

A Real-Time (Semiclassical) Picture of Vacuum Decay

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w/ Matt Johnson, Hiranya Peiris, Andrew Pontzen, and Silke Weinfurtner
1712.02356, 1806.06069, 1904.07873, and in progress

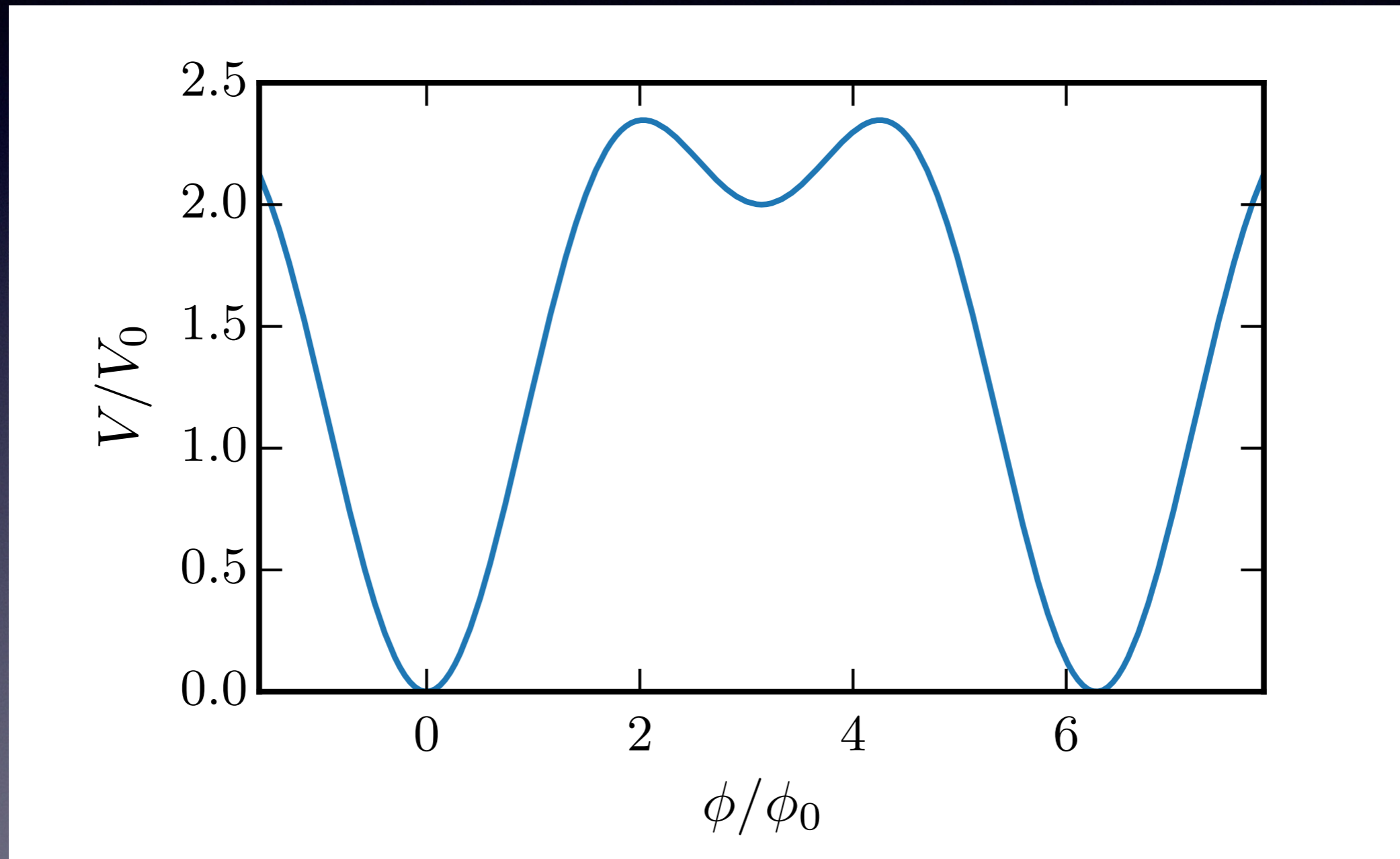
2018 Buchalter Cosmology 3rd Prize

Cosmology from Home

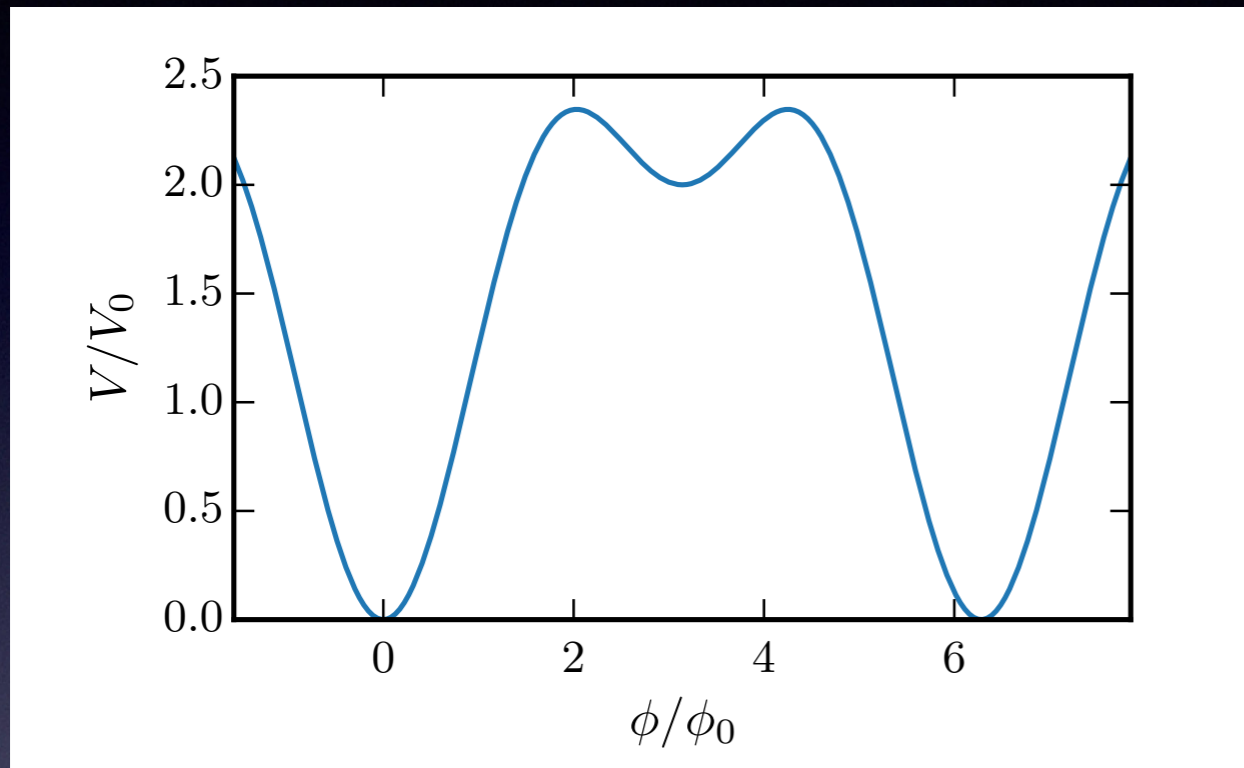
Outline

- Review of Vacuum Decay and 1st Order Phase Transitions
- Euclidean Description
- Real-Time Description of Decay
- Simulation in the Lab
- Conclusions / Novel Future Applications

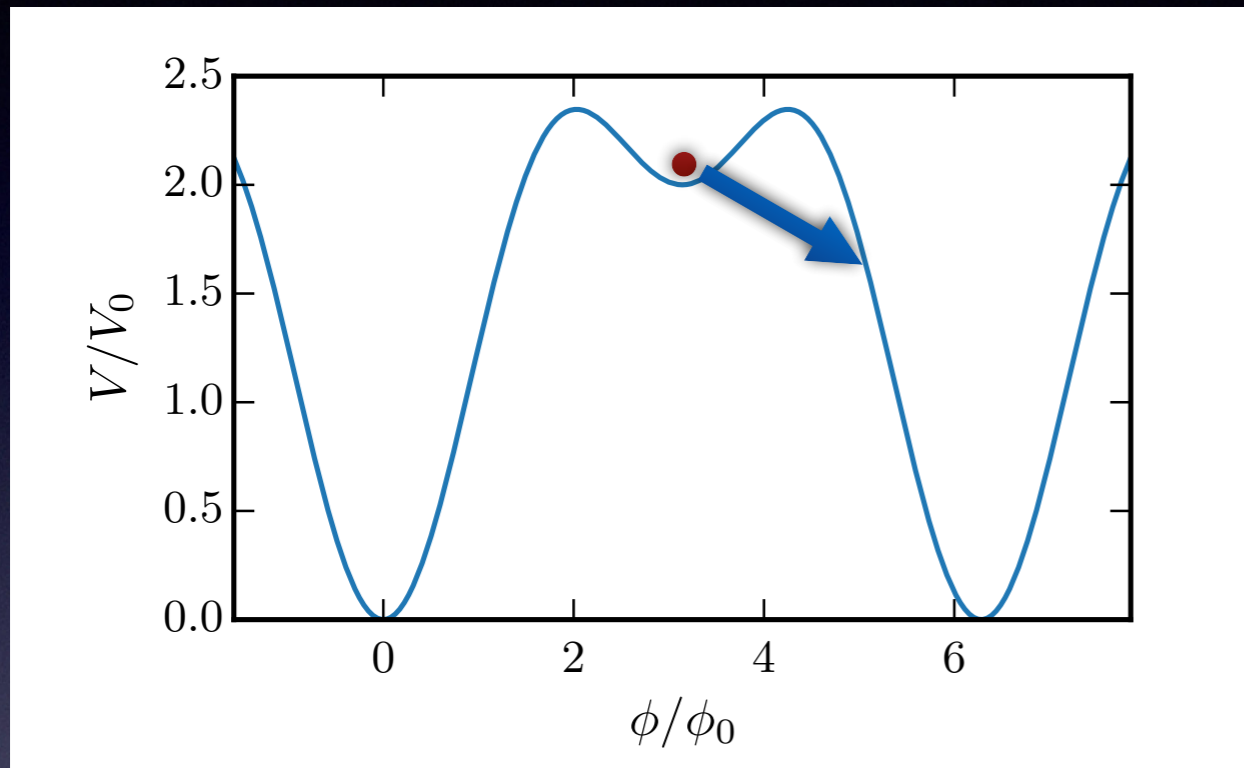
First Order Phase Transitions



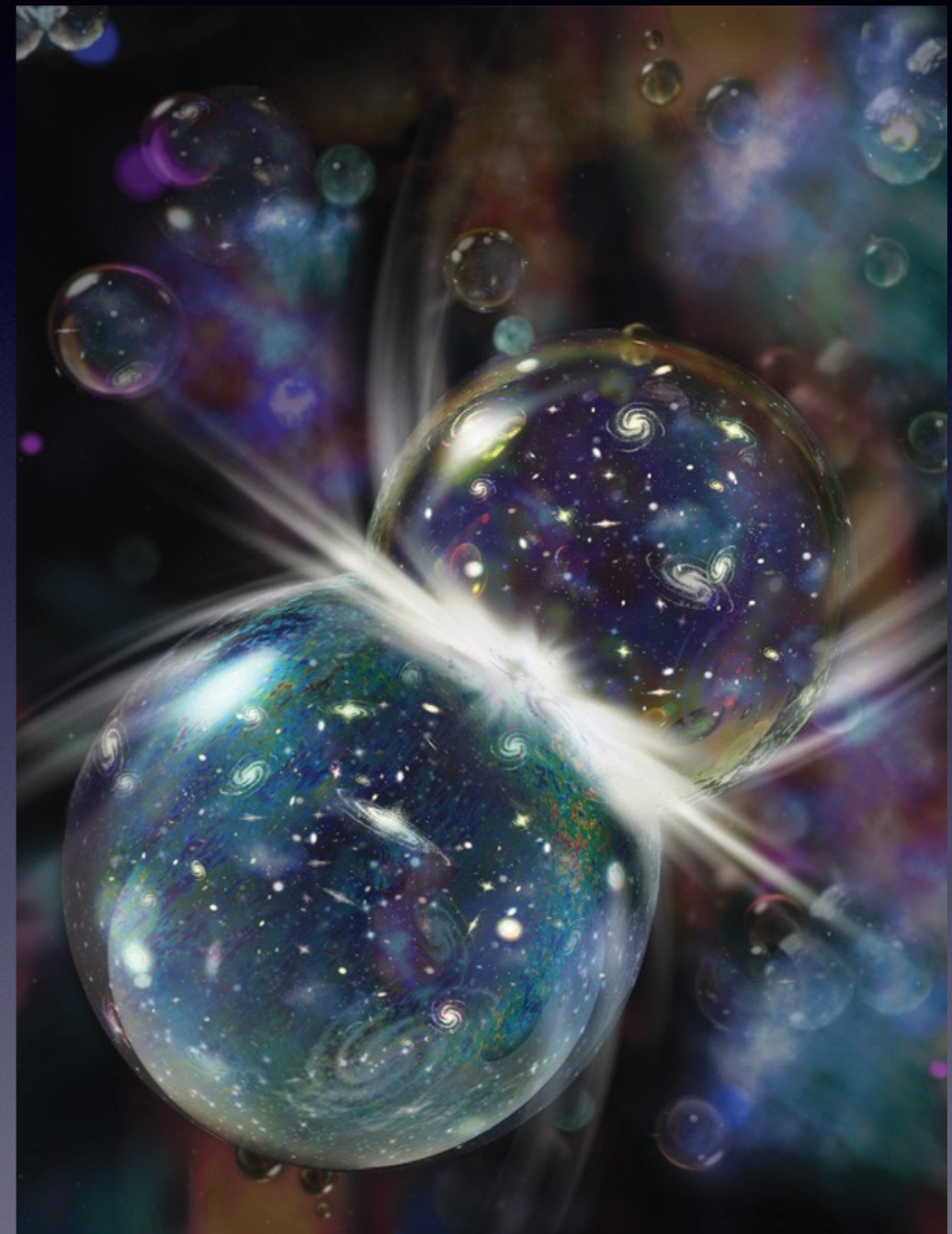
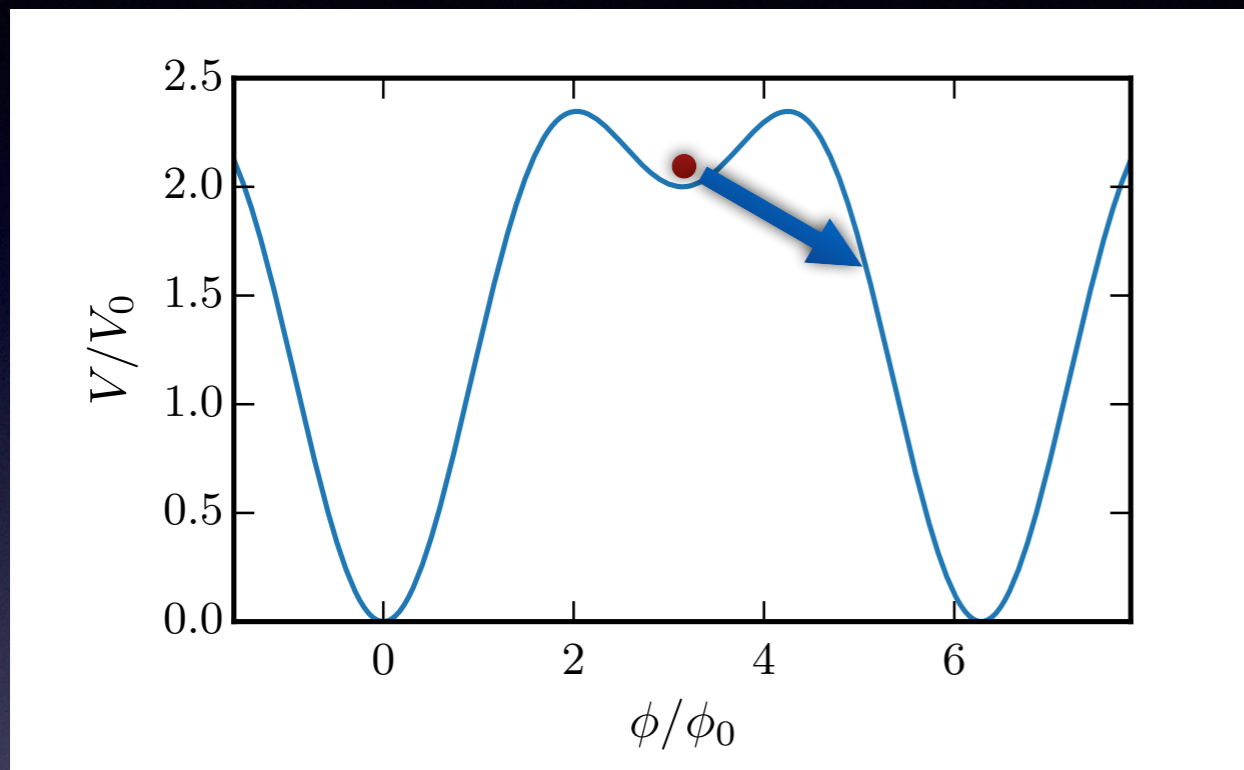
First Order Phase Transitions



First Order Phase Transitions



First Order Phase Transitions





H ATSMI

BEAUTY



0th Order Questions

- How fast does the vacuum decay?
- Do bubbles form?
- What do the bubbles look like?

Can be tackled with Euclidean bounce formalism

Bounce Formalism

[Coleman 1977 + many others]

$$P_{\text{undecayed}} = |\langle \Omega_{\text{FV}}(t) | \Omega_{\text{FV}}(t=0) \rangle|^2 \sim e^{-\Gamma t}$$

Work in Euclidean Time and Extremise Euclidean Action

$$S_{\text{E}} = A_{\text{d}+1} \int dr_{\text{E}} r_{\text{E}}^{\text{d}} \left(\frac{\phi'^2}{2} + V(\phi) \right)$$

Solve the following BVP

$$\frac{\partial^2 \phi_{\text{I}}}{\partial r_{\text{E}}^2} + \frac{d}{r_{\text{E}}} \frac{\partial \phi_{\text{I}}}{\partial r_{\text{E}}} - \frac{\partial V}{\partial \phi} = 0 \quad \frac{\partial \phi_{\text{I}}}{\partial r_{\text{E}}}(0) = 0 \quad \phi(\infty) = \phi_{\text{fv}}$$

$$r_{\text{E}}^2 = \tau^2 + \mathbf{r}^2 \quad \tau = it$$

Decay Rates

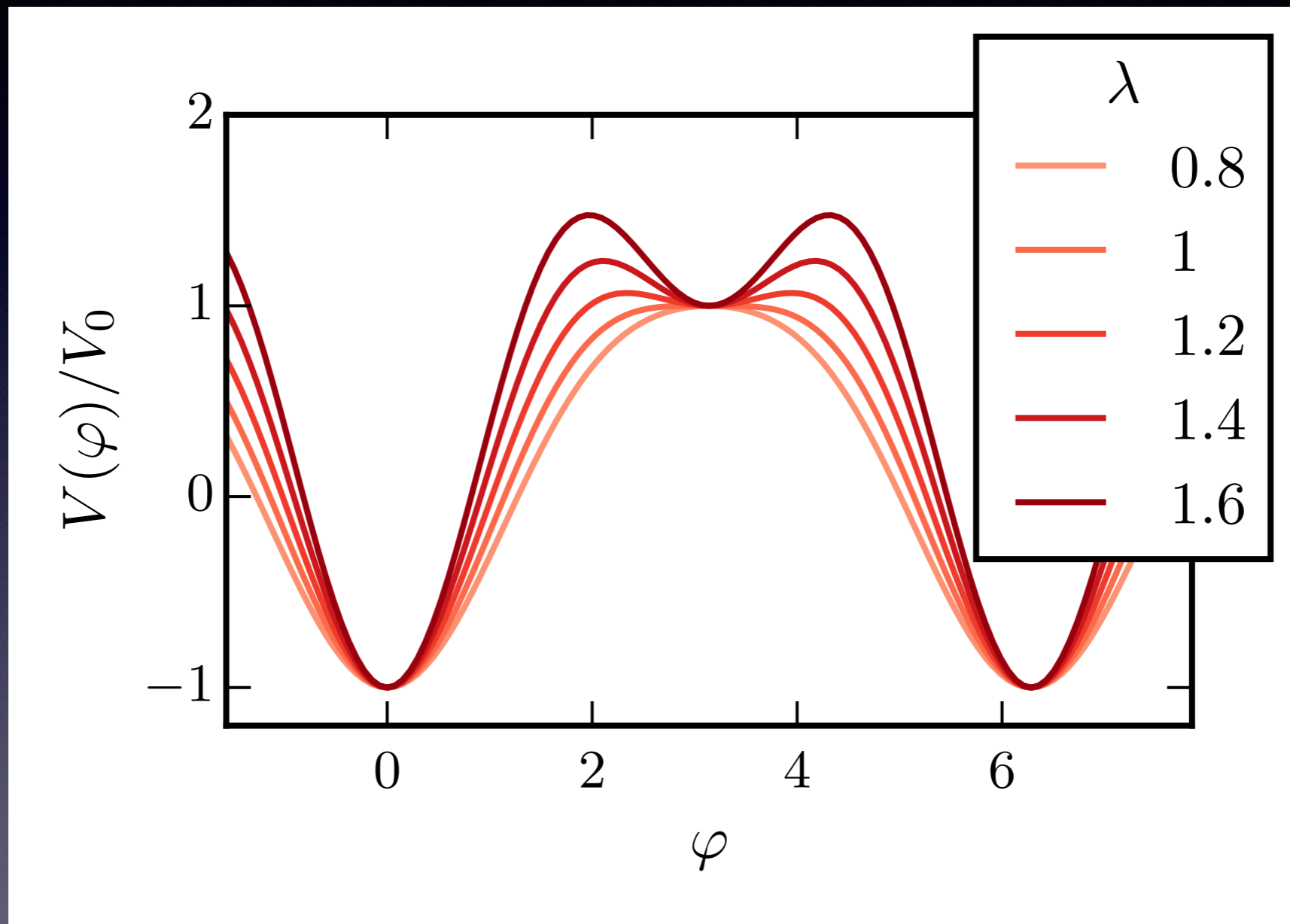
$$S_{\mathbf{E}} = A_{d+1} \int dr_{\mathbf{E}} r_{\mathbf{E}}^d \left(\frac{\phi'^2}{2} + V(\phi) \right)$$

$$S_{\mathbf{I}} = S_{\mathbf{E}}[\phi_{\mathbf{B}}] - S_{\mathbf{E}}[\phi_{\mathbf{fv}}]$$

- Single negative eigenmode

$$\frac{\Gamma}{V} = \left(\frac{S_{\mathbf{I}}}{2\pi} \right)^{D/2} \sqrt{\frac{\det \delta^2 S_{\mathbf{E}}[\phi_{\mathbf{fv}}]}{\det' \delta^2 S_{\mathbf{E}}[\phi_{\mathbf{B}}]}} e^{-S_{\mathbf{I}}} (1 + \mathcal{O}(\hbar))$$

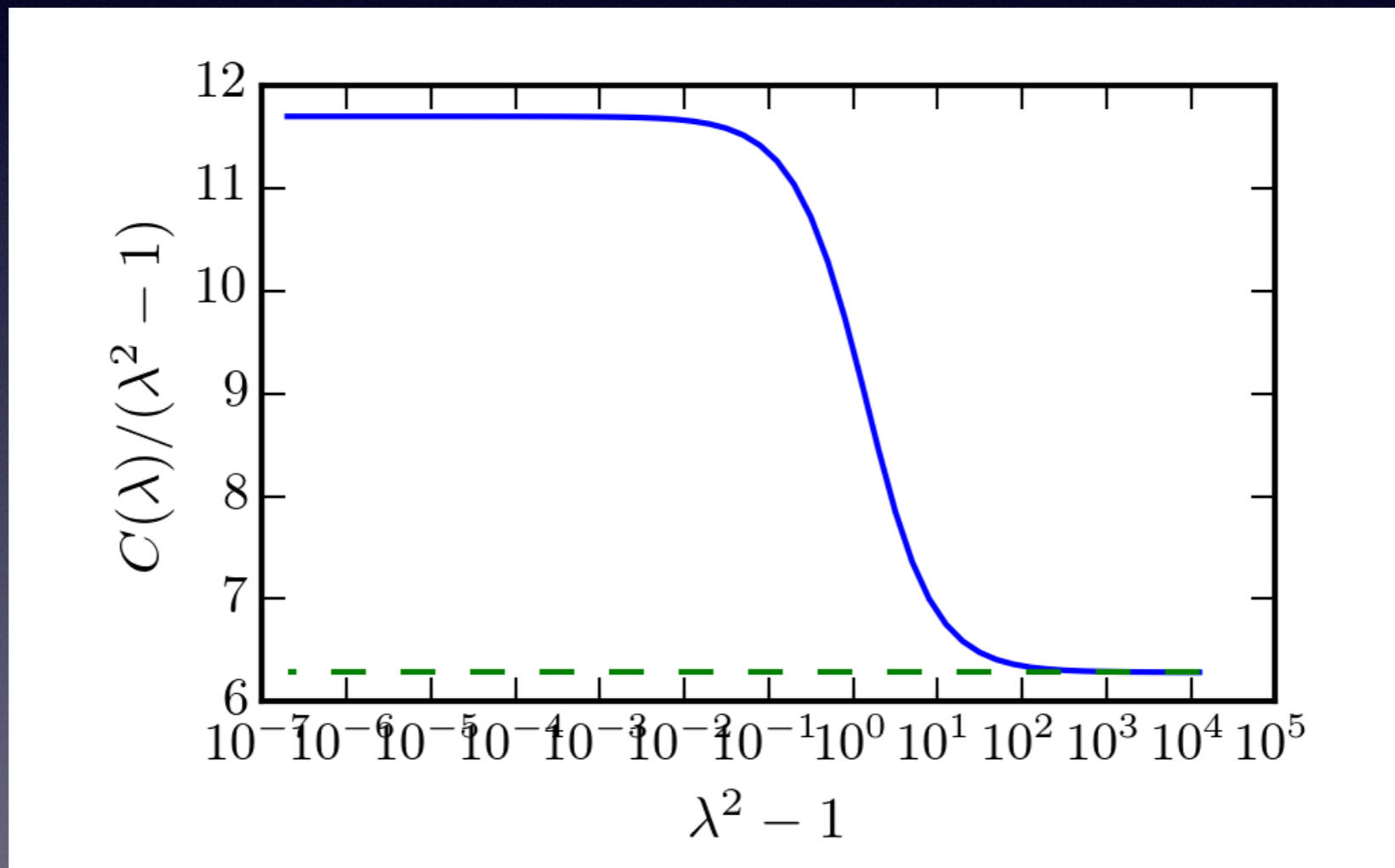
Model



$$V(\phi) = V_0 \left(-\cos \left(\frac{\phi}{\phi_0} \right) + \frac{\lambda^2}{2} \sin^2 \left(\frac{\phi}{\phi_0} \right) + 1 \right)$$

Nucleation Rates

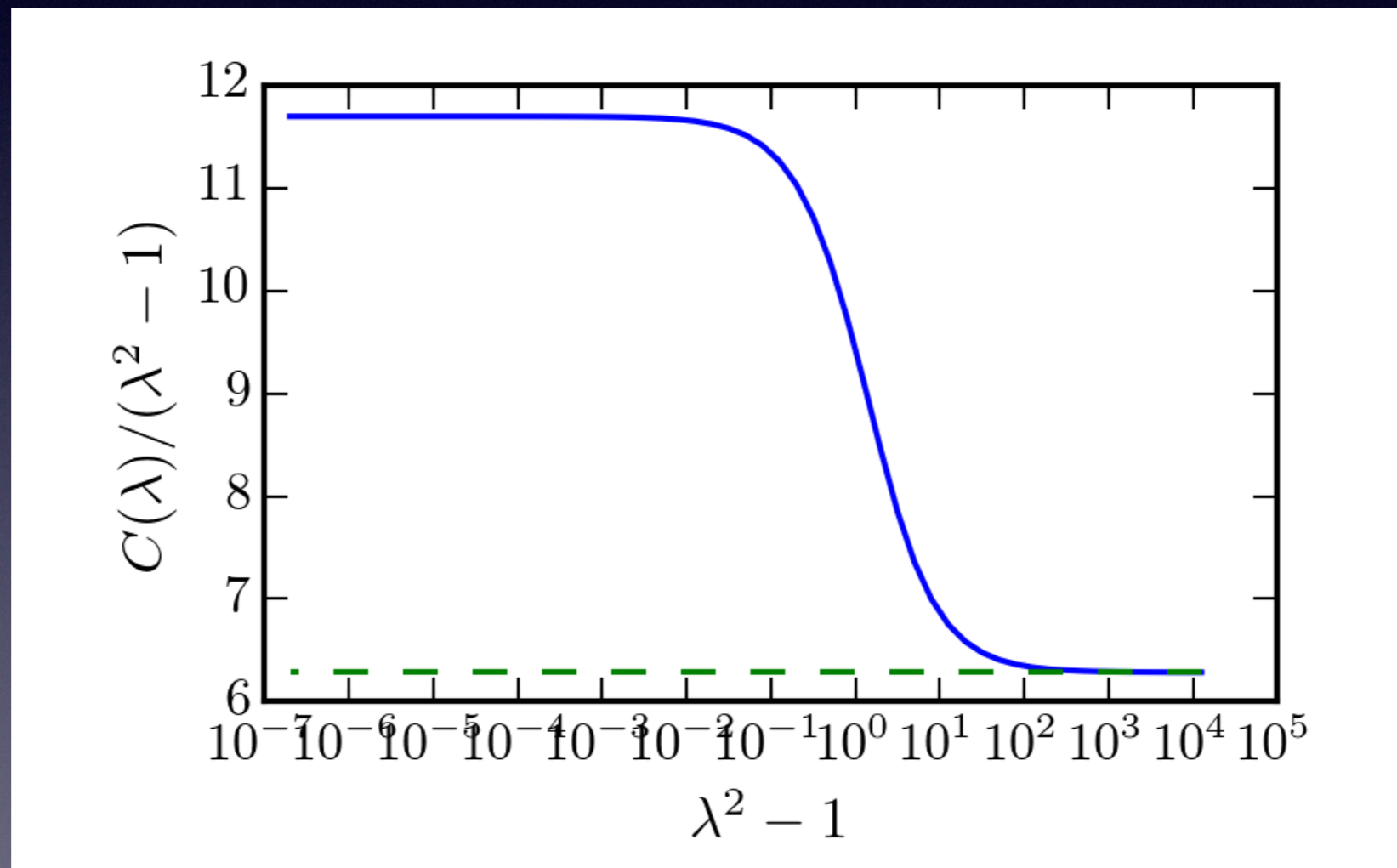
$$\frac{\Gamma}{V} \approx g(\lambda) [m_{\text{eff}}^2]^{\frac{D}{2}} \left(\frac{S_I}{2\pi} \right)^{\frac{D}{2}} e^{-S_I}$$



$$V(\phi) = V_0 \left(-\cos \left(\frac{\phi}{\phi_0} \right) + \frac{\lambda^2}{2} \sin^2 \left(\frac{\phi}{\phi_0} \right) + 1 \right)$$

Nucleation Rates

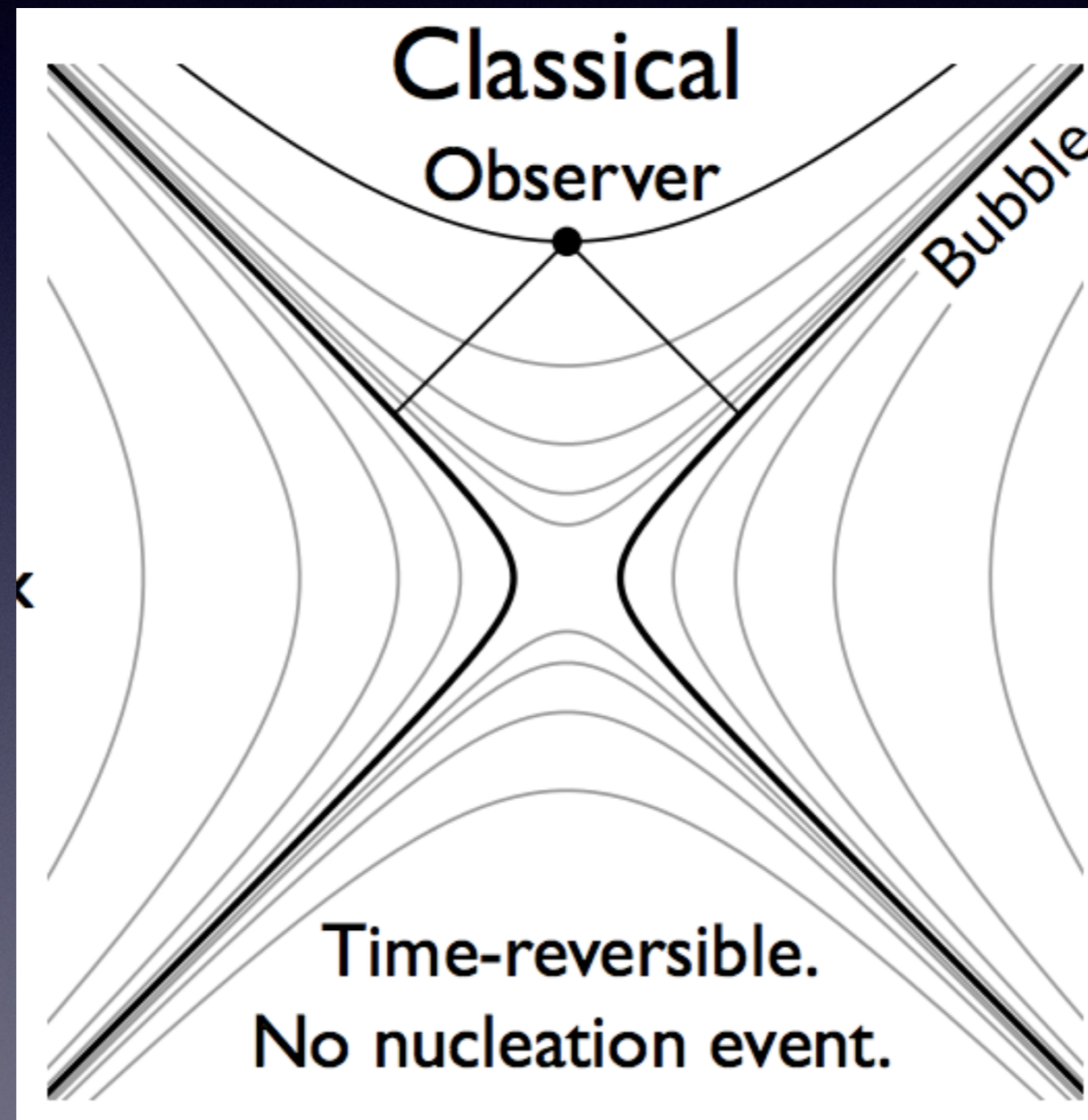
$$S_I = 2\pi\phi_0^2 C(\lambda)$$



$$V(\phi) = V_0 \left(-\cos \left(\frac{\phi}{\phi_0} \right) + \frac{\lambda^2}{2} \sin^2 \left(\frac{\phi}{\phi_0} \right) + 1 \right)$$

Real-Time Interpretation

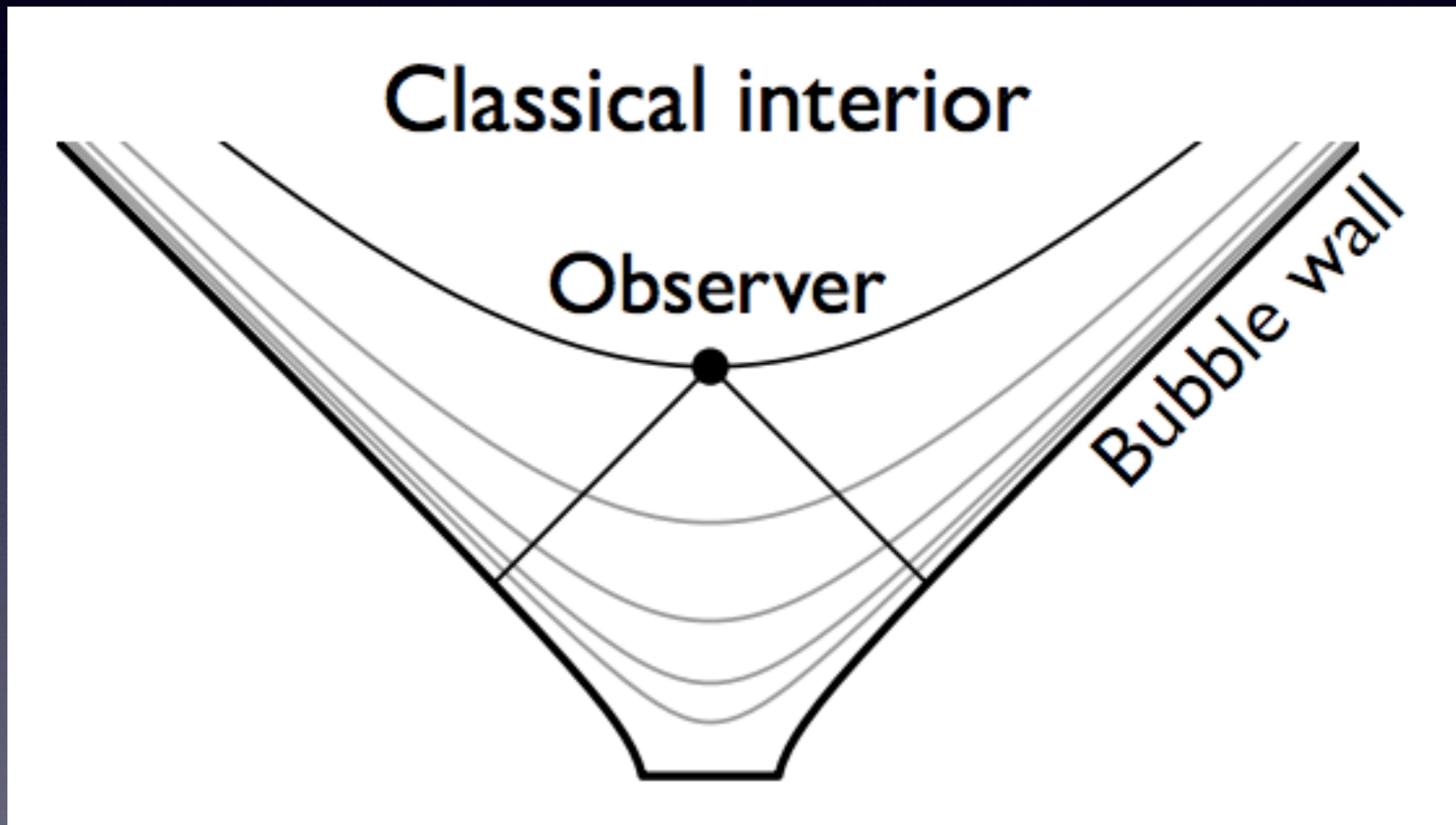
$$\phi(x, t) = \phi_I(\sqrt{x^2 - t^2})$$



[Figure courtesy of Andrew Pontzen]

Ad-Hoc Nucleation

$$\phi(\mathbf{x}, t = 0) = \phi_I(|\mathbf{x}|)$$



No real-time classical description

[Figure courtesy of Andrew Pontzen]

Some Questions

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- Time-dependent description of nucleation

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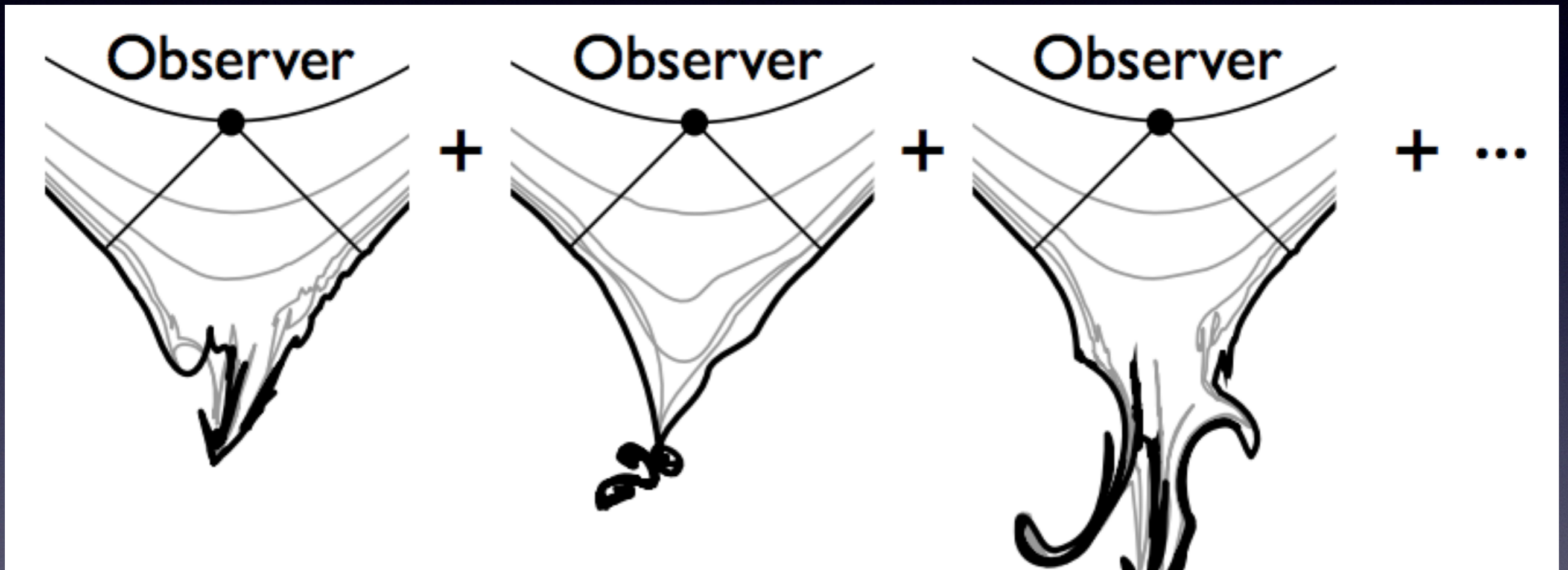
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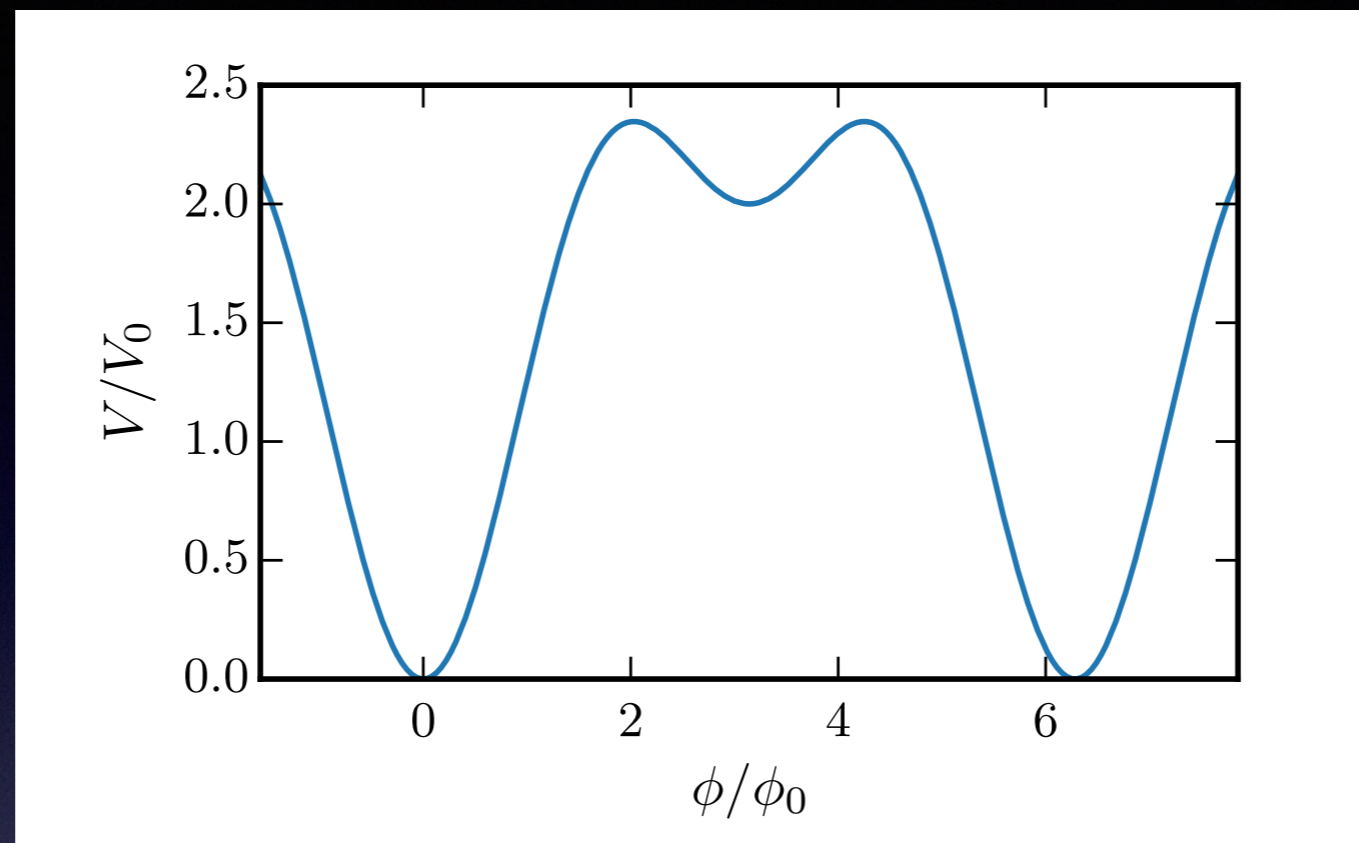
- Time-dependent description of nucleation
 - Bubble precursor? Initial nucleation
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**QFT exponentially complex.
Need Approximations!**

Full Evolution?

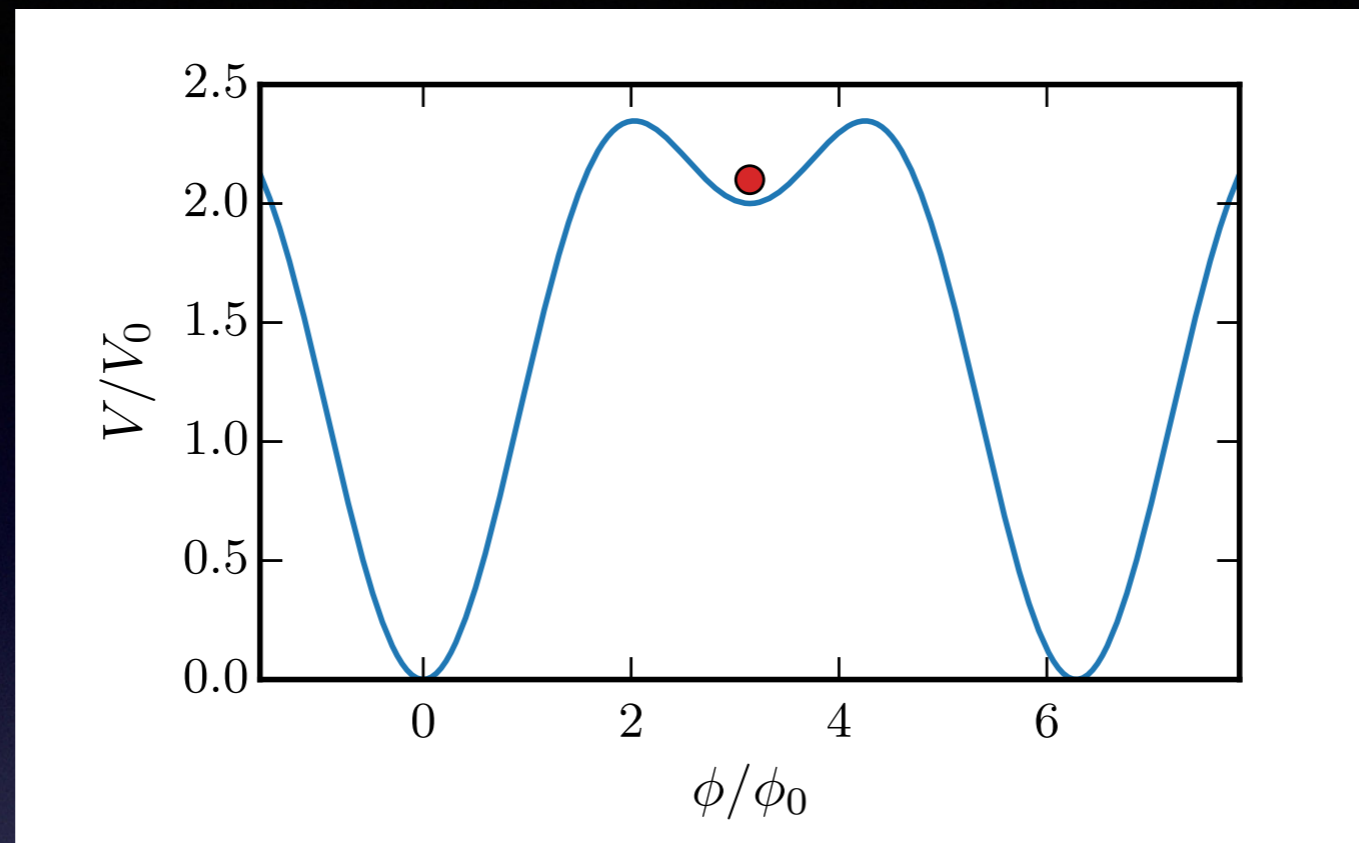


[Figure courtesy of Andrew Pontzen]



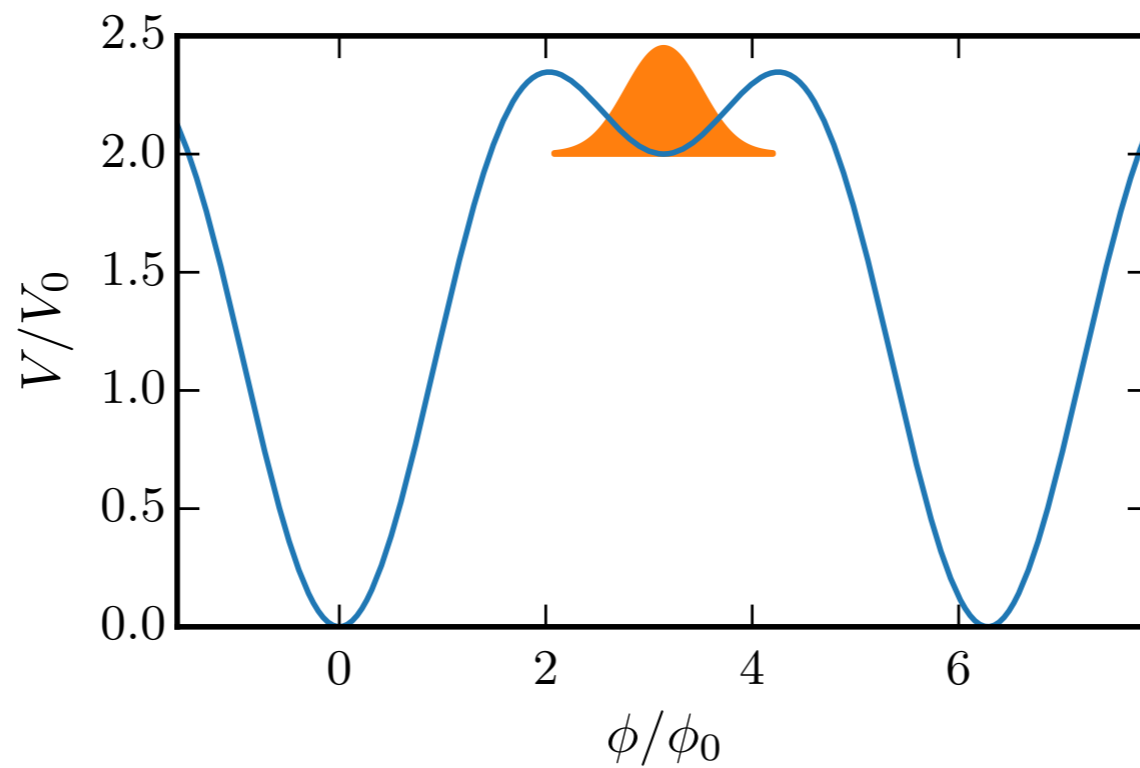
$$\phi =$$

$$\Pi =$$



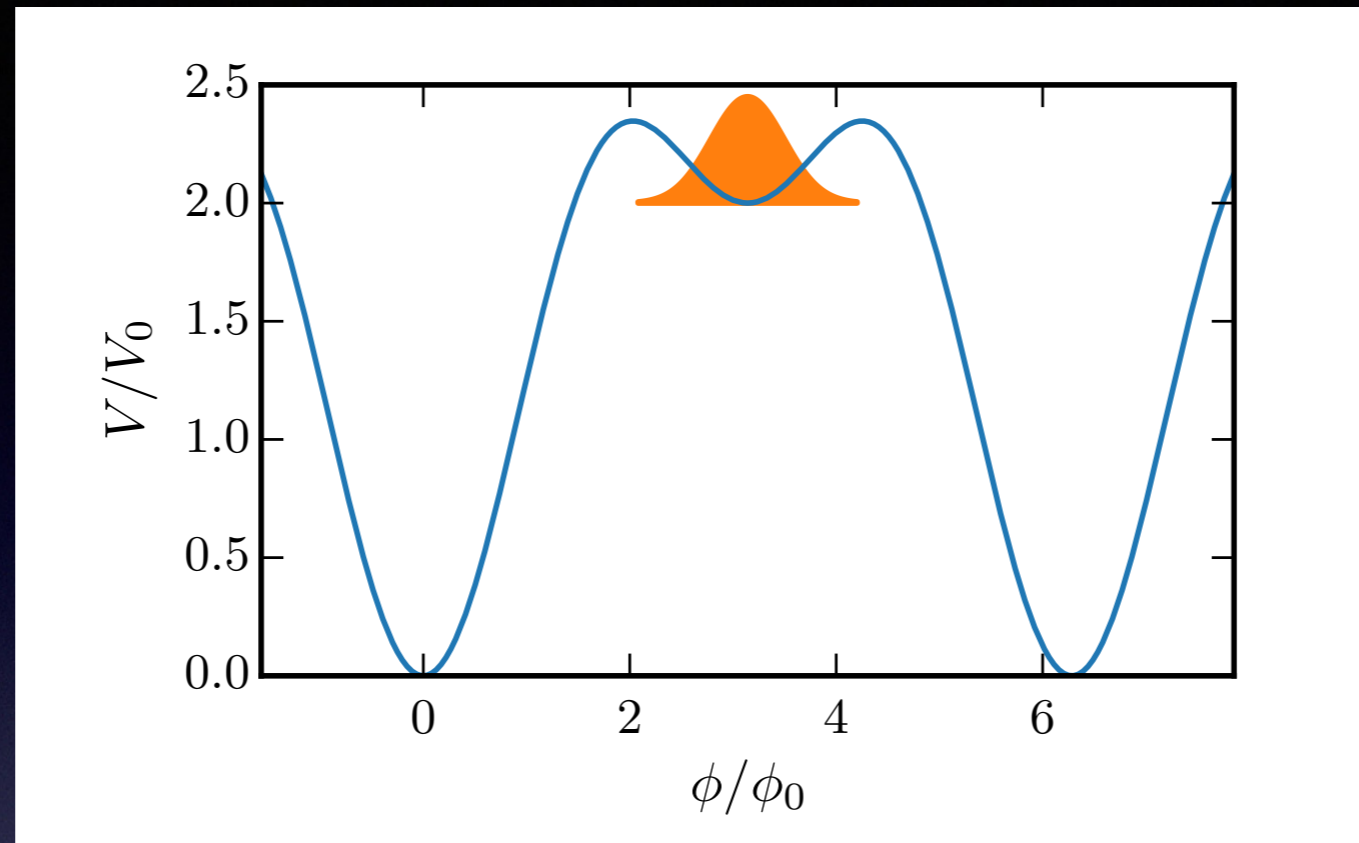
$$\phi = \phi_{fv}$$

$$\Pi = 0$$



$$\phi = \phi_{fv} + \delta\hat{\phi}(\mathbf{x}, t)$$

$$\Pi = 0 + \delta\hat{\Pi}(\mathbf{x}, t)$$

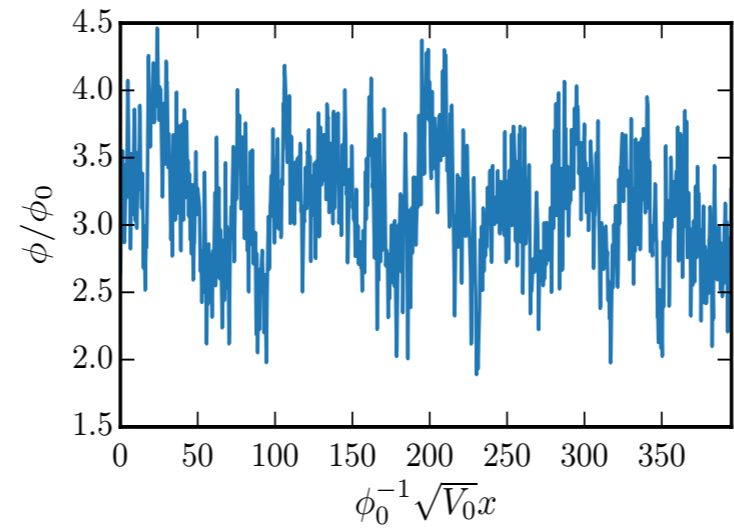


$$\phi = \phi_{\text{fv}} + \delta\hat{\phi}(\mathbf{x}, t)$$

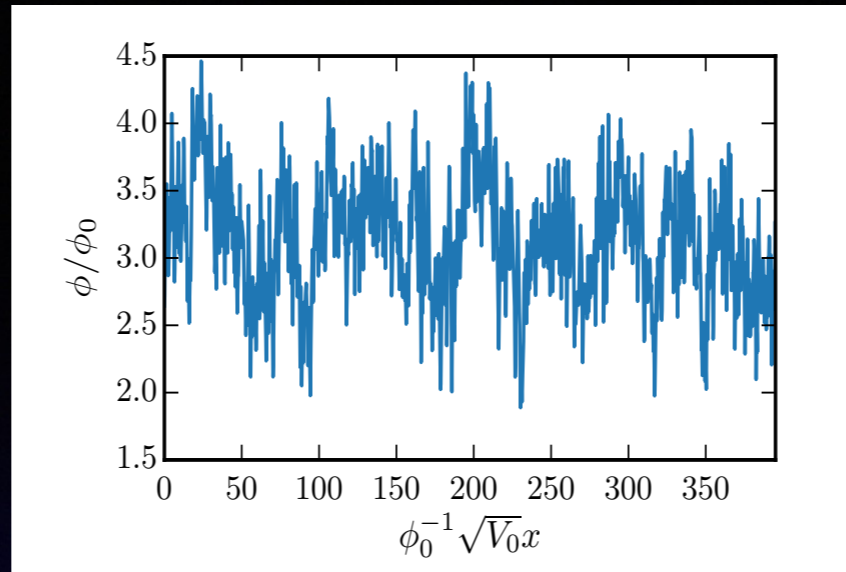
$$\Pi = 0 + \delta\hat{\Pi}(\mathbf{x}, t)$$

$$\langle \delta\tilde{\phi}_k \delta\tilde{\phi}_p^* \rangle = \frac{1}{2\omega_k} \delta(k - p) \quad \langle \delta\tilde{\Pi}_k \delta\tilde{\Pi}_p^* \rangle = \frac{\omega_k}{2} \delta(k - p)$$

Quantum Commutators



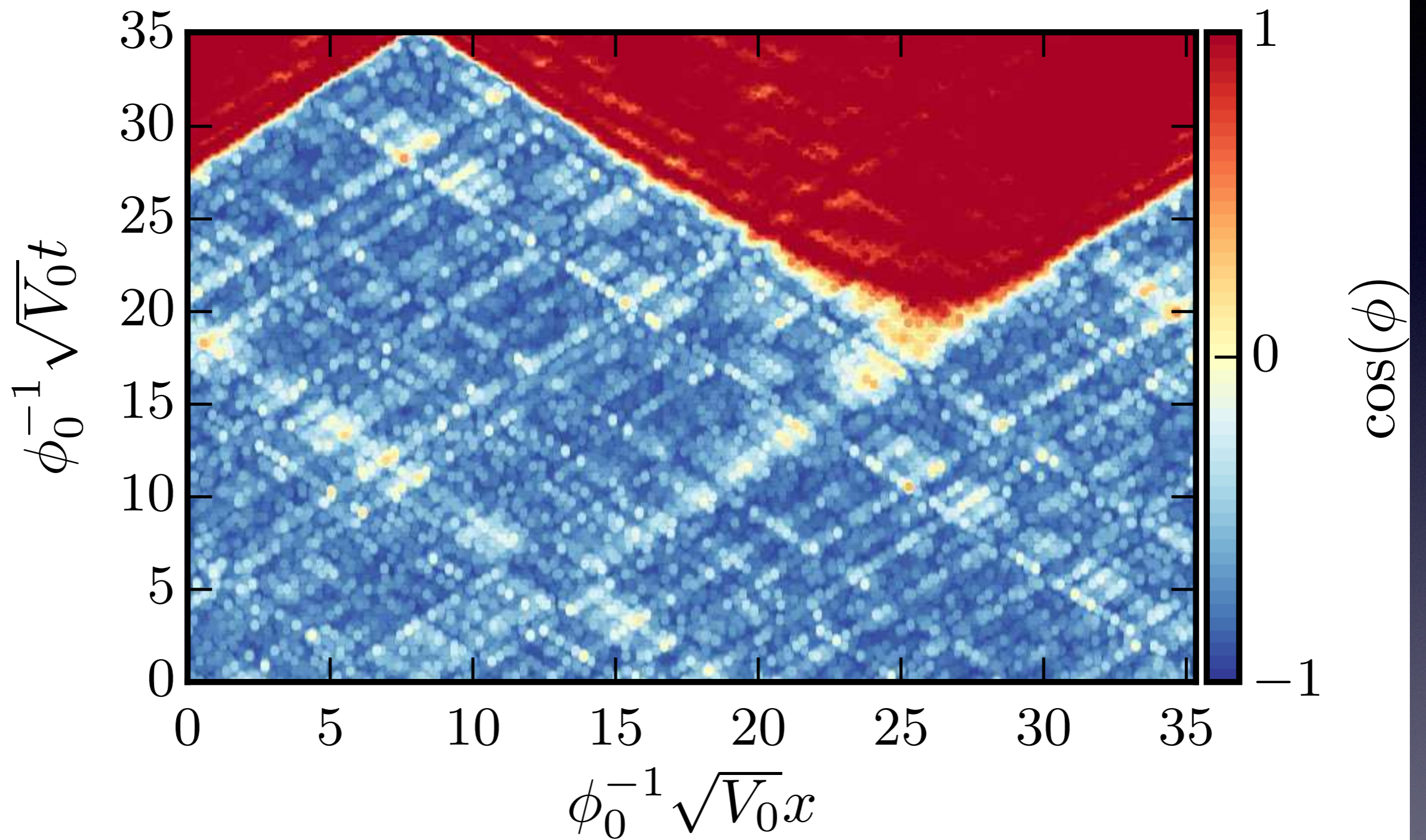
Quantum Commutators



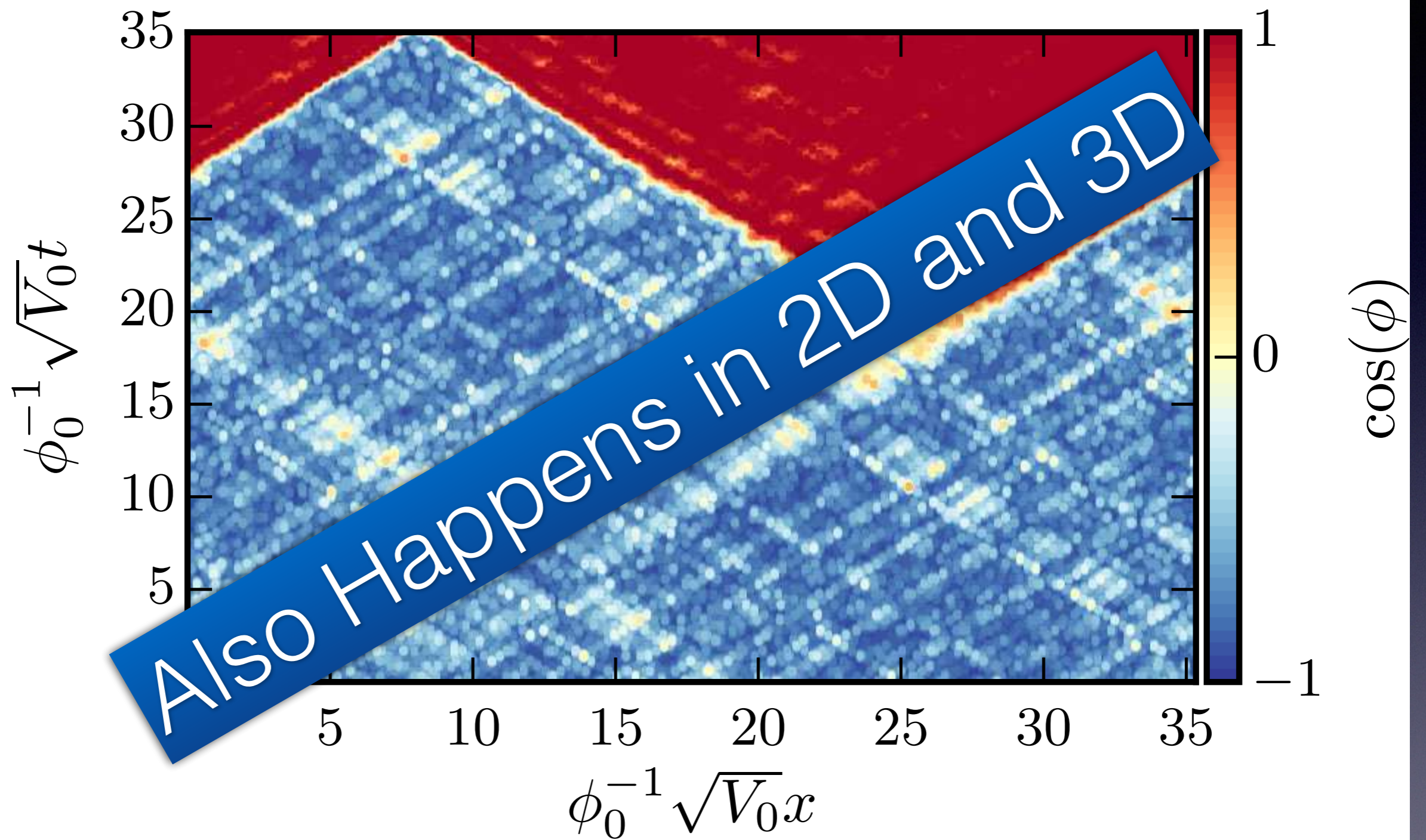


$$\ddot{\phi} - \nabla^2 \phi + V'(\phi) = 0$$



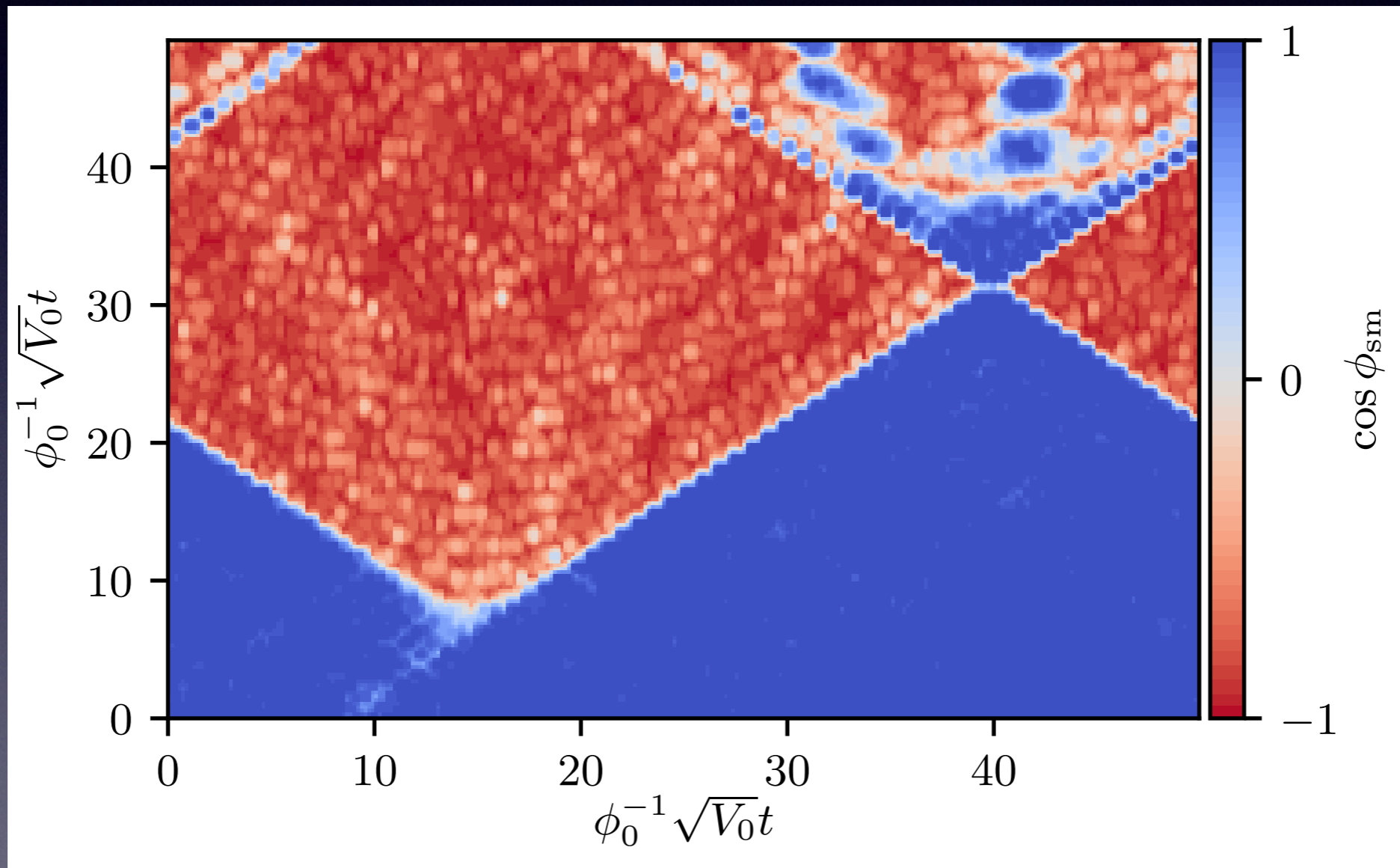


Classically-Allowed Vacuum Decay

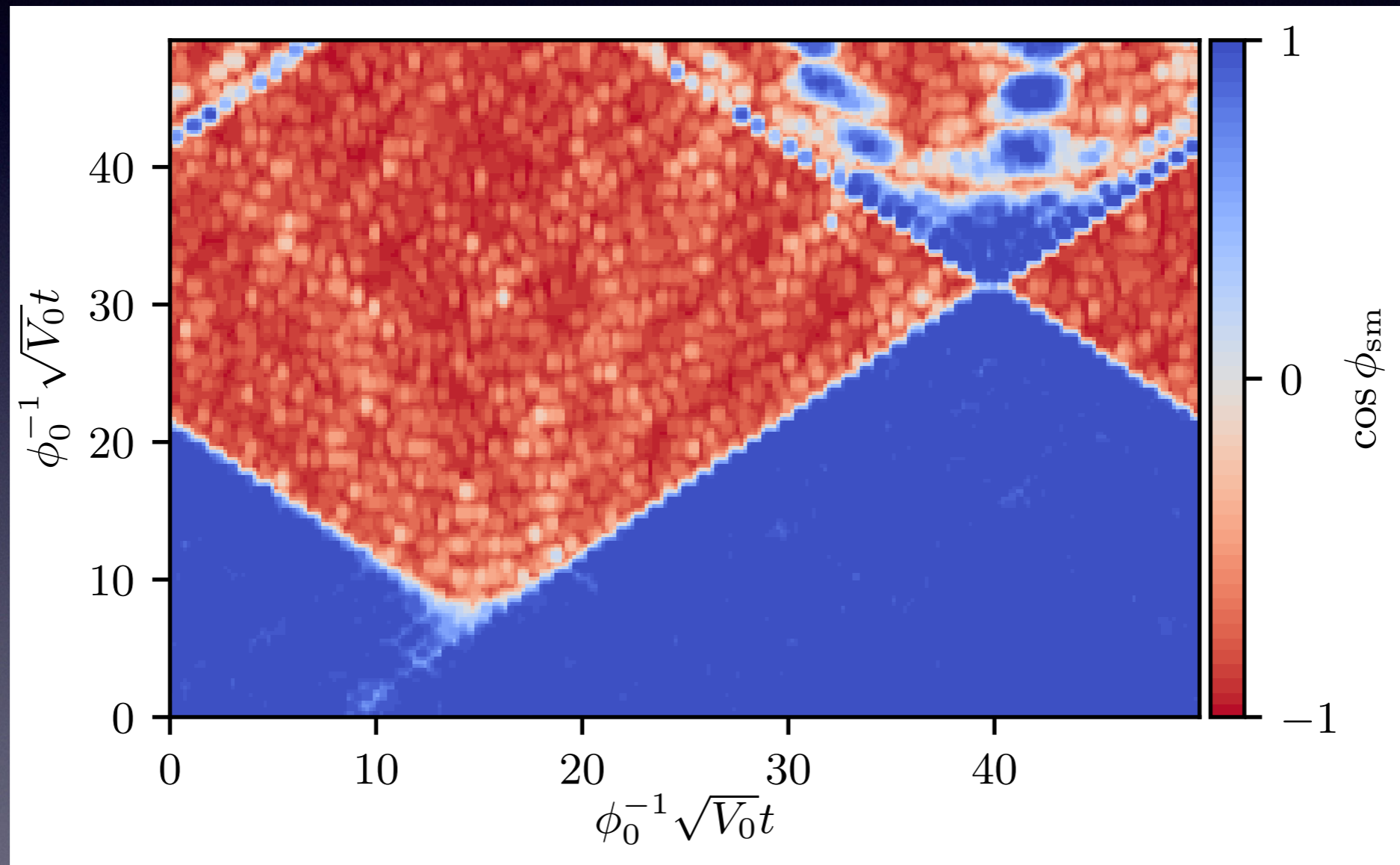


Classically-Allowed Vacuum Decay

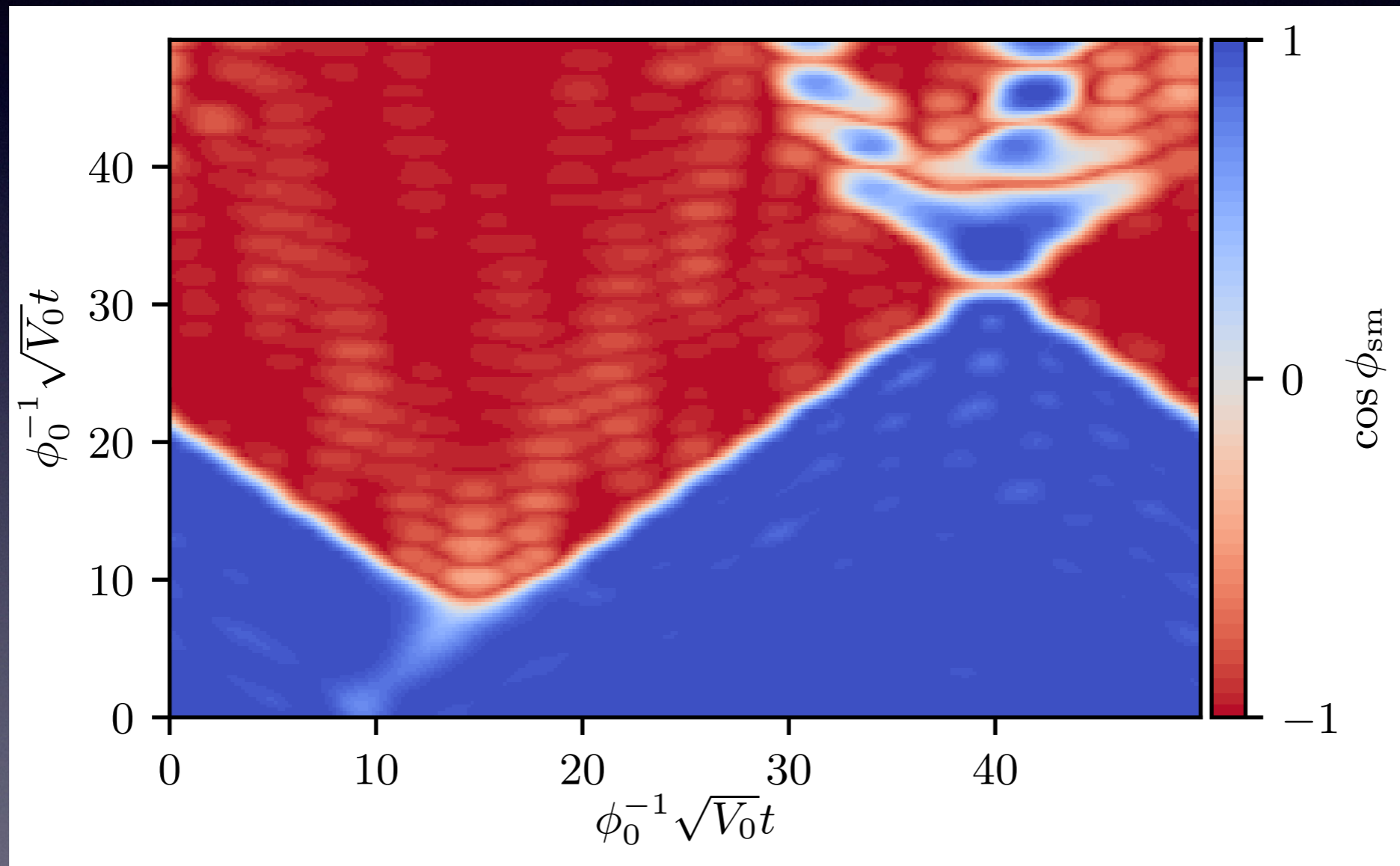
Energy of Bubble



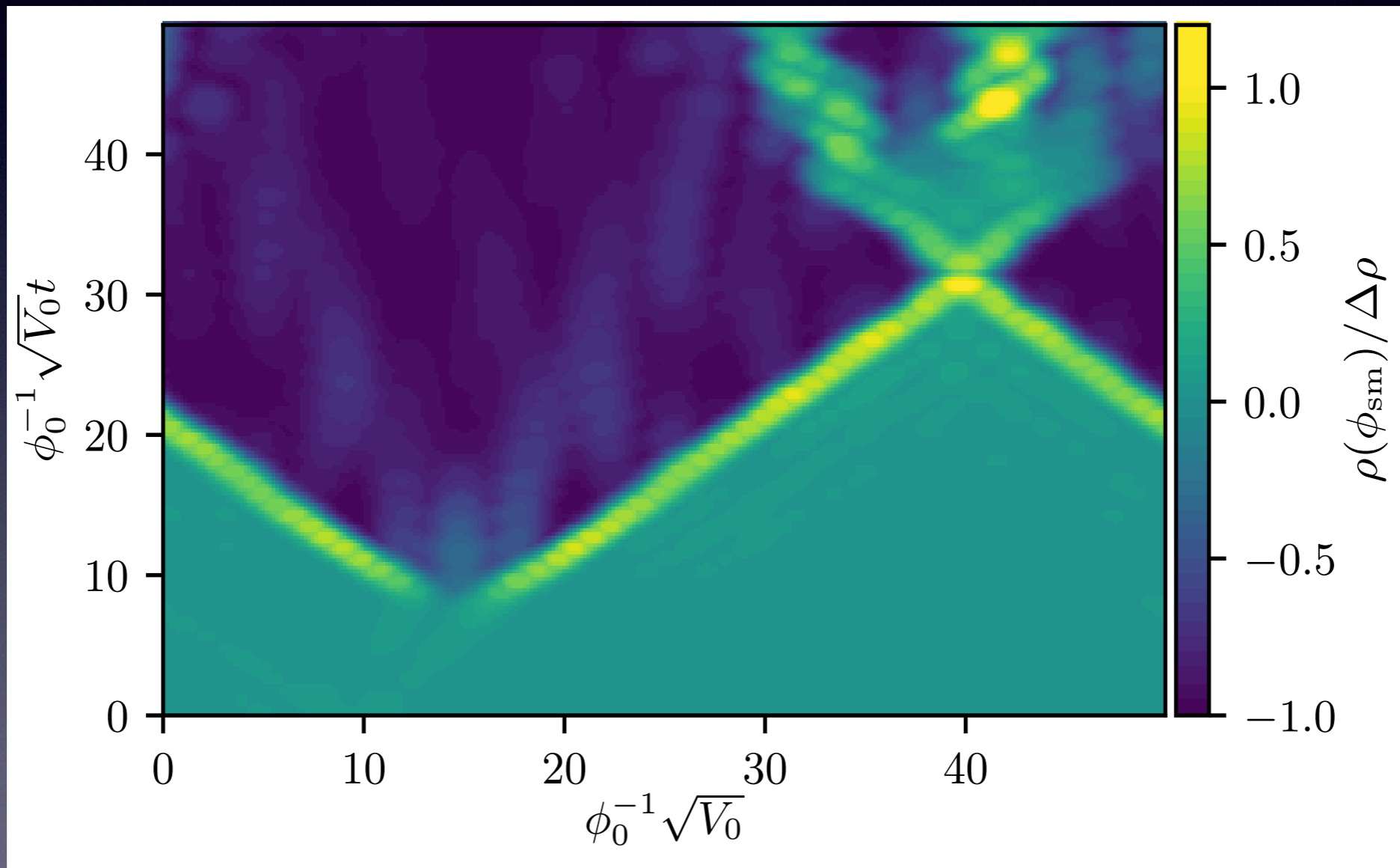
Restrict to long wavelength field modes instead



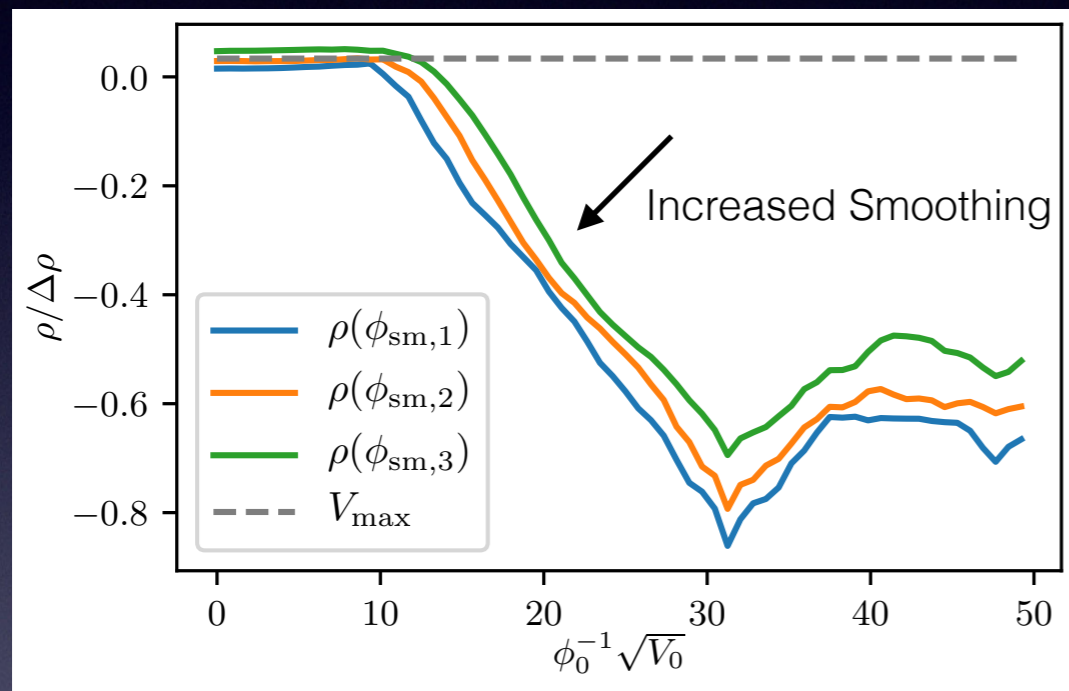
Restrict to long wavelength field modes instead



Energy of smoothed fields

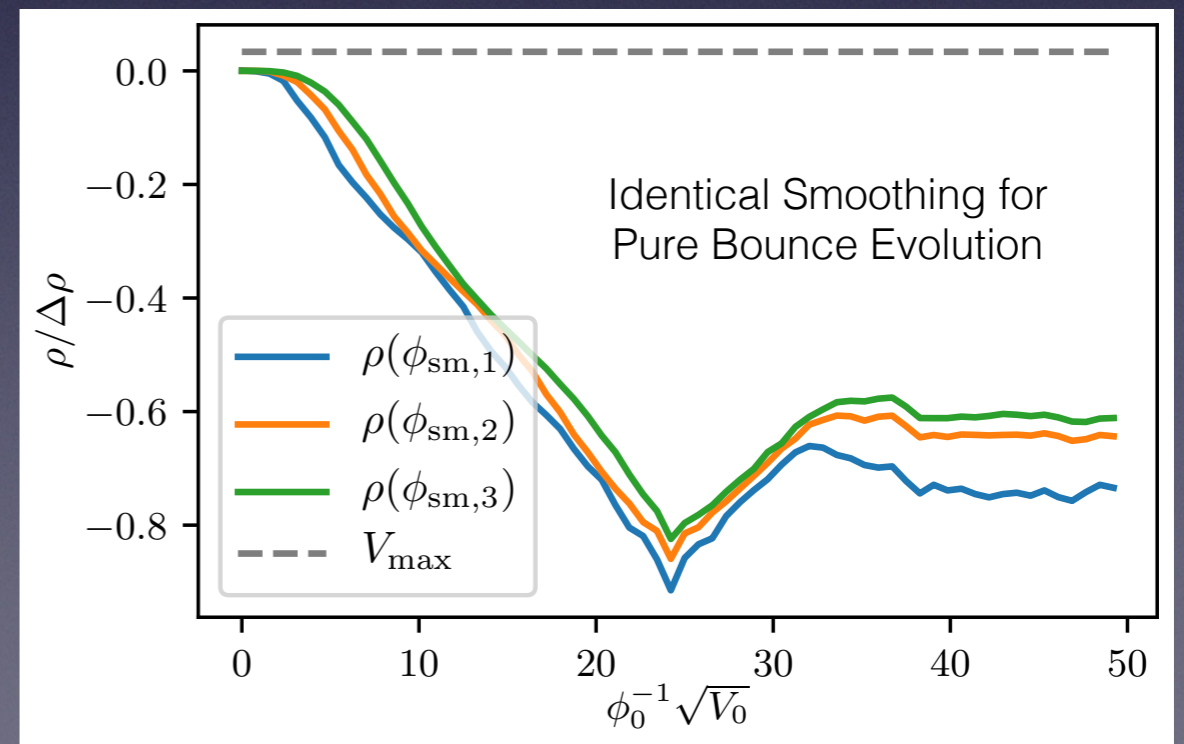


Energy of long wave fields matches instanton



Energy in long-wave modes as smoothing is varied

Compares well to bounce bubble initialized at $t=0$



Decay Rates?

Prediction

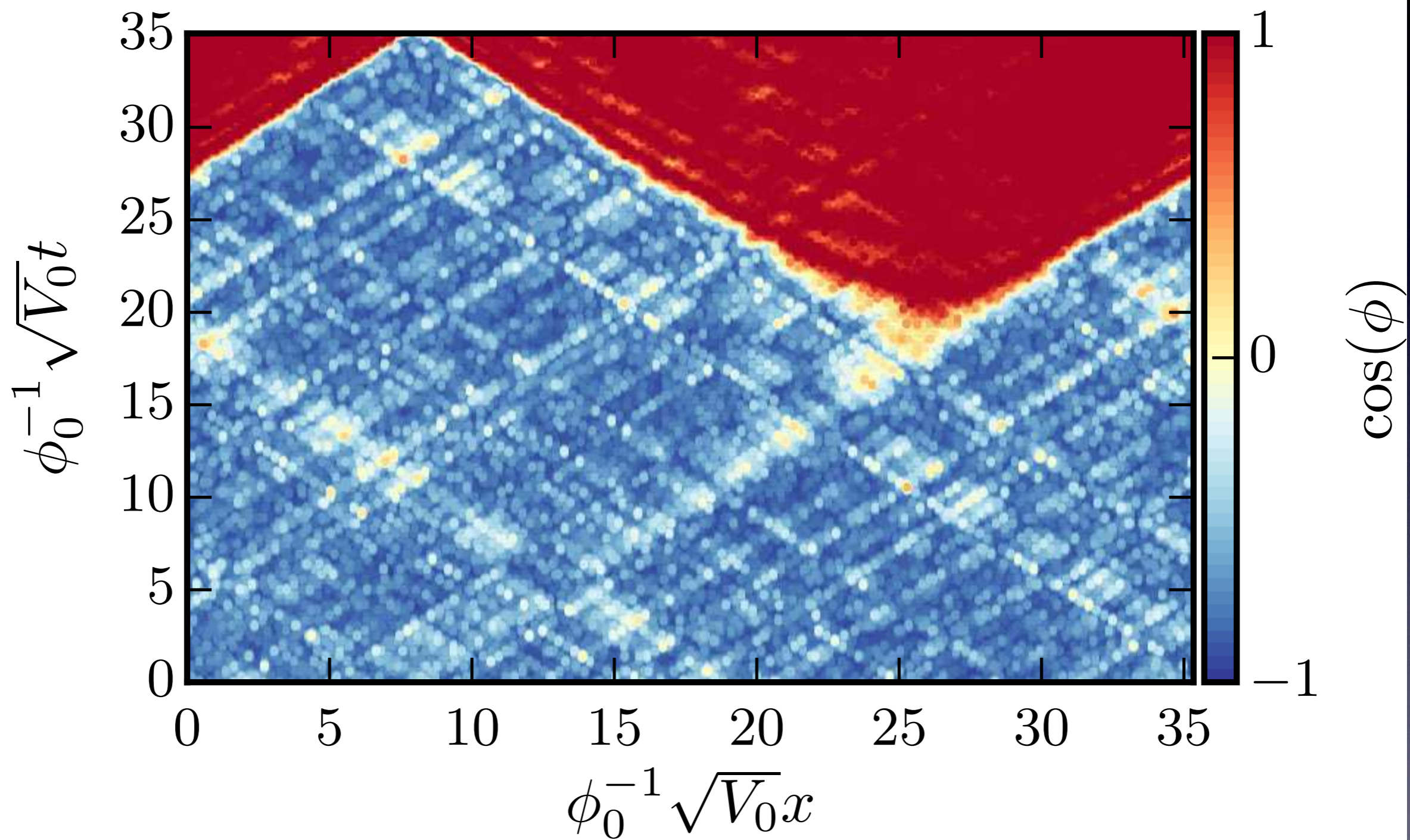
$$\frac{\Gamma_I^{(1+1)}}{L} \approx g(\lambda, V_0, \phi_0) m_{\text{eff}}^2 \phi_0^2 C(\lambda) e^{-2\pi \phi_0^2 C(\lambda)}$$

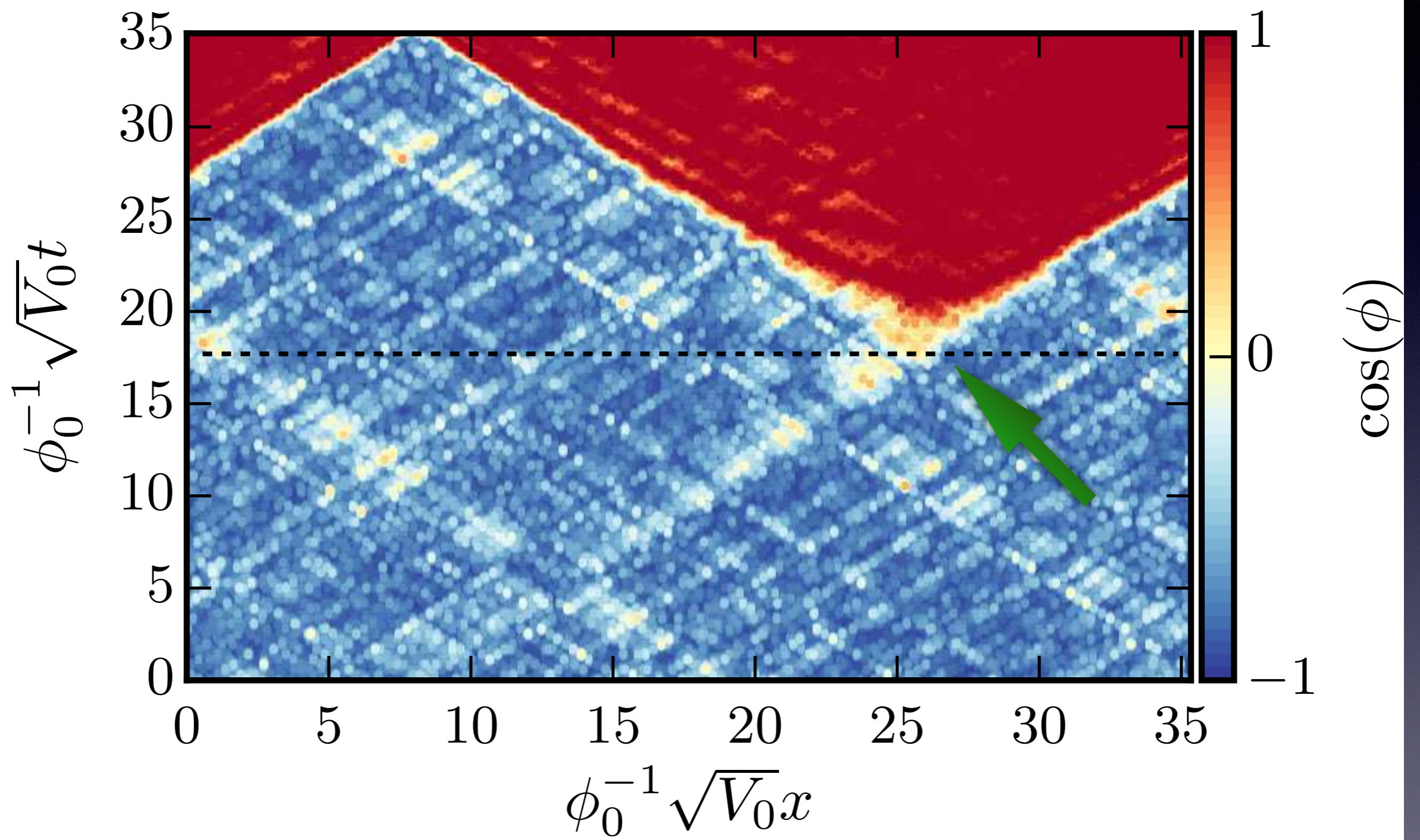
$\mathcal{O}(1)$

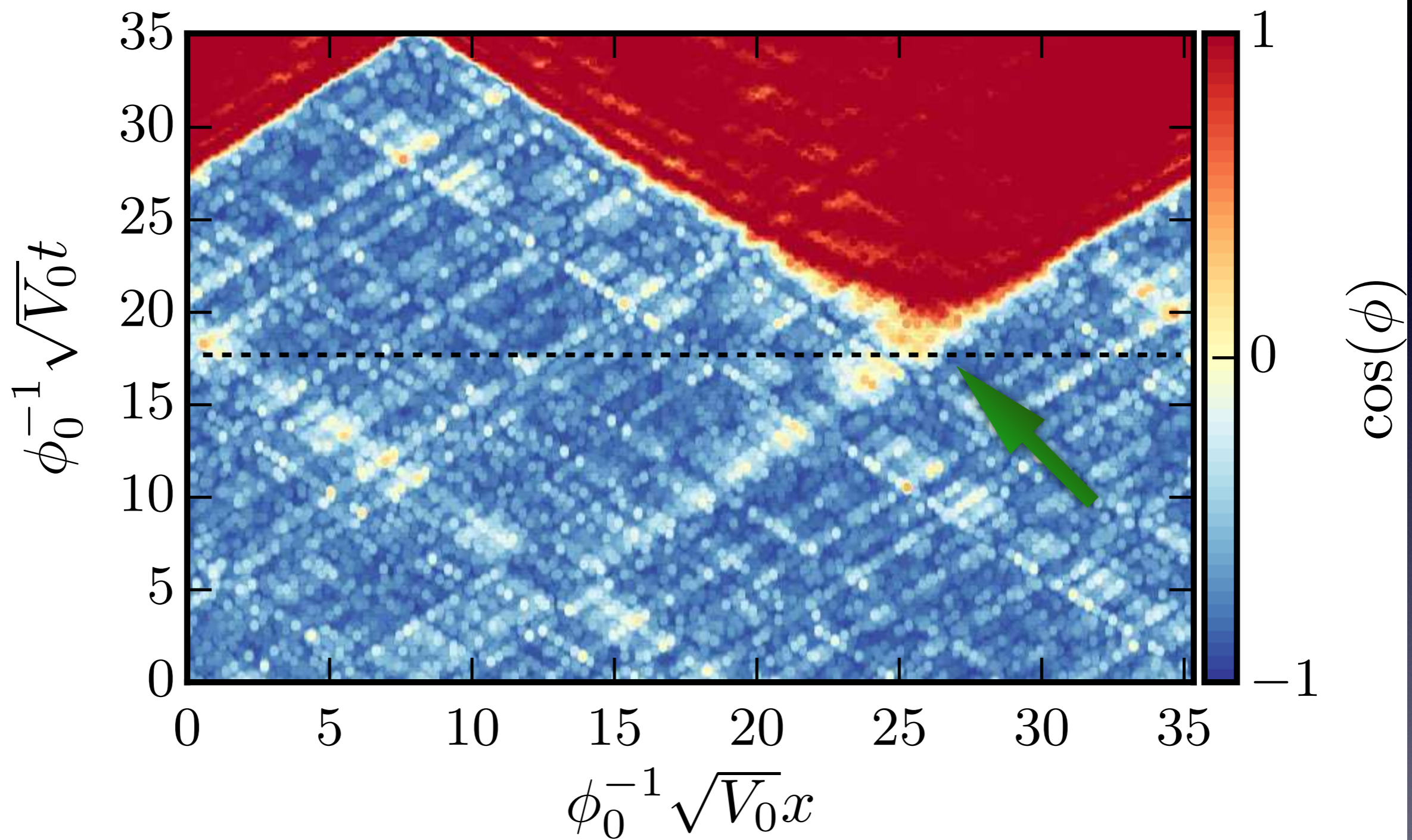
$\sim V''(\phi_{\text{fv}})$

Instanton

$$V(\phi) = V_0 \left(-\cos \left(\frac{\phi}{\phi_0} \right) + \frac{\lambda^2}{2} \sin^2 \left(\frac{\phi}{\phi_0} \right) + 1 \right)$$

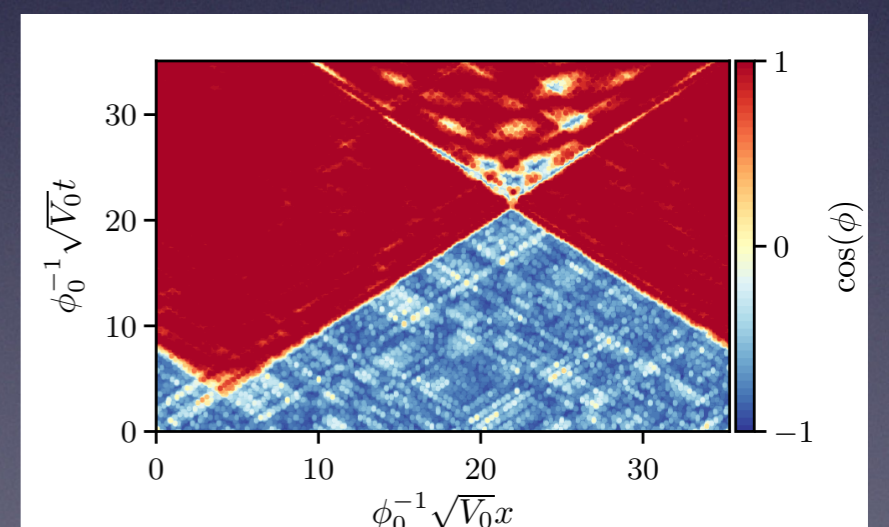
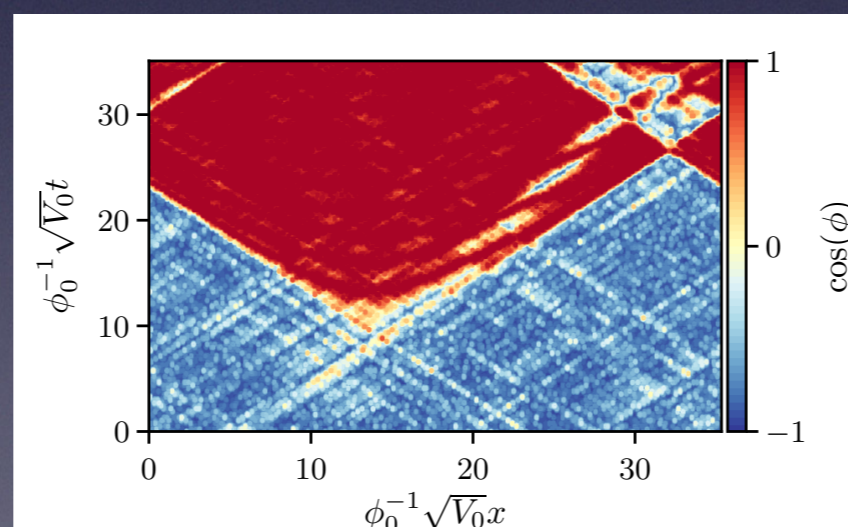
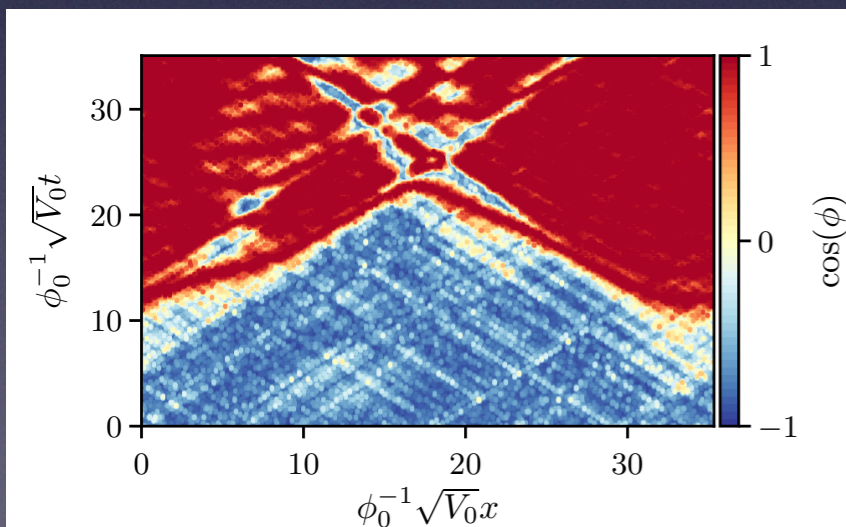
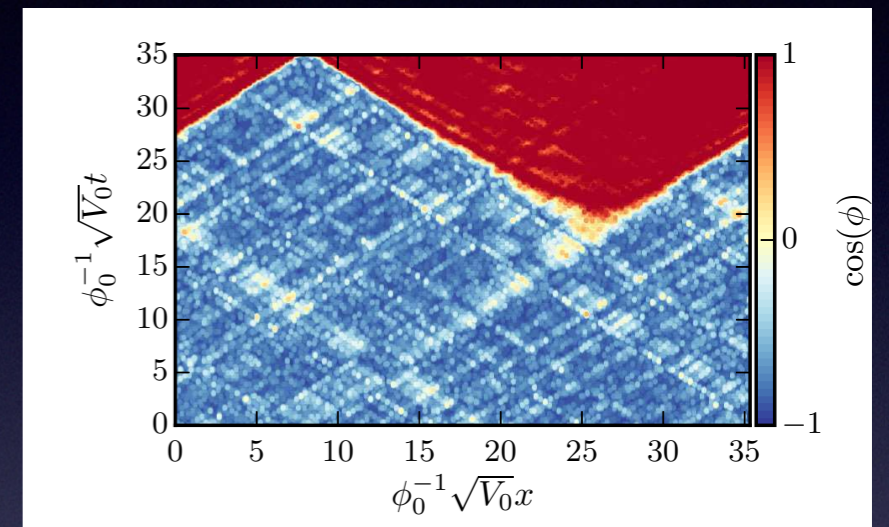
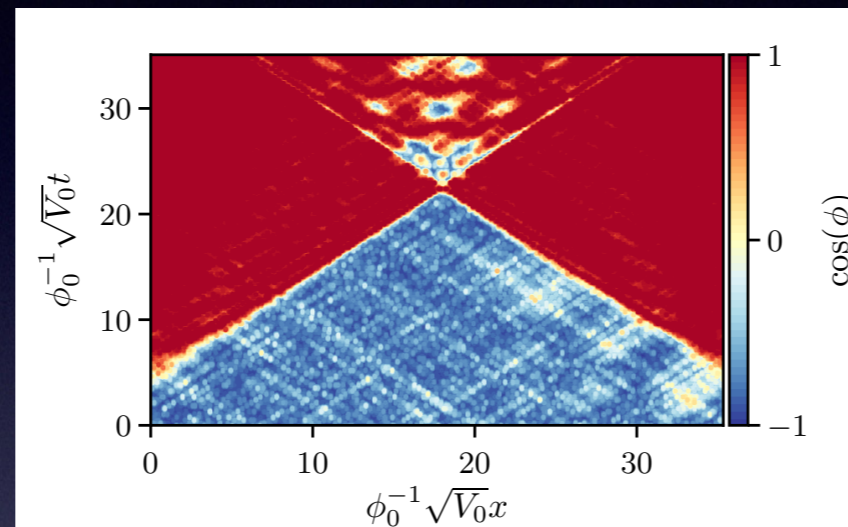
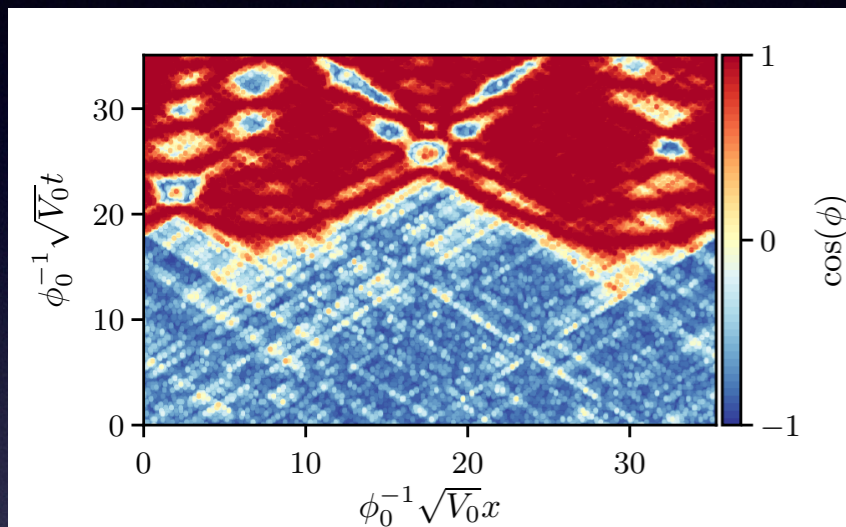






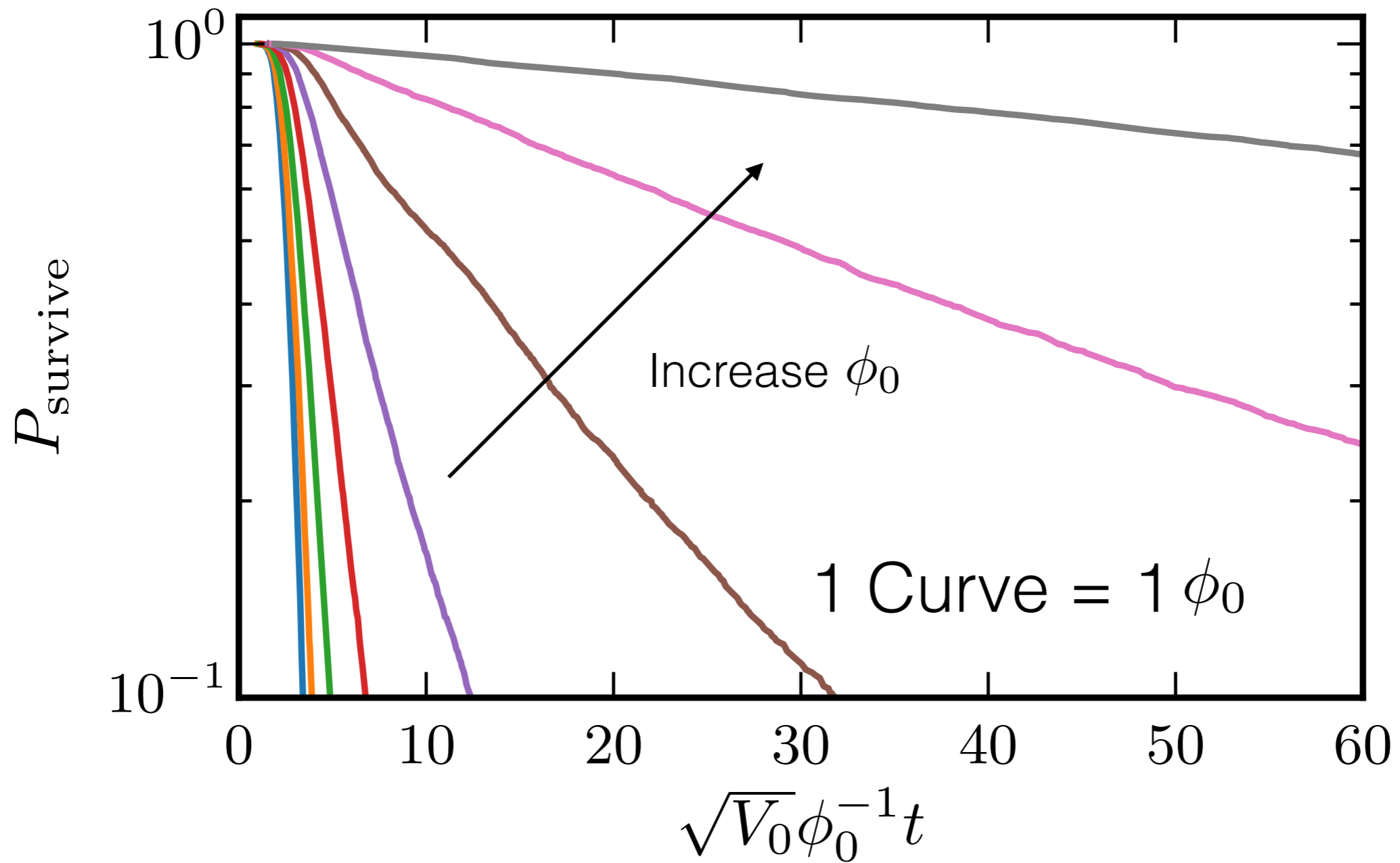
$t_{\text{decay}}^{(i)}$

Monte Carlo Over ICs



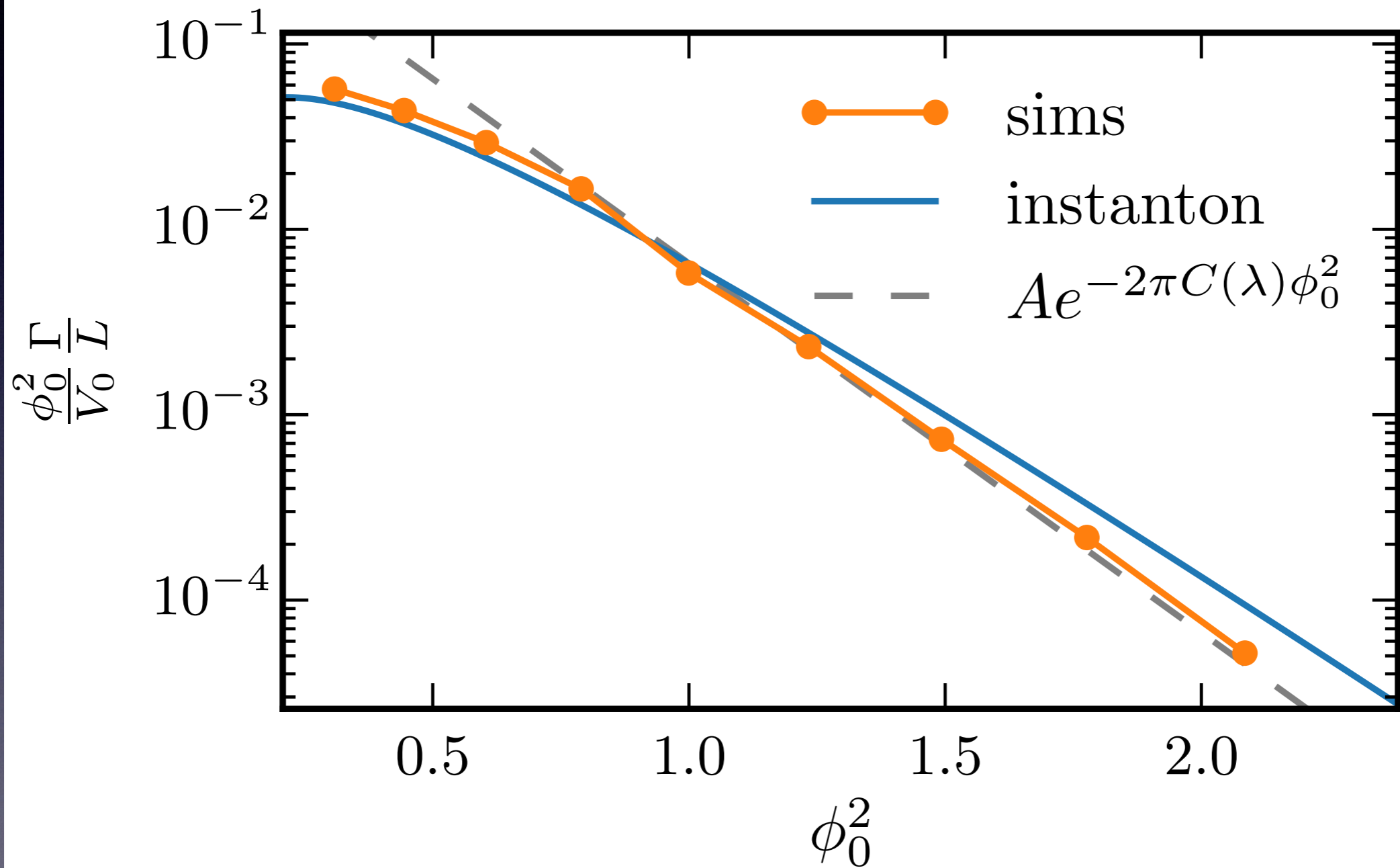
10000 sims / model x 1000 models \sim 10 million sims

$$P_{\text{survive}} \sim e^{-\Gamma(t-t_0)}$$



Sanity Check : $\Gamma \propto L$

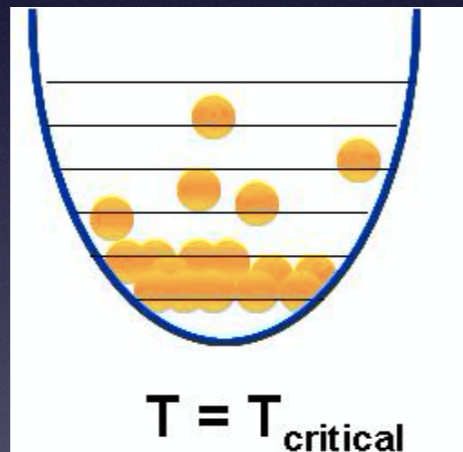
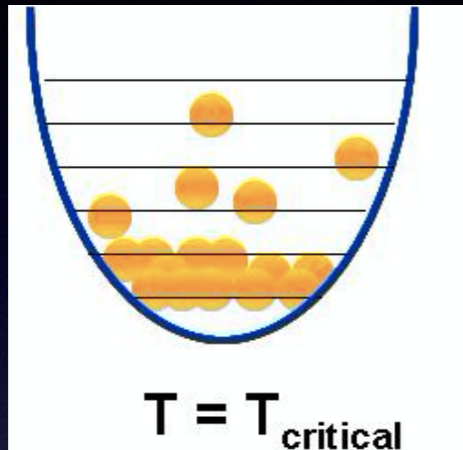
$$\frac{\Gamma_I^{(1+1)}}{L} = g(\lambda, \phi_0) m_{\text{eff}}^2 \phi_0^2 e^{-2\pi \phi_0^2 C(\lambda)}$$



Can We Test This?

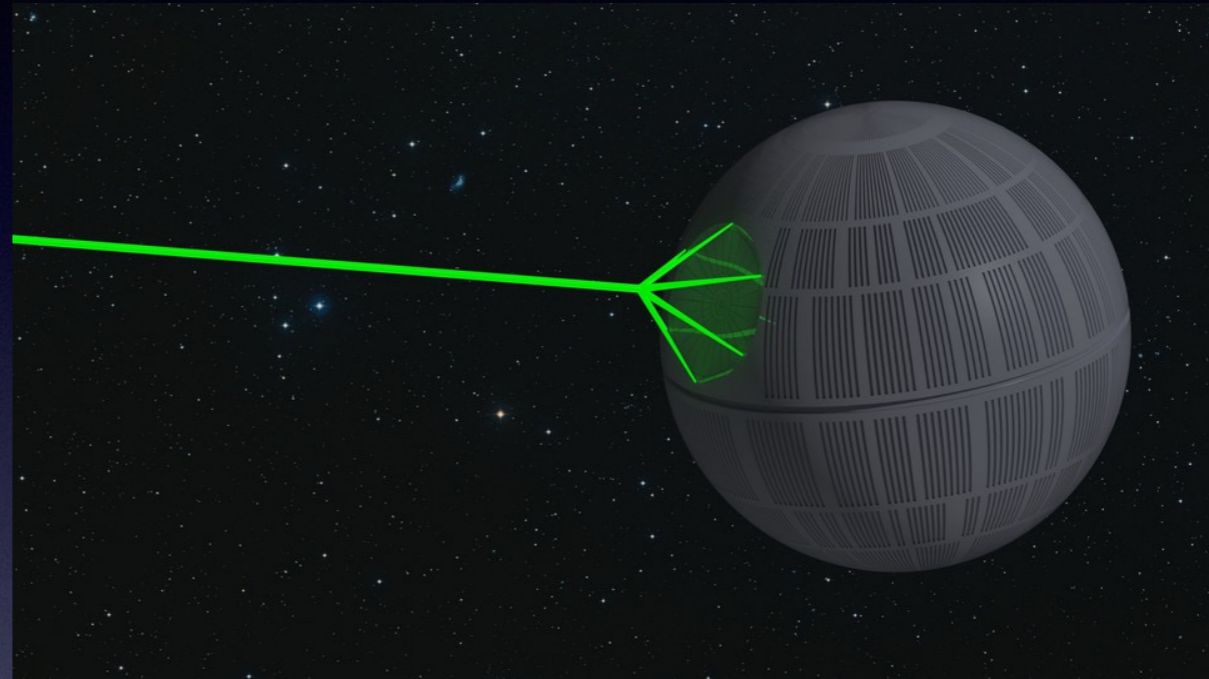
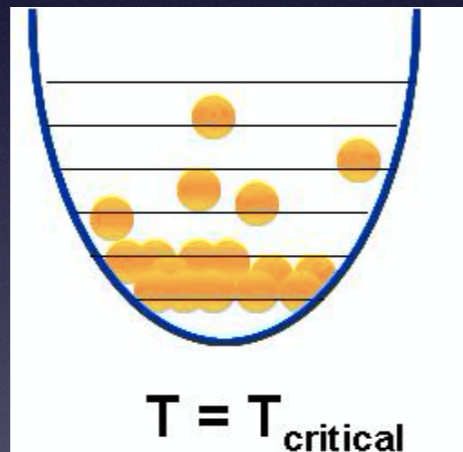
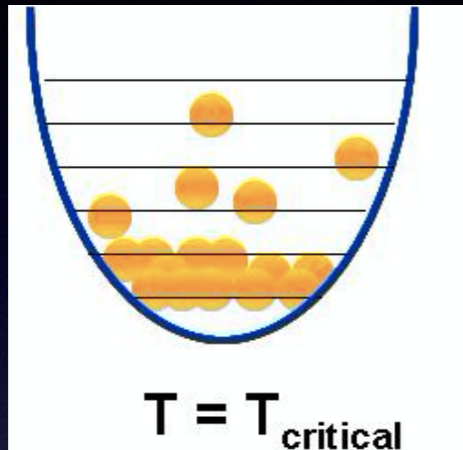
Analog Cold Atom BEC

[JB, Johnson, Peiris, Weinfurtner, 1712.02356, 1904.07873
Fialko et al 1408.1163, 1607.01460]



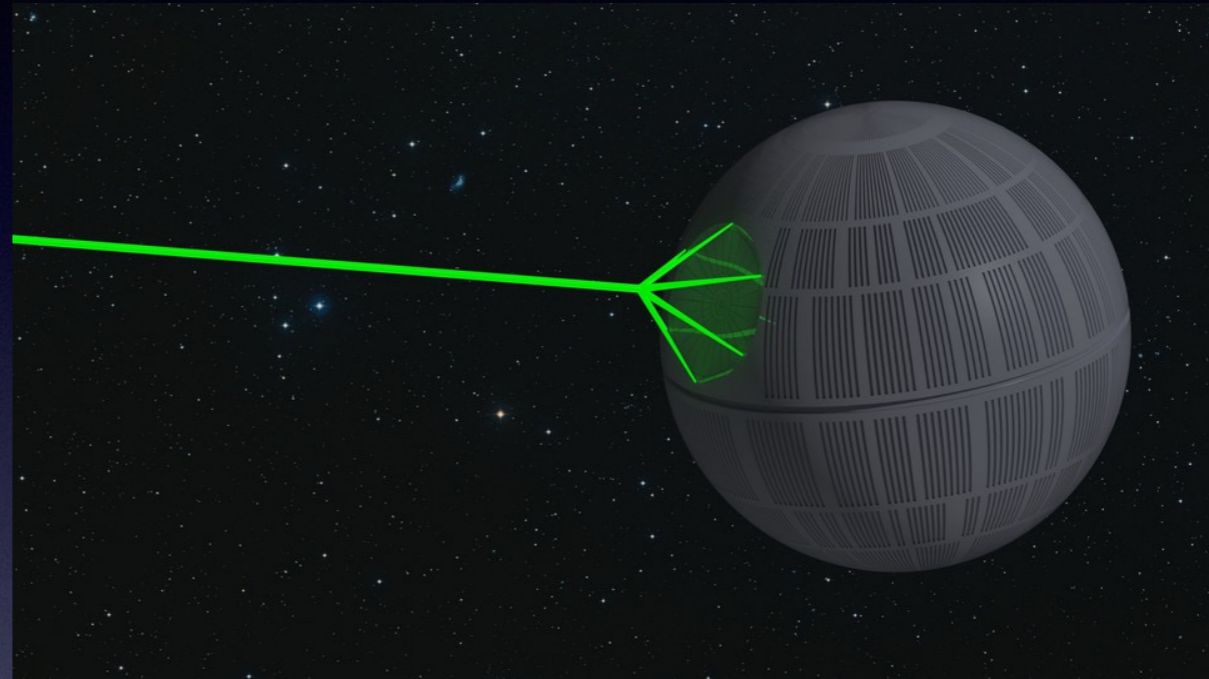
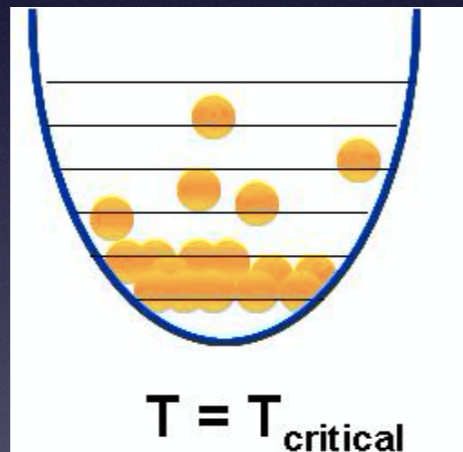
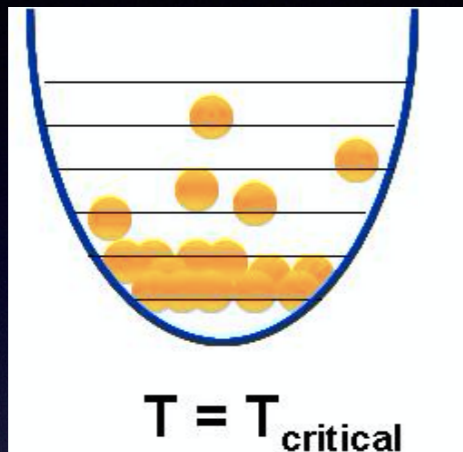
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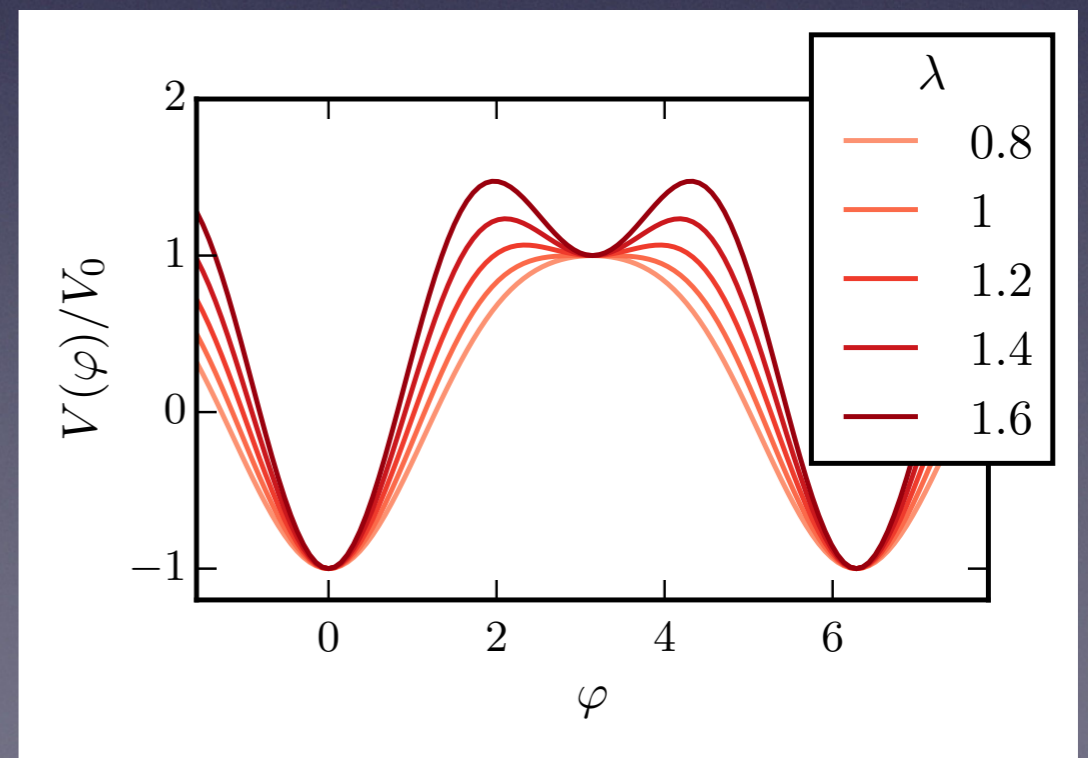


Analog Cold Atom BEC

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Dynamics of relative phase
is a relativistic field
with periodic potential

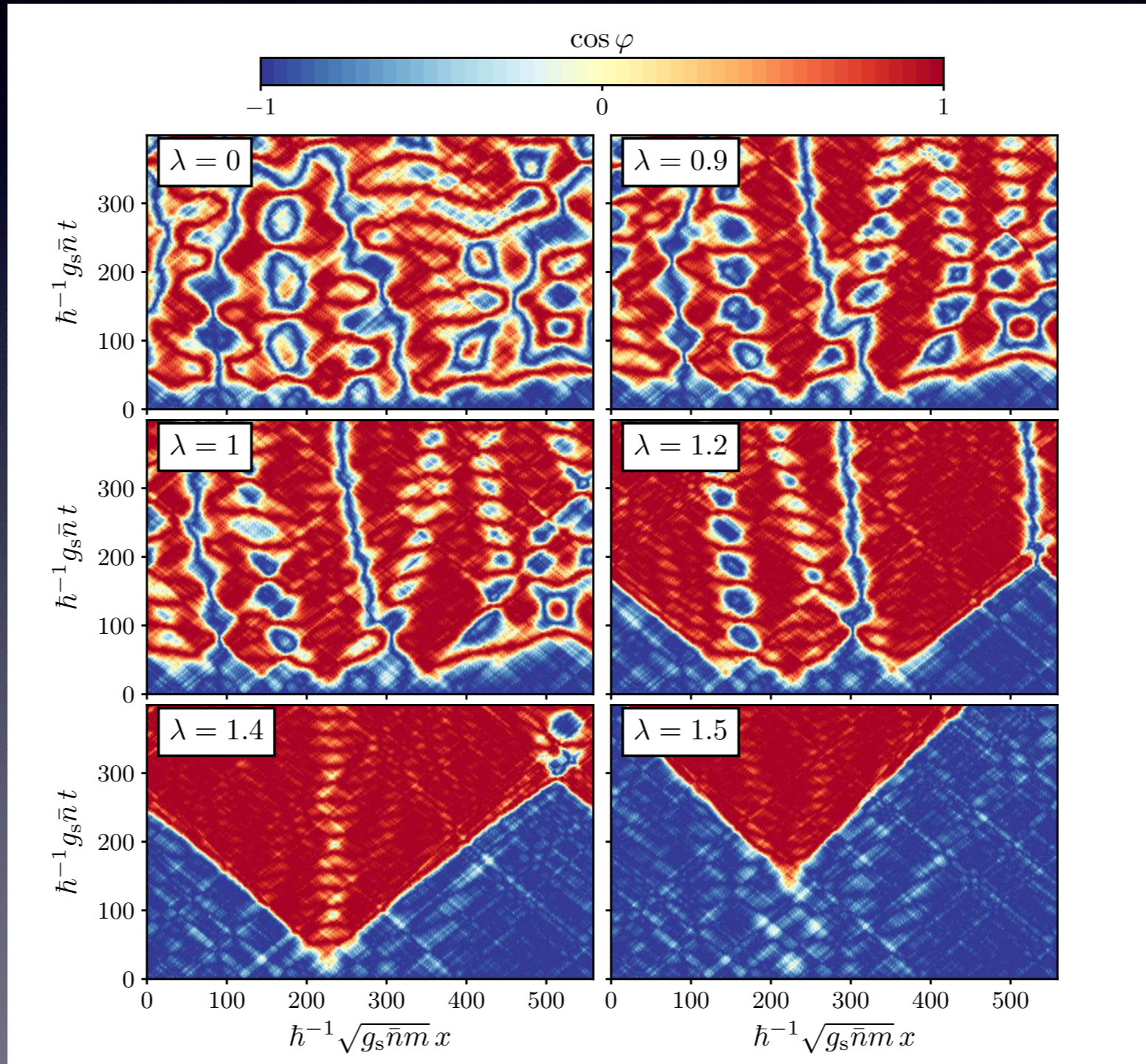


BEC Simulations

2nd-Order
Phase Transition

Rapid 1st-Order
Phase Transition

Slower 1st-Order
Phase Transition



Increase
Depth of
Minimum



Team

World-leading researchers in the following STFC and EPSRC areas:

■ Experimentalist

■ Theorist

■ Experimentalist/Theorist

Cosmology, Gravity and non-equilibrium Field Theory

- Jonathan Braden (Canada, CITA)
- Hiranya Peiris (UK, UCL)
- Andrew Pontzen (UK, UCL)
- Mathew Johnson (Canada, Perimeter Institute)
- Ian Moss (UK, Newcastle)
- Ruth Gregory (UK, Durham)
- Jorma Louko (UK, Nottingham)
- Ralf Schuetzhold (Germany, Helmholtz Centre)
- Bill Unruh (Canada, Vancouver)
- Silke Weinfurter (UK, Nottingham)

Ultracold Atoms

- Thomas Billam (UK, Newcastle)
- Zoran Hadzibabic (UK, Cambridge)
- Joerg Schmiedmayer (Austria, Vienna)

Superfluid 4He

- Carlo Barenghi (UK, Newcastle)
- John Owers-Bradley (UK, Nottingham)

Superfluid Nanofabrication

- Gregoire Ithier (UK, Royal Holloway London)
- Xavier Rojas (UK, Royal Holloway London)

Opto-mechanics

- Pierre Verlot (UK, Nottingham University)

Quantum Optics

- Friedrich Koenig (UK, St. Andrews University)

More info at
www.qsimfp.org

Conclusions

Physical Process: False Vacuum Decay

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- False Vacuum decay **can** occur via classical time-evolution (quantum is in initial state)

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Physical Process: False Vacuum Decay

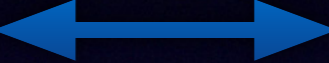
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- Exciting possibility to test in the lab

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Physical Process: False Vacuum Decay

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 - Complimentary to instanton (Euclidean rate wrong)
 - (Magic cancellation of amplitudes)
- Exciting possibility to test in the lab

Current/Future Work

- Real-time  Instanton
 - Renormalisation, Fluc. Determinant, Wigner
- Mean bubble profile = instanton? (preliminary yes)
- Bubble-bubble correlations?
- Time-dependent background or potential
- Non-vacuum initial states (pure or mixed)
- Application to many fields (more robust than instanton)

THANK YOU!

QFT in Phase Space

[See Braden et al, also Hertzberg]

Consider the Wigner functional

$$W[\phi, \Pi] \equiv \int \mathcal{D}\eta e^{-\frac{i}{\hbar} \int d^d x \Pi(\mathbf{x}) \eta(\mathbf{x})} \left\langle \phi + \frac{\eta}{2} \middle| \Psi \right\rangle \left\langle \Psi \middle| \phi - \frac{\eta}{2} \right\rangle$$

Important Properties

$$\int \mathcal{D}\phi \mathcal{D}\Pi W[\phi, \Pi] = 1$$

$$\langle \hat{\mathcal{O}}(\hat{\phi}, \hat{\Pi}) \rangle = \int \mathcal{D}\phi \mathcal{D}\Pi W(\phi, \Pi) \mathcal{O}_W(\phi, \Pi)$$

$W \sim$ quantum probability distribution

(caveat: Not positive definite in general,
but is for Gaussian states)

Wigner Approach

[see also Hertzberg and Yamada]

$$W[\phi, \Pi] \equiv \int \mathcal{D}\eta e^{-\frac{i}{\hbar} \int d^d x \Pi(\mathbf{x}) \eta(\mathbf{x})} \left\langle \phi + \frac{\eta}{2} \middle| \Psi \right\rangle \left\langle \Psi \middle| \phi - \frac{\eta}{2} \right\rangle$$

$$\left[\frac{\partial}{\partial t} + \int d^d x \left(\Pi \frac{\delta}{\delta \phi} + \nabla^2 \phi \frac{\delta}{\delta \Pi} - \frac{2}{i\hbar} V(\phi) \sin \left(\overleftarrow{\nabla}_\phi \frac{i\hbar}{2} \overrightarrow{\frac{\partial}{\partial \Pi}} \right) \right) \right] W[\phi(x), \Pi(x); t] = 0$$

Wigner Approach

[see also Hertzberg and Yamada]

$$W[\phi, \Pi] \equiv \int \mathcal{D}\eta e^{-\frac{i}{\hbar} \int d^d x \Pi(\mathbf{x}) \eta(\mathbf{x})} \left\langle \phi + \frac{\eta}{2} \middle| \Psi \right\rangle \left\langle \Psi \middle| \phi - \frac{\eta}{2} \right\rangle$$

$$\left[\frac{\partial}{\partial t} + \int d^d x \left(\dot{\phi} \frac{\delta}{\delta \phi} + \dot{\Pi} \frac{\delta}{\delta \Pi} \right) + \mathcal{O} \left(\hbar^2 V'''(\phi) \frac{\delta^3}{\delta \Pi^3} \right) \right] W[\phi(x), \Pi(x); t] = 0$$

Wigner Approach

[see also Hertzberg and Yamada]

$$W[\phi, \Pi] \equiv \int \mathcal{D}\eta e^{-\frac{i}{\hbar} \int d^d x \Pi(\mathbf{x}) \eta(\mathbf{x})} \left\langle \phi + \frac{\eta}{2} \middle| \Psi \right\rangle \left\langle \Psi \middle| \phi - \frac{\eta}{2} \right\rangle$$

$$\left[\frac{\partial}{\partial t} + \int d^d x \left(\dot{\phi} \frac{\delta}{\delta \phi} + \dot{\Pi} \frac{\delta}{\delta \Pi} \right) + \mathcal{O} \left(\hbar^2 V'''(\phi) \frac{\delta^3}{\delta \Pi^3} \right) \right] W[\phi(x), \Pi(x); t] = 0$$

Initial State (t=0)

Wigner Approach

[see also Hertzberg and Yamada]

$$W[\phi, \Pi] \equiv \int \mathcal{D}\eta e^{-\frac{i}{\hbar} \int d^d x \Pi(\mathbf{x}) \eta(\mathbf{x})} \left\langle \phi + \frac{\eta}{2} \middle| \Psi \right\rangle \left\langle \Psi \middle| \phi - \frac{\eta}{2} \right\rangle$$

$$\left[\frac{\partial}{\partial t} + \int d^d x \left(\dot{\phi} \frac{\delta}{\delta \phi} + \dot{\Pi} \frac{\delta}{\delta \Pi} \right) + \mathcal{O} \left(\hbar^2 V'''(\phi) \frac{\delta^3}{\delta \Pi^3} \right) \right] W[\phi(x), \Pi(x); t] = 0$$

Classical Evolution

Initial State (t=0)

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[see also Hertzberg and Yamada]

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