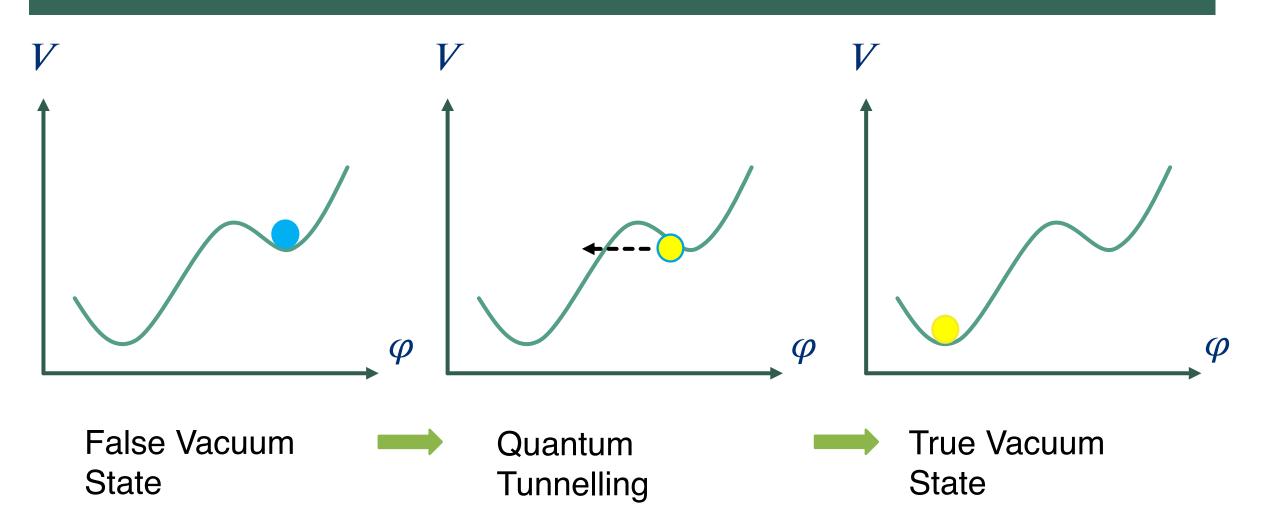
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BOSE EINSTEIN CONDENSATION



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VACUUM DECAY

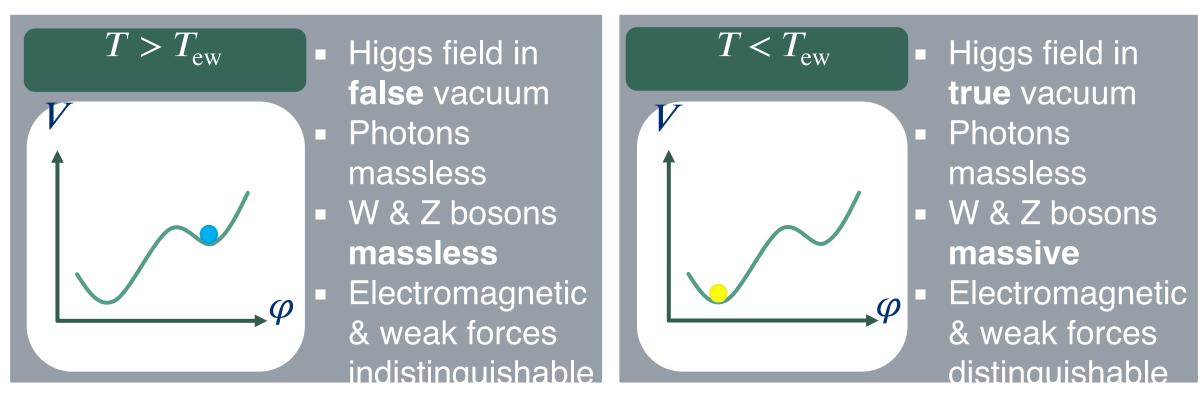






VACUUM DECAY – THE ELECTROWEAK PHASE TRANSITION

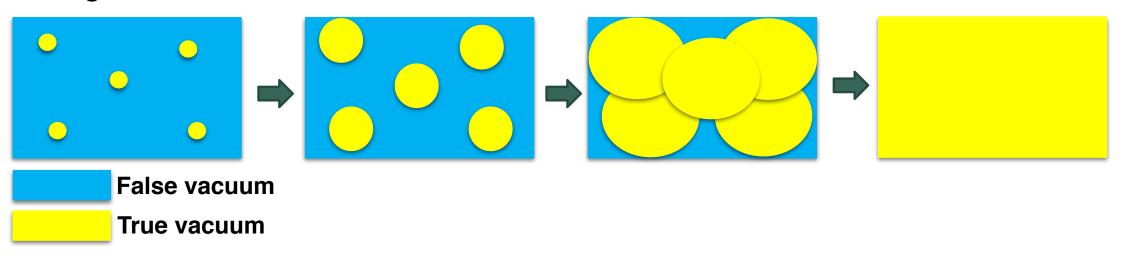
- Weak force and electromagnetic force differentiated
- Occurred ≈ 1 ns after the Big Bang, $T_{\rm ew} \approx 10^{15} {\rm K}$





VACUUM DECAY – THE ELECTROWEAK PHASE TRANSITION

- Standard Model: 2nd order, continuous
- Beyond the Standard Model: 1st order
- Universe converted from false vacuum to true vacuum via bubbles
- Observational consequences: baryon asymmetry, gravitational waves, magnetic fields



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ANALOGUE QUANTUM-FLUID SYSTEM

- We can model bubble growth using cold atoms
- We need a system of two components:

$$\psi_0 = \sqrt{n_0} e^{i\theta_0}, \qquad \psi_1 = \sqrt{n_1} e^{i\theta_1}$$

Static interaction potential:

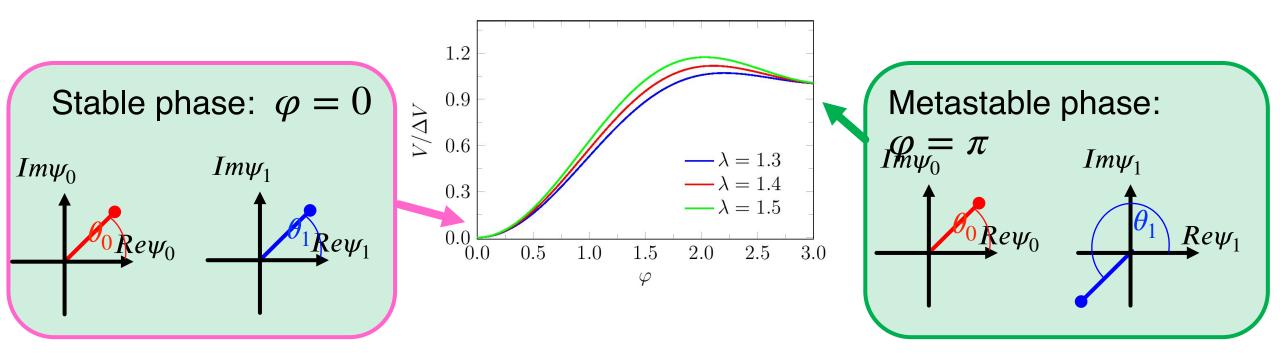
$$V = \frac{g}{2} \sum_{i} \left(\psi_i^* \psi_i \right)^2 - \mu \psi^* \psi - \mu \epsilon^2 \psi^* \sigma_x \psi + \frac{g}{4} \lambda^2 \epsilon^2 \left(\psi^* \sigma_y \psi \right)^2$$

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ANALOGUE QUANTUM-FLUID SYSTEM

- Scalar of interest: $\varphi = \theta_0 \theta_1$
- Interaction potential becomes:

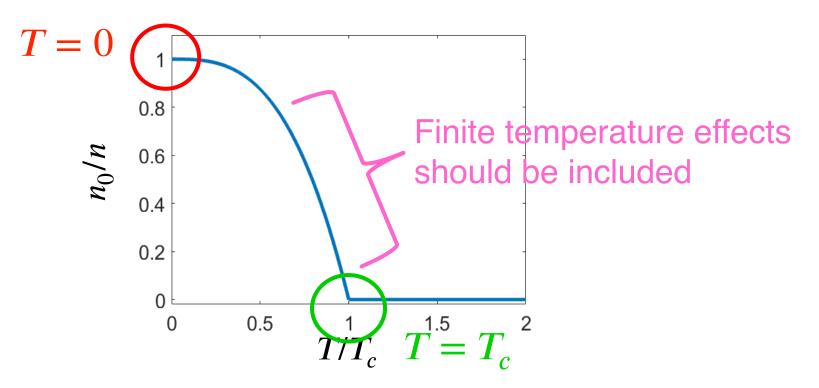
 $V \approx -2\epsilon^2 - 2\epsilon^2 \cos(\varphi) + \epsilon^2 \lambda^2 \sin^2(\varphi)$



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NUMERICAL MODEL

- We use a finite temperature approach
- Interested in all highly occupied, low energy modes
- Not just the ground state



NUMERICAL MODEL

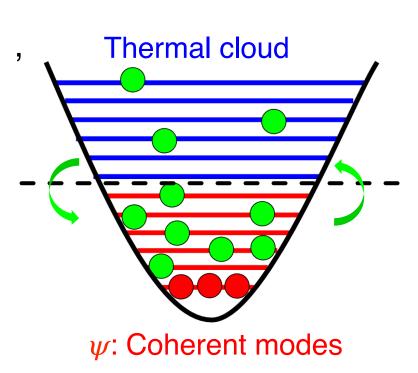
- Stochastic-Projected Gross-Pitaevskii Equation (SPGPE):
- Two coupled equations:

$$i\partial_t \boldsymbol{\psi}_j = \wp \left[(1 - i\gamma) \left\{ -\frac{1}{2} \nabla^2 \boldsymbol{\psi}_j + \frac{\partial V}{\partial \boldsymbol{\psi}_j^*} \right\} + \boldsymbol{\eta}_j \\ = 0, 1$$

- γ : Dissipation
- η : Gaussian noise with correlations:

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\left< \eta_i(x,t)\eta_j(x',t') \right> = 2\gamma T\delta(x-x')\delta(t-t')\delta_{ij}
```

D. Projector eliminates high wavelength modes.





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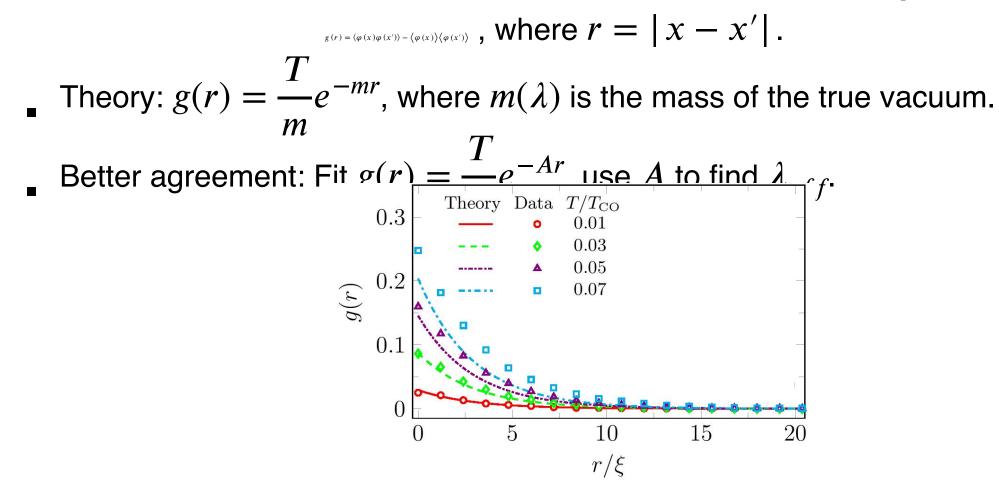
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EQUILIBRIUM

- System: 1D box with periodic boundaries
- Protocol: initiate in stable "true vacuum" state and let system equilibrate.
- Temperature range: T = 0.01 0.1
- $\lambda = 1.4$

CORRELATION FUNCTION

• Consider spatial correlation function of the **phase difference** ϕ

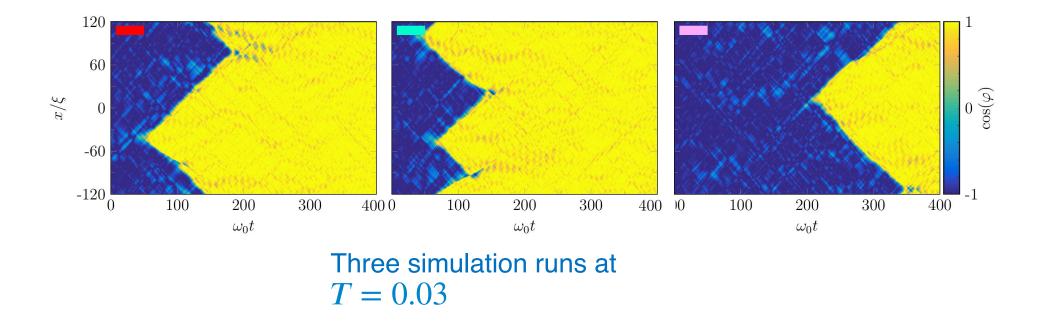


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OBSERVING BUBBLES

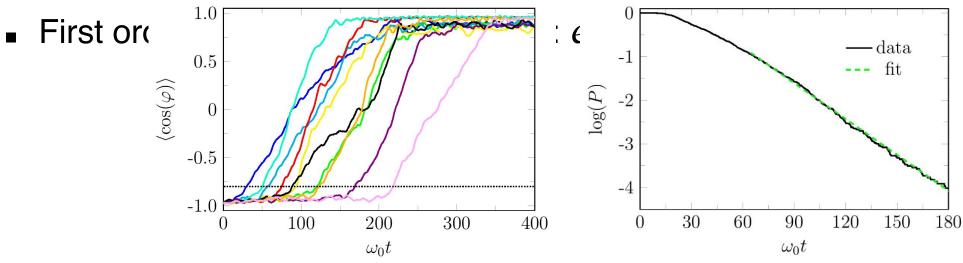
- System: 1D box with periodic boundaries
- Protocol: Initiate in metastable state and wait for bubbles to form
- Temperature range: T = 0.015 0.03



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OBSERVING BUBBLES

- We're interested in the probability, P, of remaining in the false vacuum
- Consider many trajectories of $\langle \cos(\varphi) \rangle$
- We say a bubble has formed when $\langle \cos(\varphi) \rangle$ reaches -0.8.
- P(t): The proportion of trajectories for which $\langle \cos(\varphi) \rangle < -0.8$



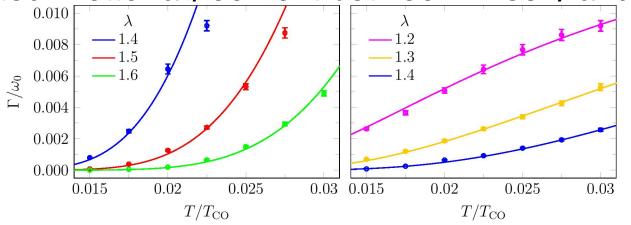
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OBSERVING BUBBLES

• Theory:

 $\Gamma = A \cdot B^{0.5} e^{-B},$ Where A is constant and $B = \frac{\alpha(\lambda)\epsilon}{T}$

- Left: A fitted, α taken from Gutierrez-Abed et al. (2020) [1]. Good agreement between Theory and Data.
- Right: A α fitted. Better agreement between Theory and Data.



[1] arxiv: 2006.06289

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FUTURE PLANS

- Replace time averaged potential with oscillating potential Zero temperature: Instability (Braden et al., JHEP, 2018) SPGPE: Can we damp this out?
- Upgrade from 1D to 2D

Full reference list: Arxiv 2006.09820

