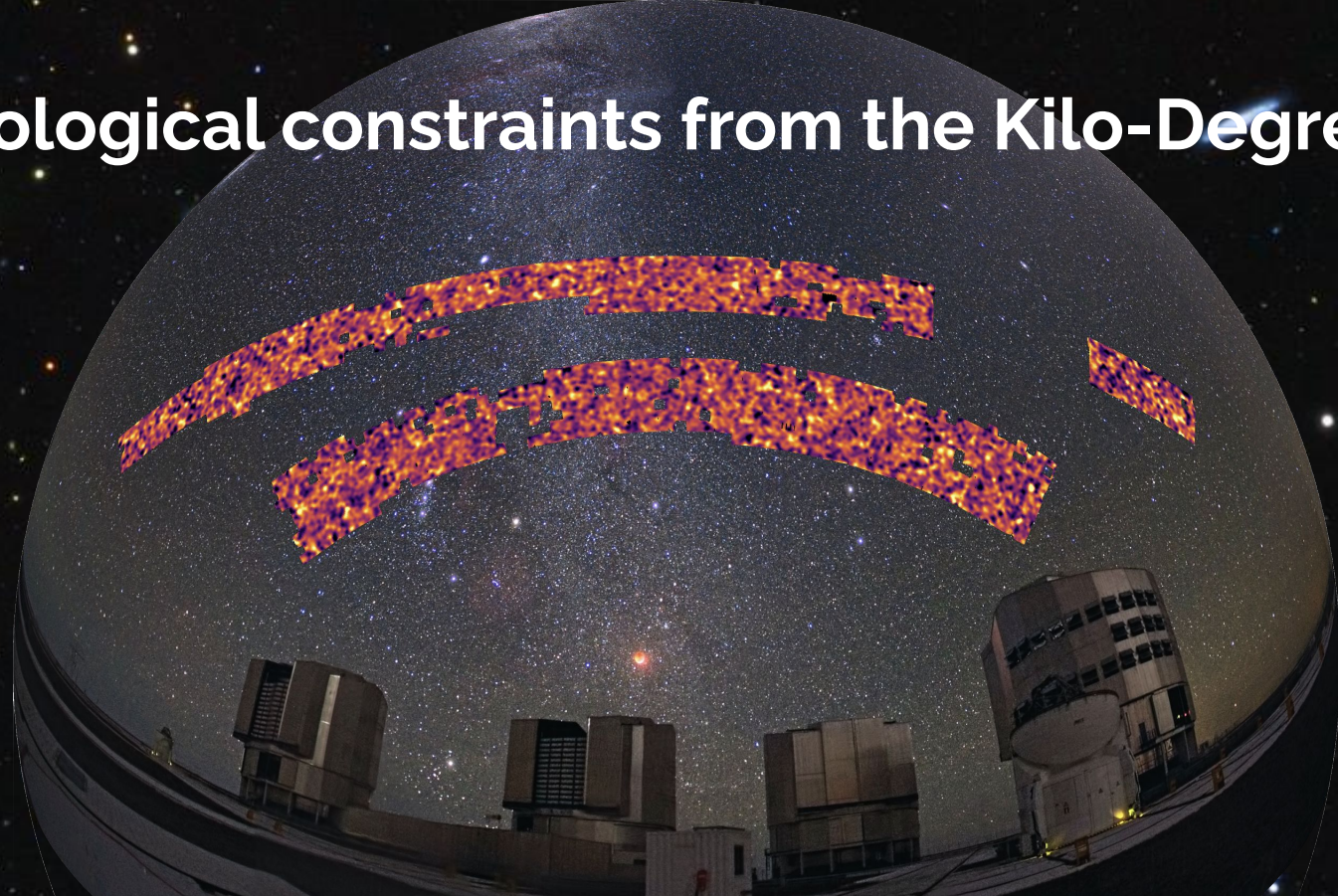


Cosmological constraints from the Kilo-Degree Survey



Catherine Heymans and Tilman Tröster

University of Edinburgh, Scotland
German Centre for Cosmological Lensing, RUB, Germany

KiDS-1000 Publications

Methodology: Joachimi, Lin, Asgari, Tröster, Heymans et al.

Photometric Redshifts: Hildebrandt, van den Busch, Wright et al.

Shear Measurements: Giblin, Heymans, Asgari et al.

Cosmic Shear Cosmology: Asgari, Lin, Joachimi et al.

3x2pt Cosmology: Heymans, Tröster et al.



Benjamin Joachimi



Chieh-An Lin

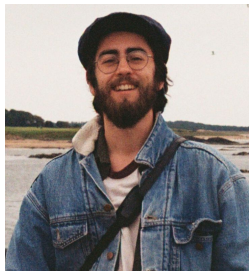


Jan Luca van den Busch



Angus H Wright

KiDS



Ben Giblin



Marika Asgari



Tilman Tröster



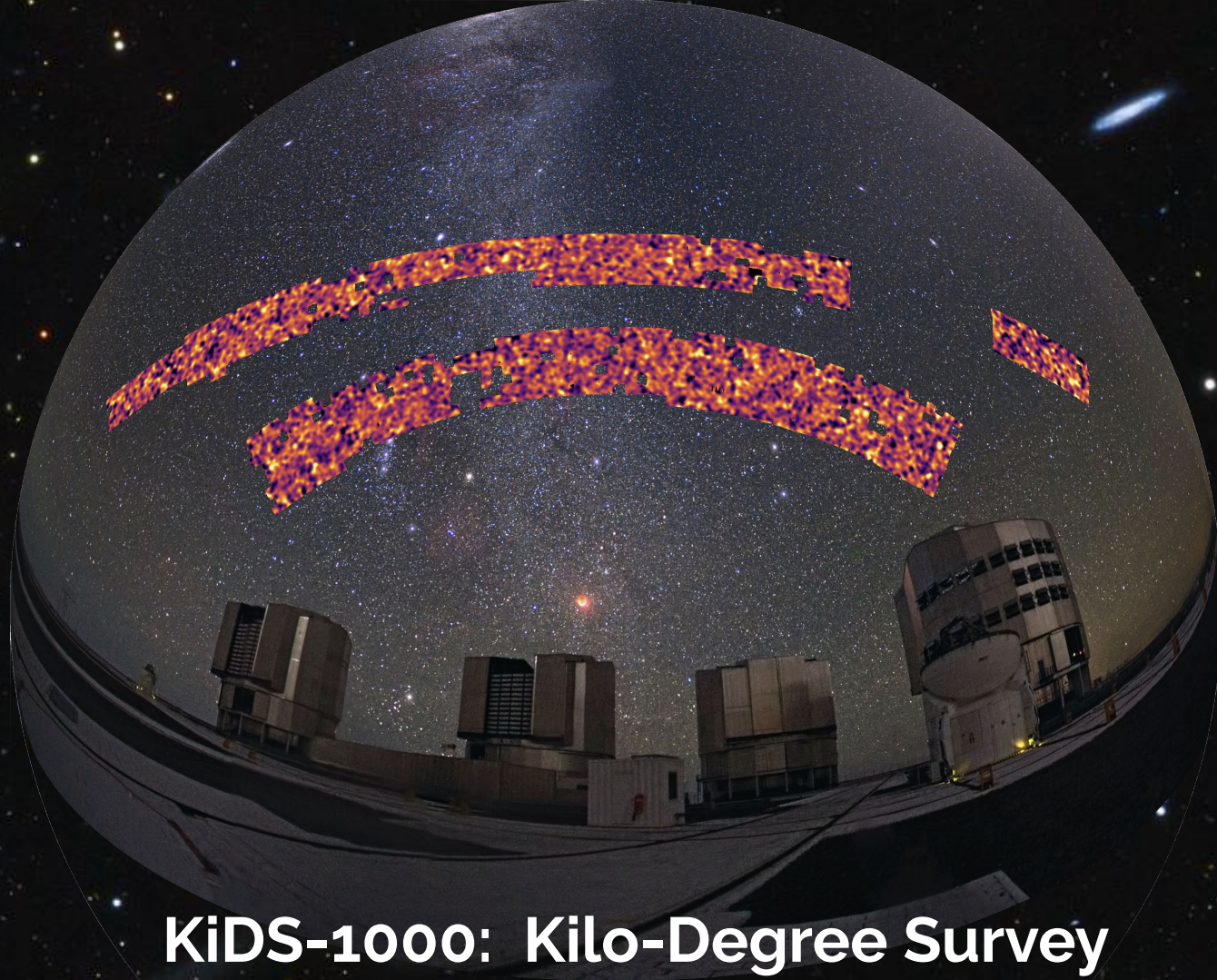
Konrad Kuijken



Hendrik Hildebrandt

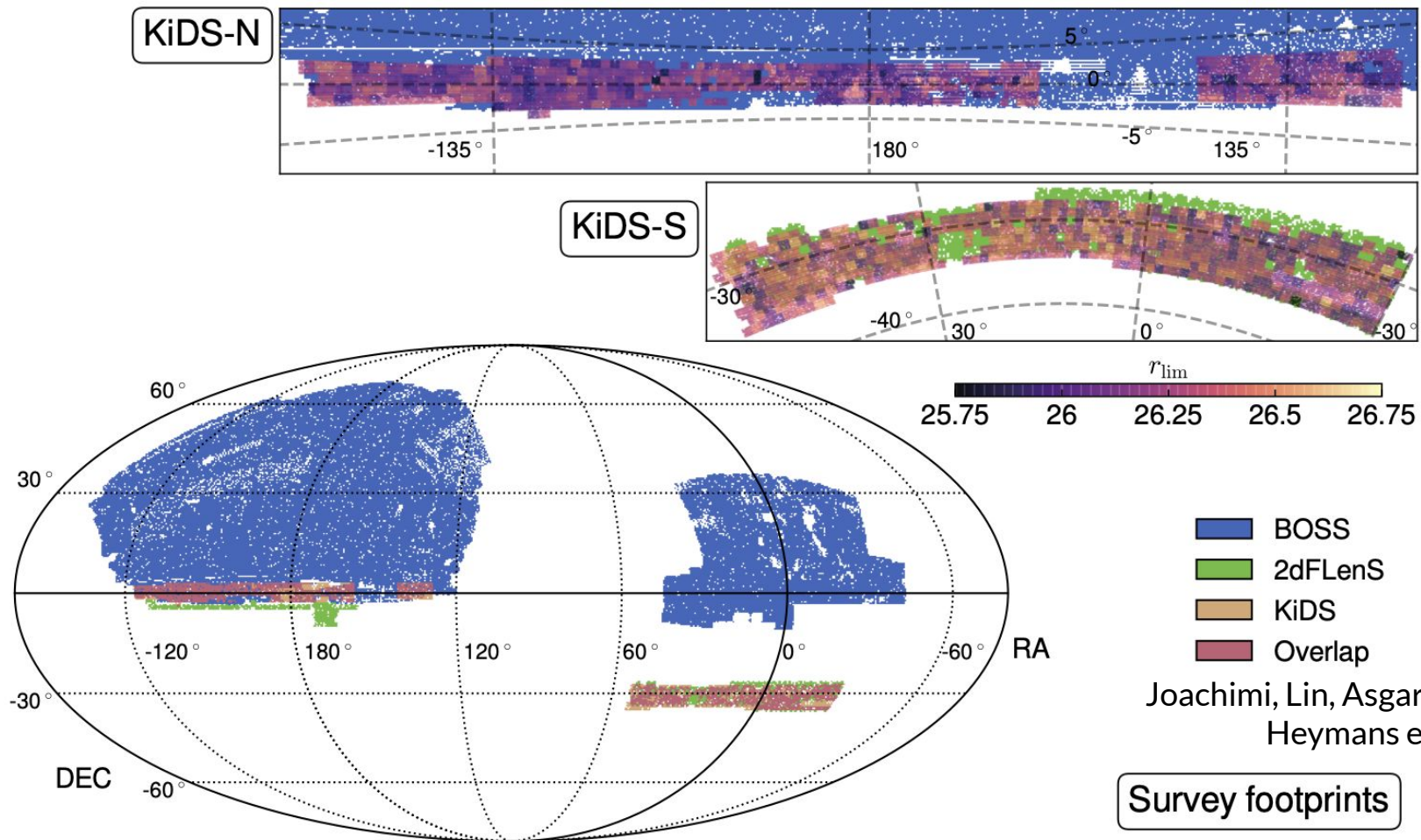


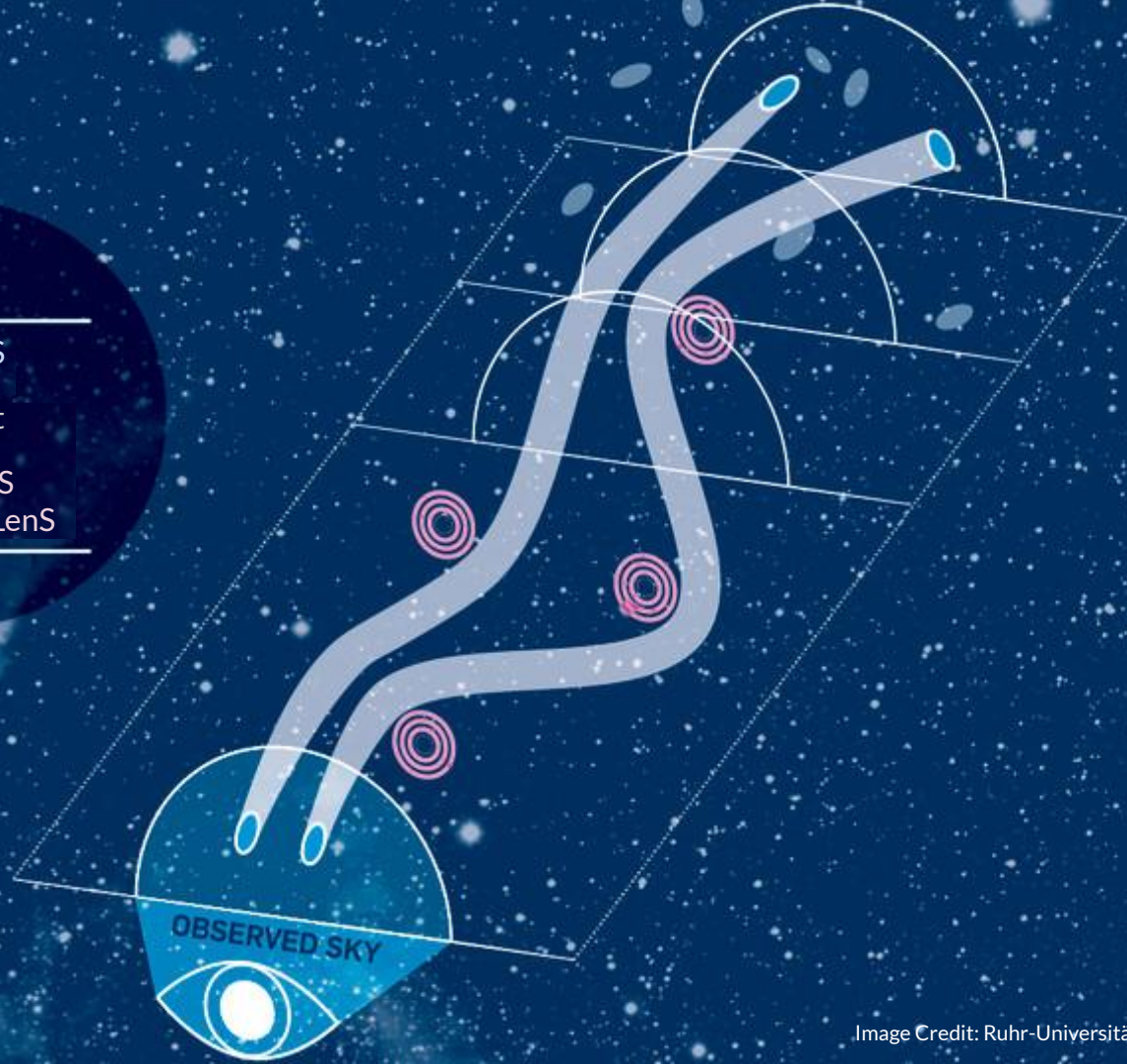
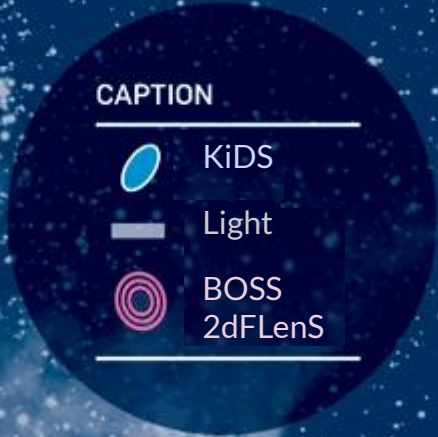
Catherine Heymans



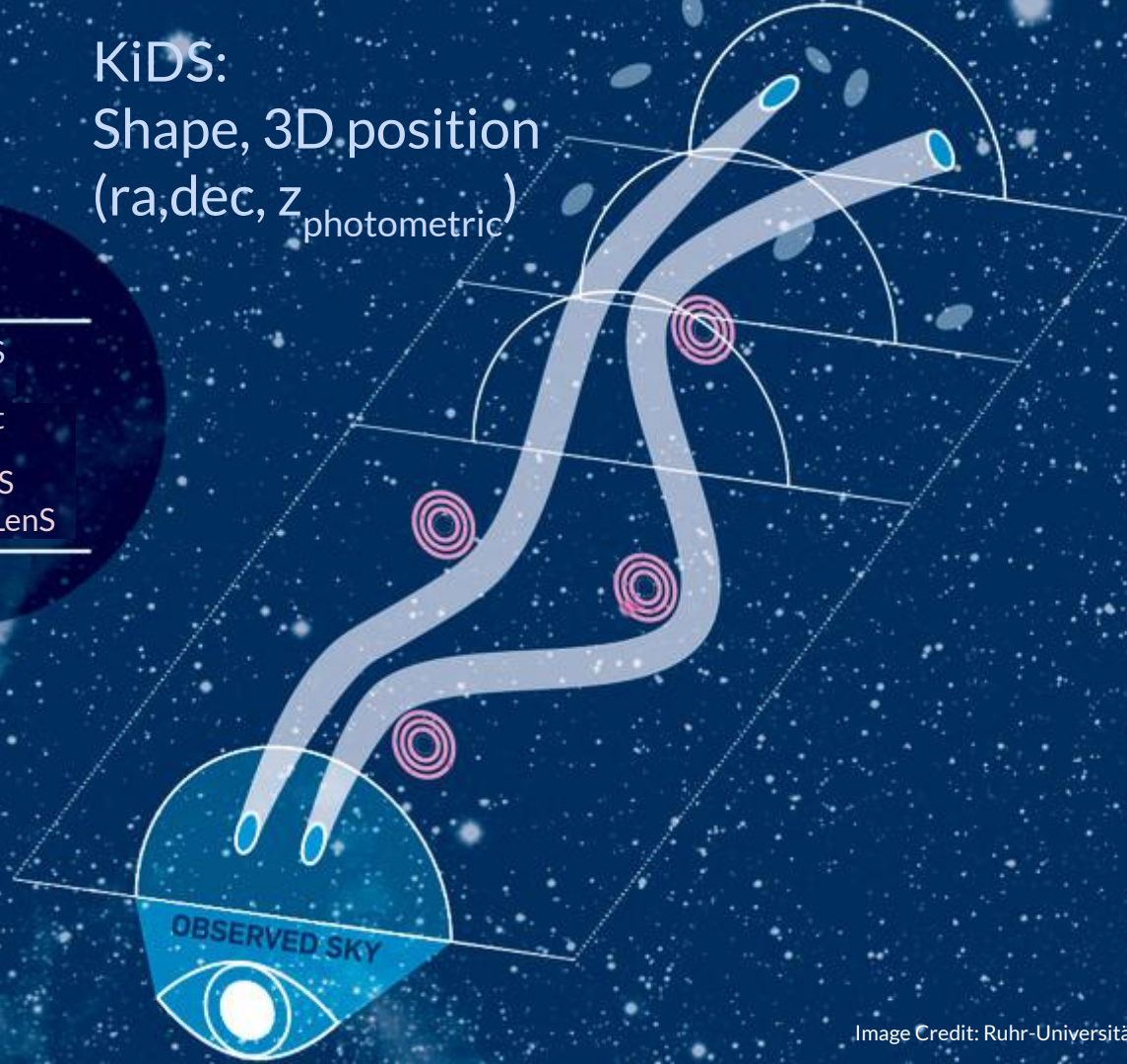
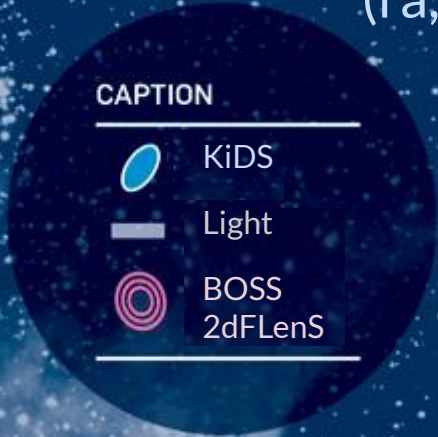
KiDS-1000: Kilo-Degree Survey

KiDS, BOSS and 2dFLenS

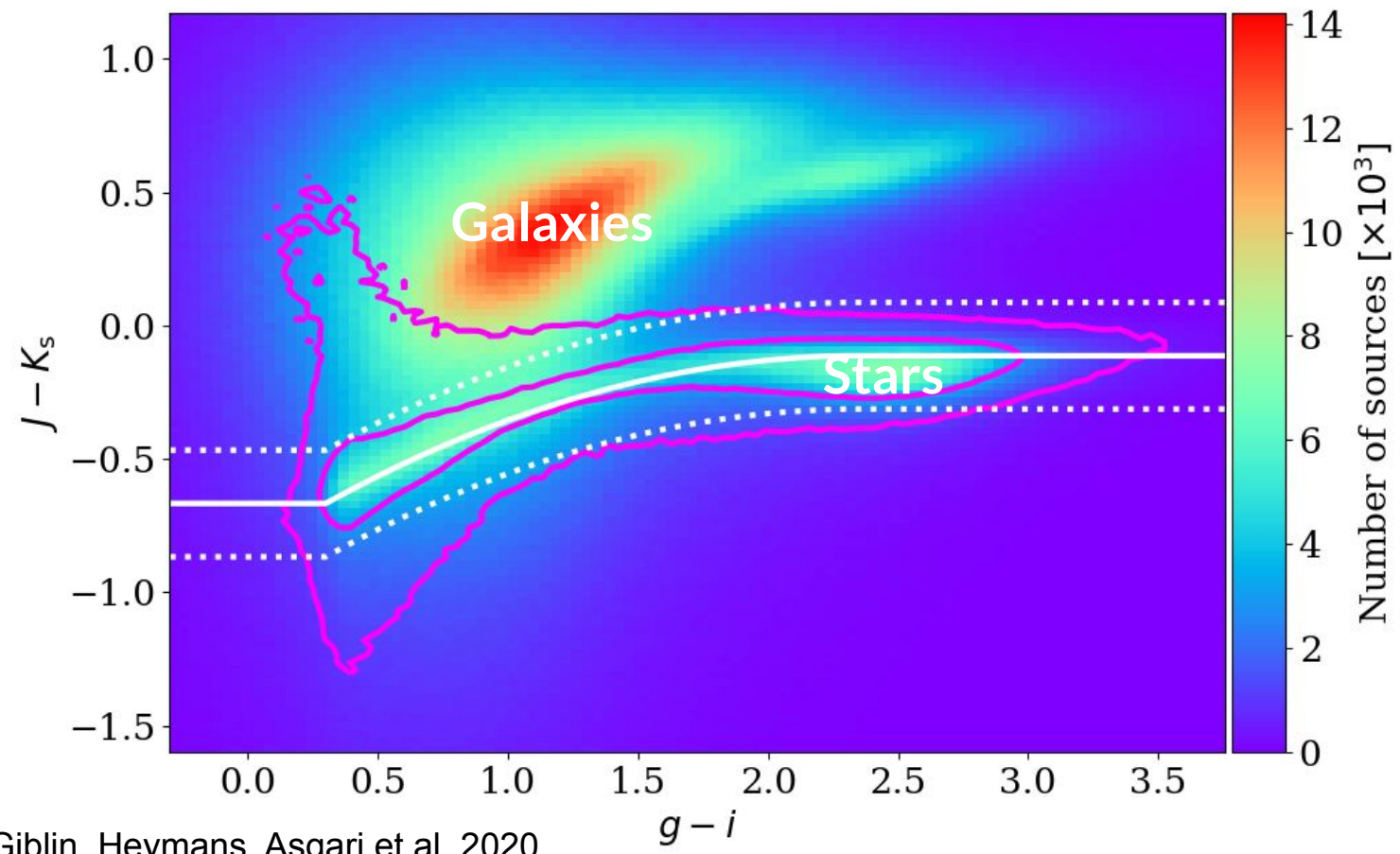




KiDS:
Shape, 3D position
($ra, dec, z_{\text{photometric}}$)



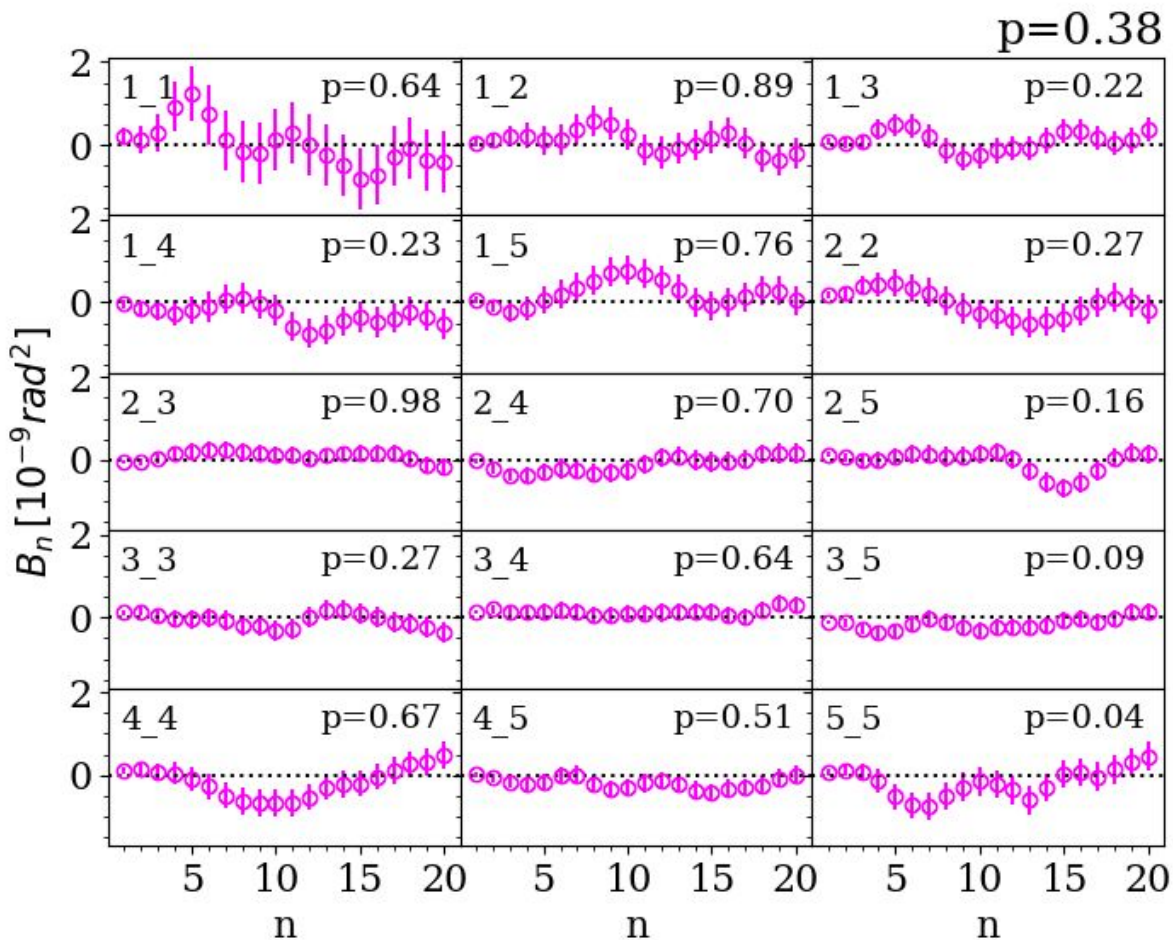
Stars for the PSF, Galaxies for the Cosmology



Giblin, Heymans, Asgari et al. 2020

Shear null tests

- B-modes consistent with pure noise
- PSF model accuracy size/shape requirements easily met
- Instrumental defects quantified
- Shear-ratio test passed

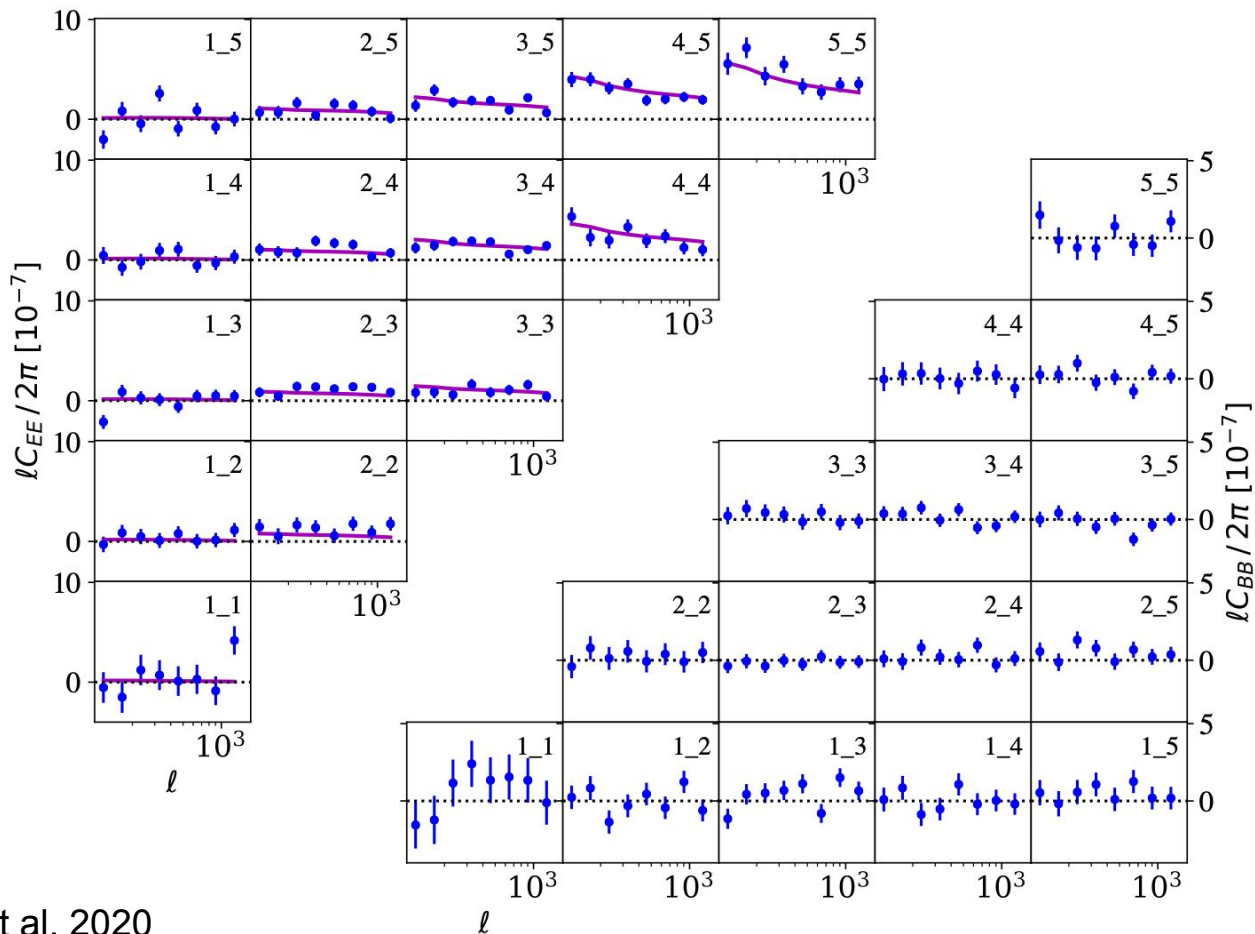


Cosmic Shear

Full Seminar: Check out Cosmology Talks YouTube Channel ⁹

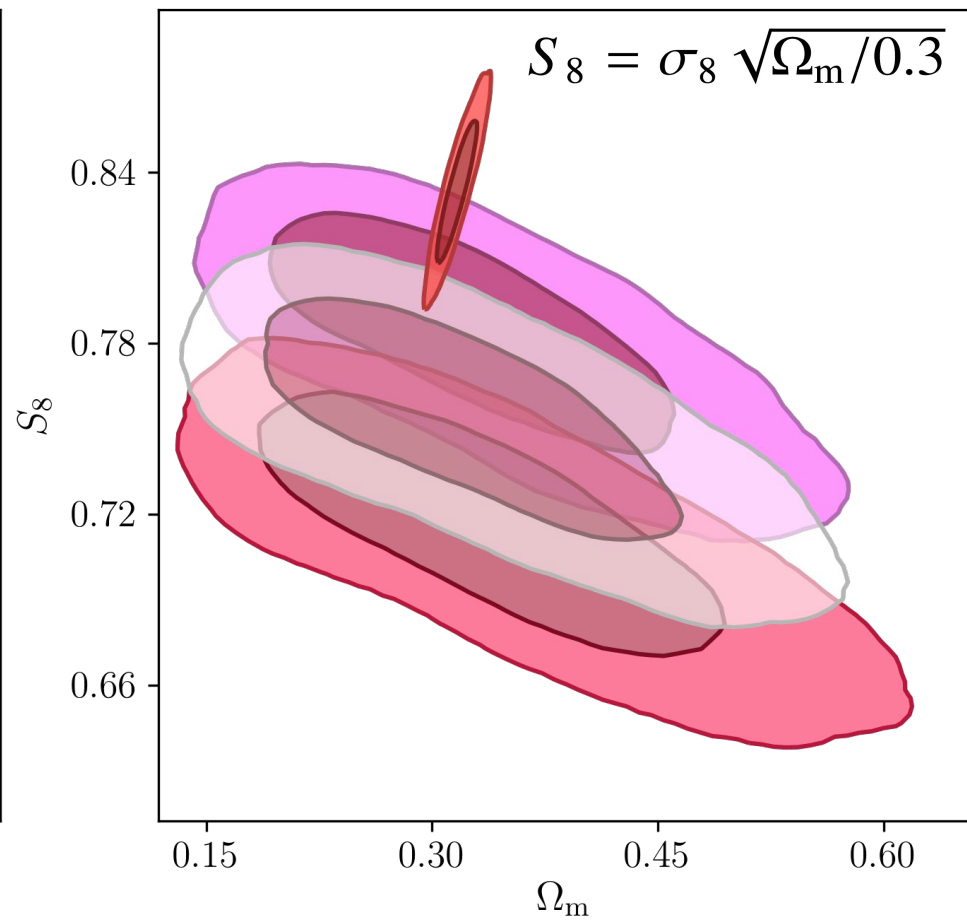
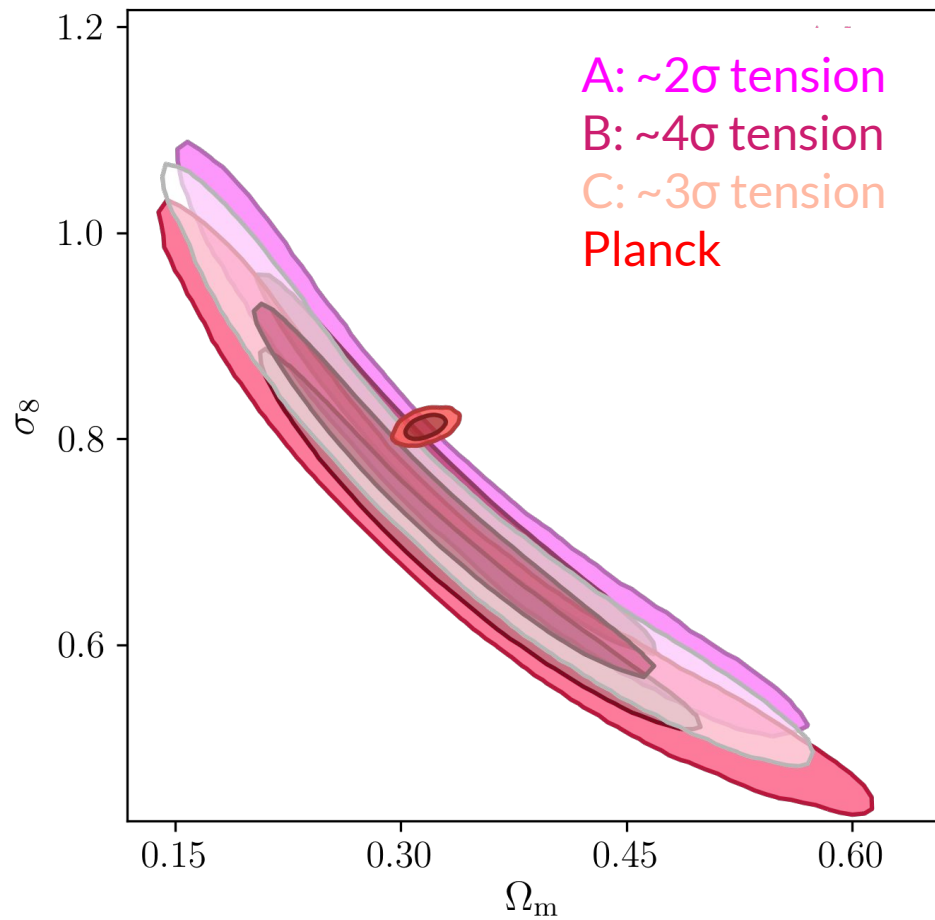
The theoretical model includes

- Flat Λ CDM
- Intrinsic Galaxy Alignments
- Baryon Feedback
- Photometric Redshift Calibration Uncertainty
- Shear Calibration Uncertainty



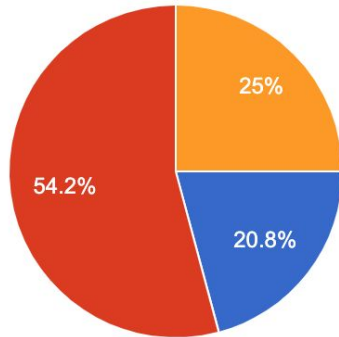
Asgari, Lin, Joachimi et al. 2020

Blind Analysis



Choose the one that you would like to be the truth.

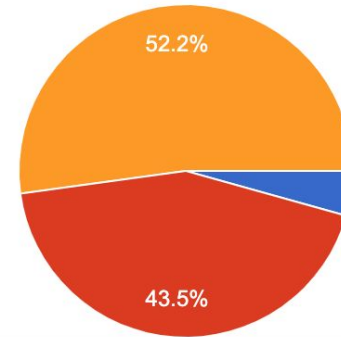
24 responses



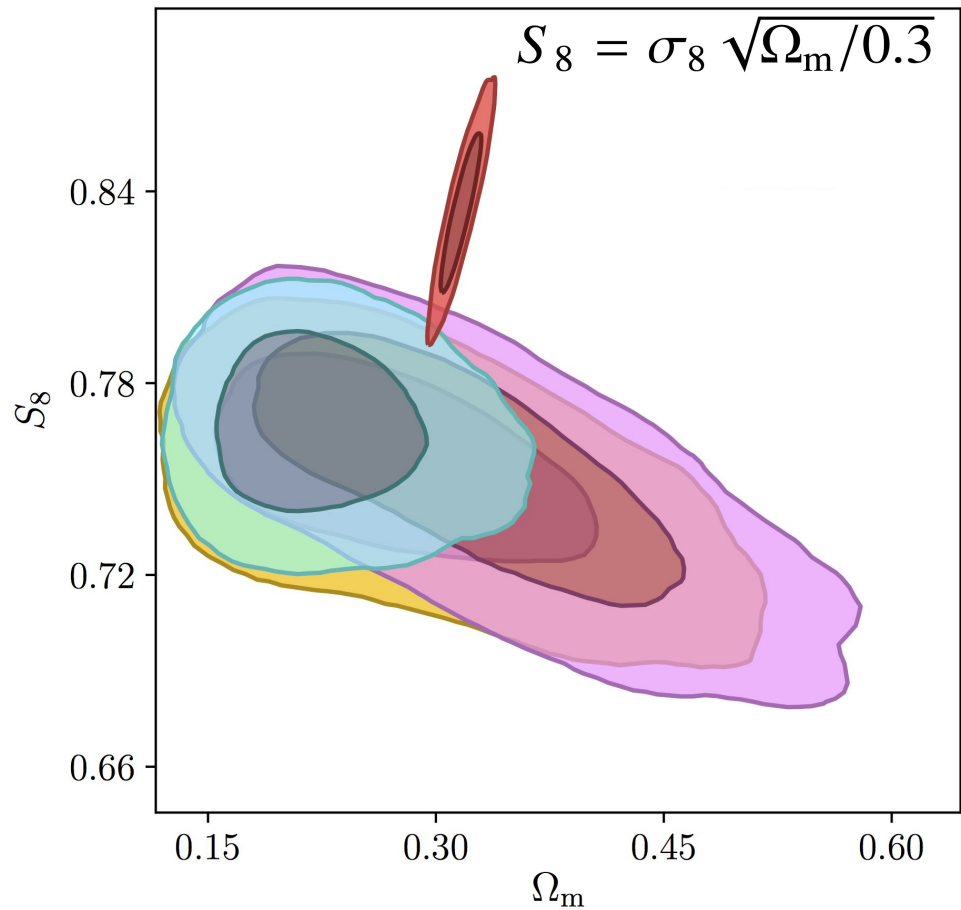
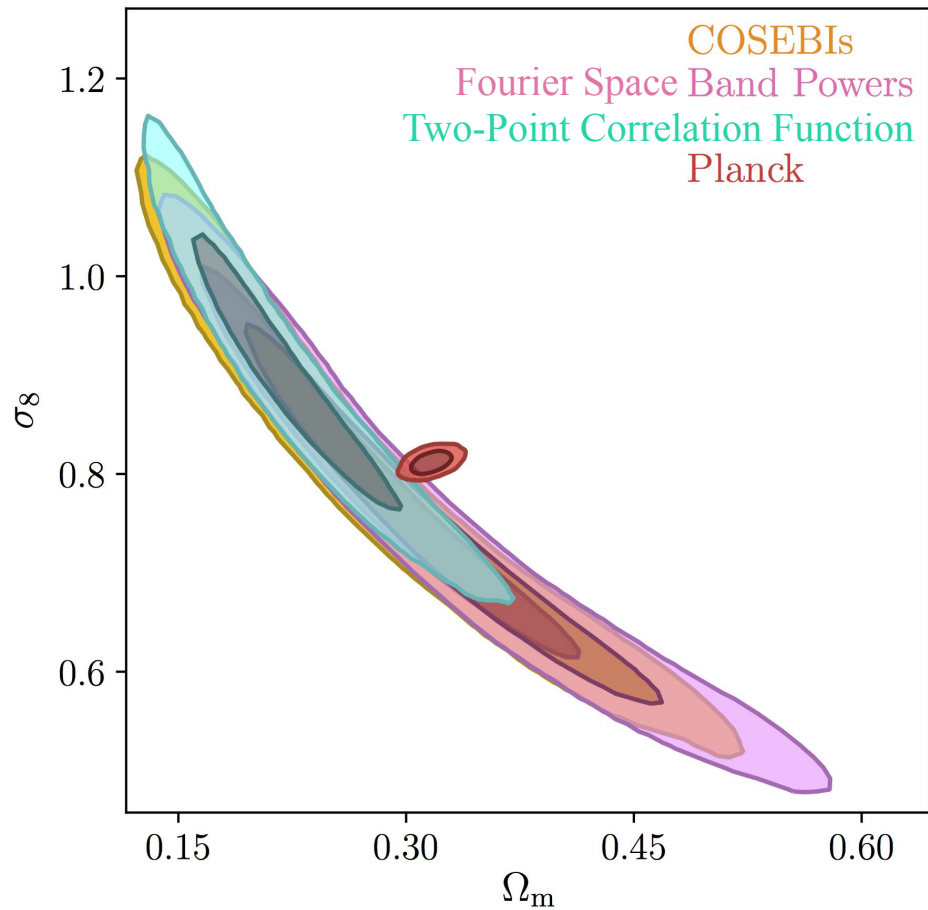
● A
● B
● C

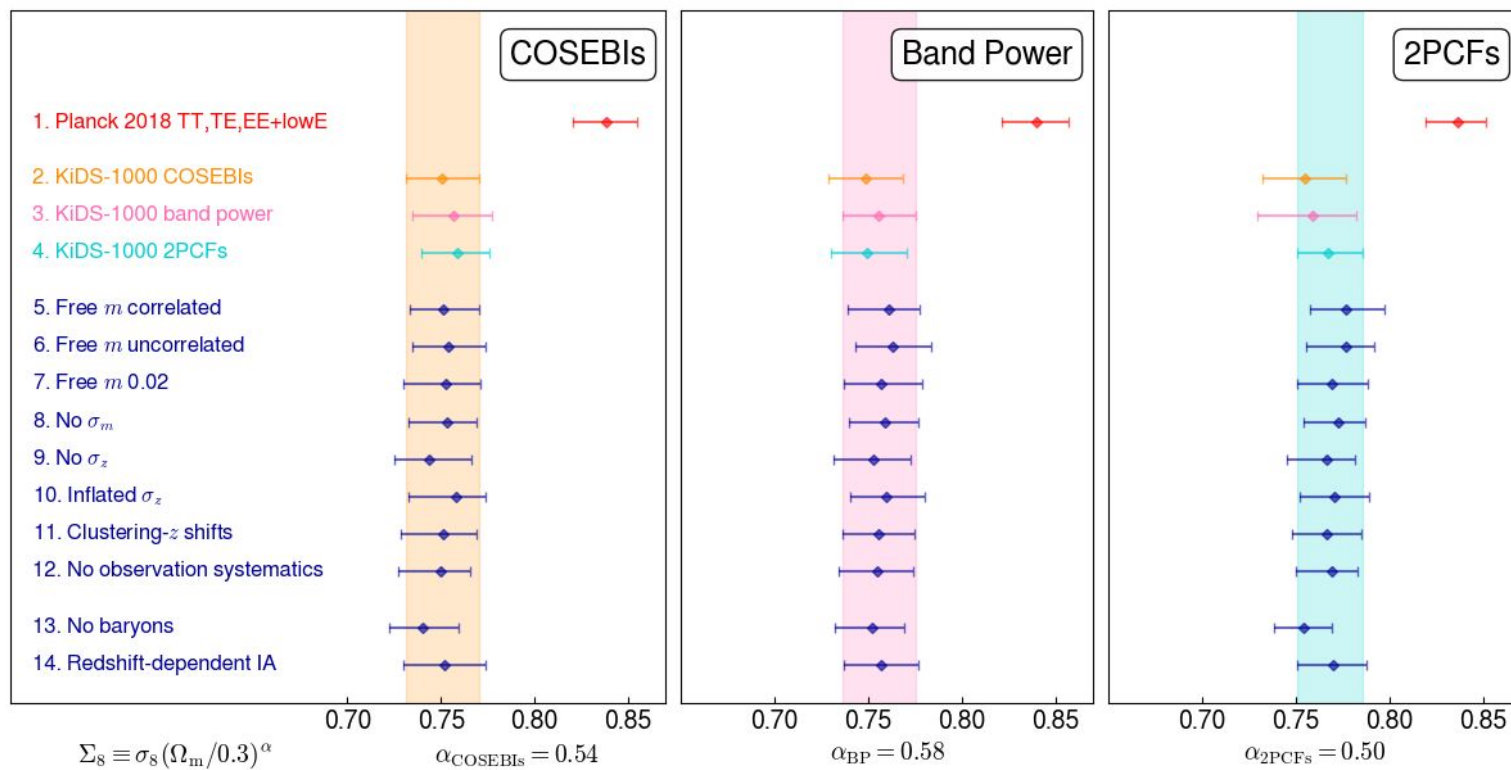
Which one do you think is the truth?

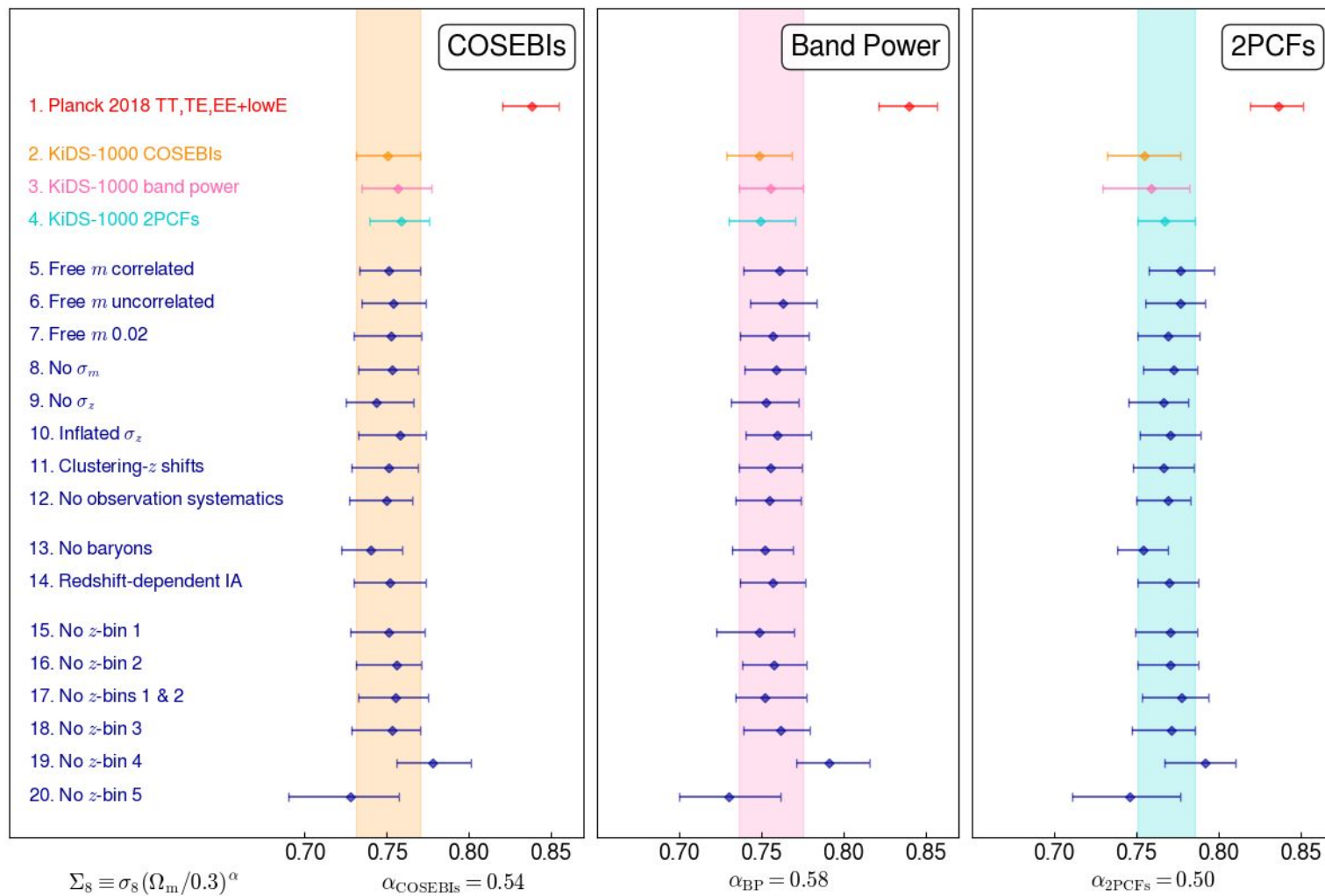
23 responses



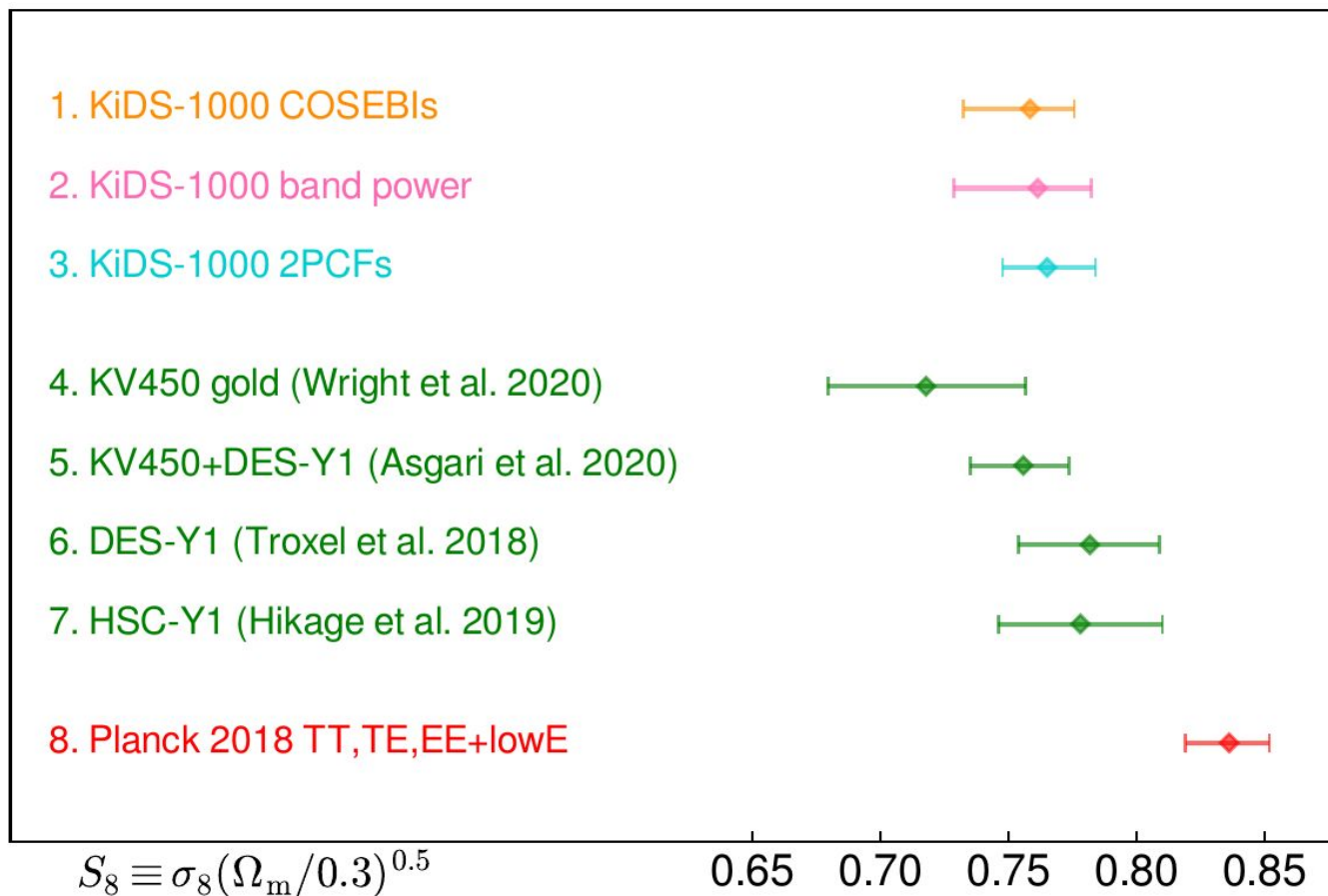
● A
● B
● C



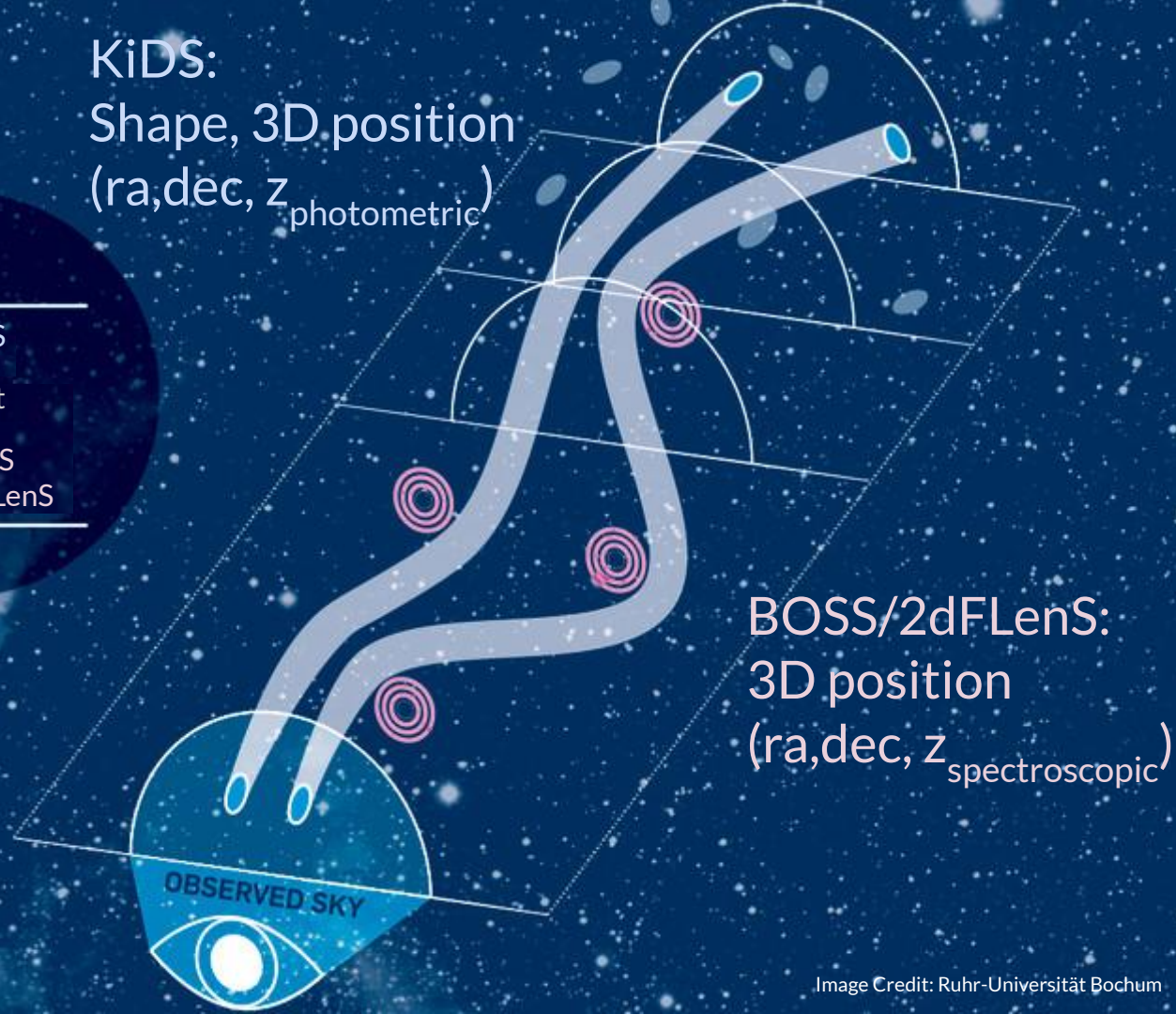
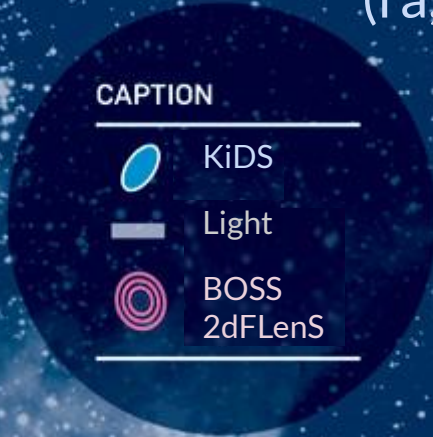




Comparison with Planck, DES and HSC

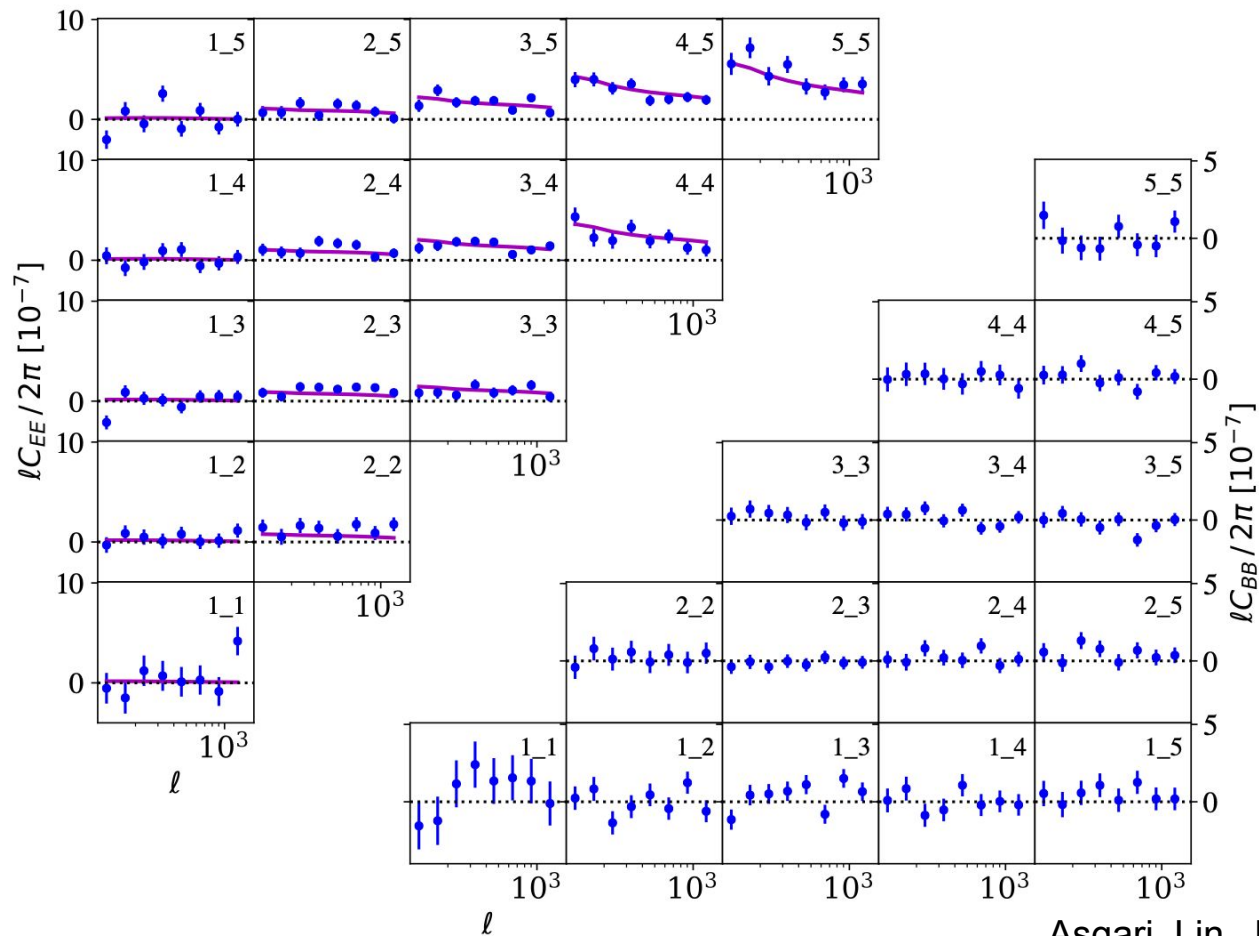


KiDS:
Shape, 3D position
($ra, dec, z_{\text{photometric}}$)



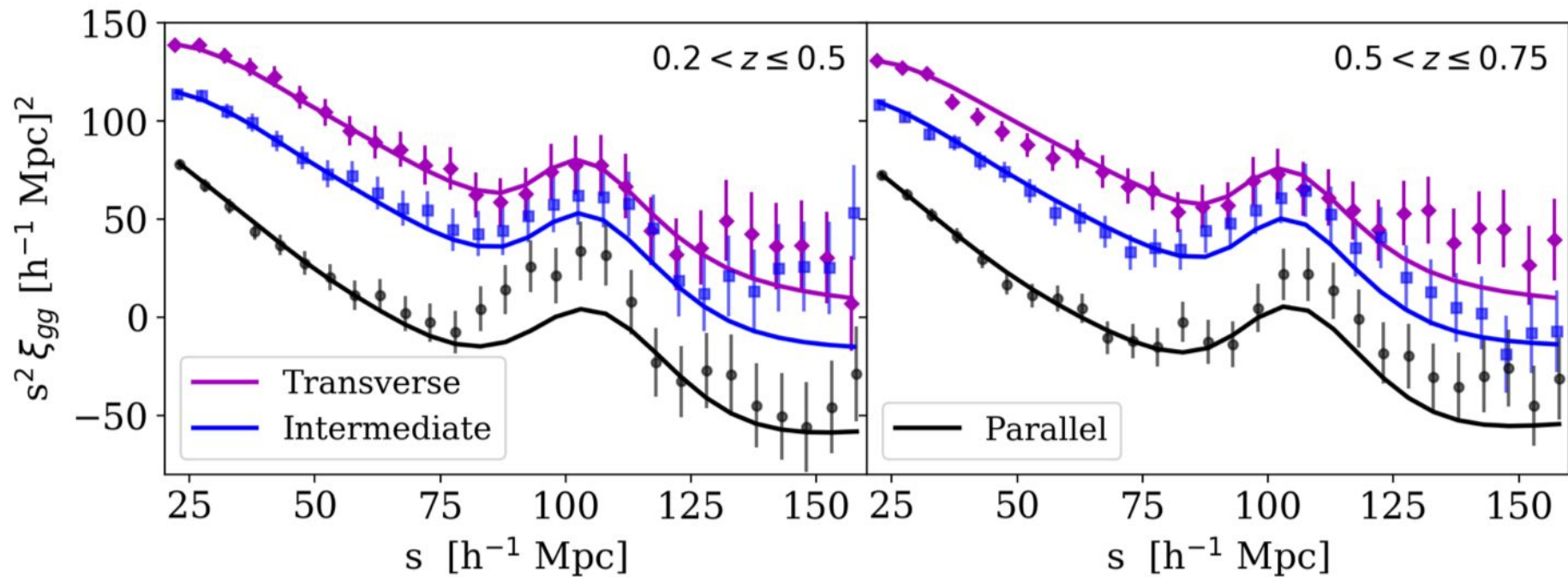
3x2pt: Cosmic Shear +

Band-Powers



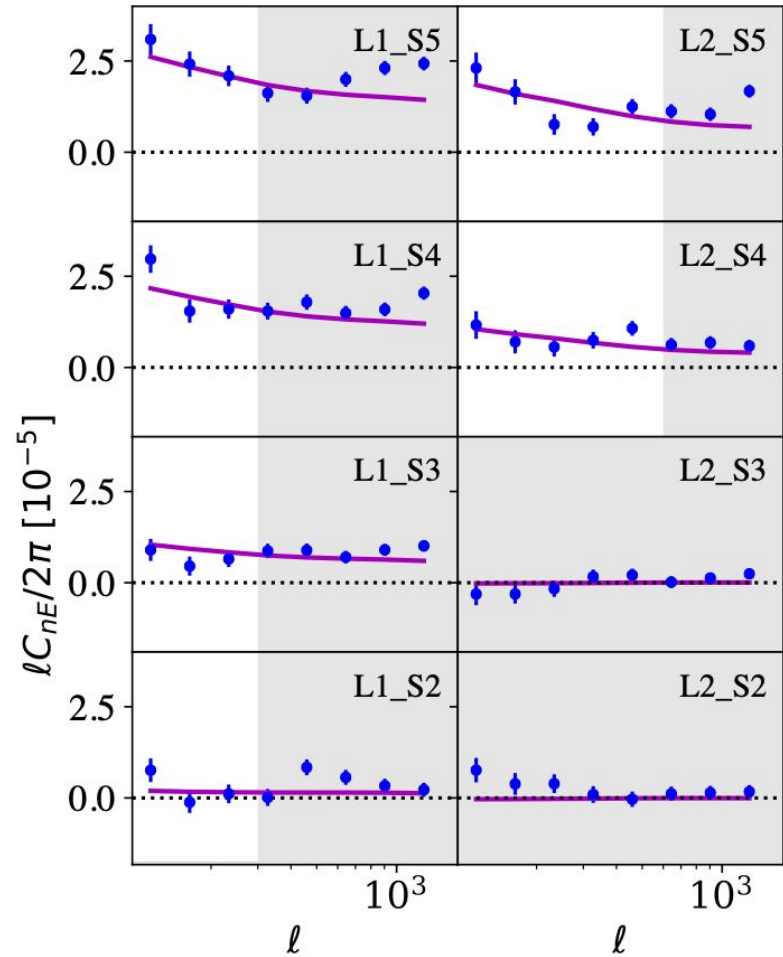
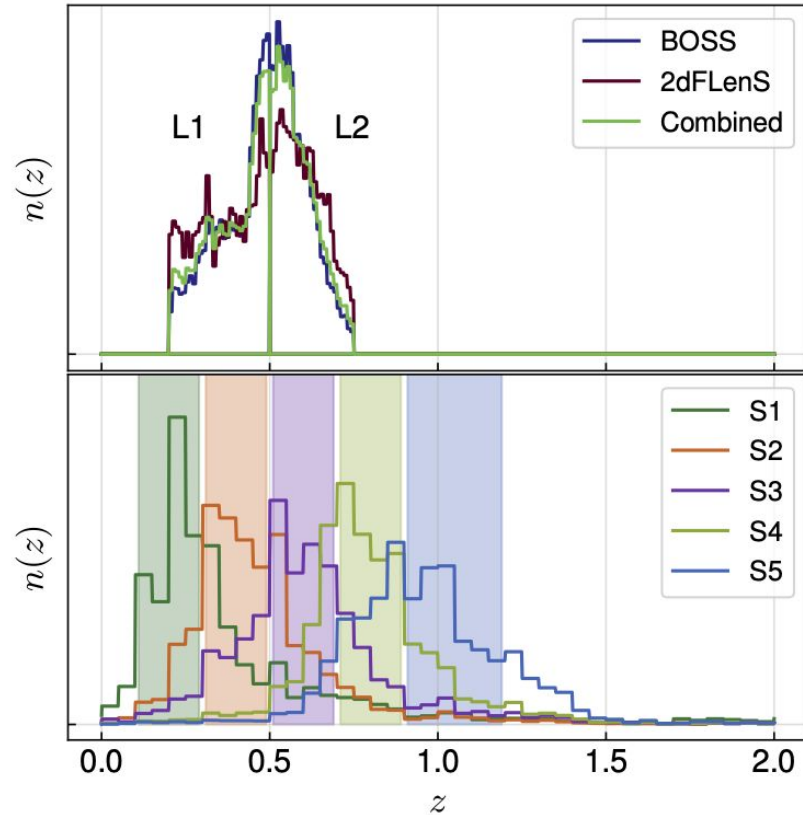
3x2pt: Cosmic Shear + Clustering +

Anisotropic Galaxy Clustering: RSD + BAO



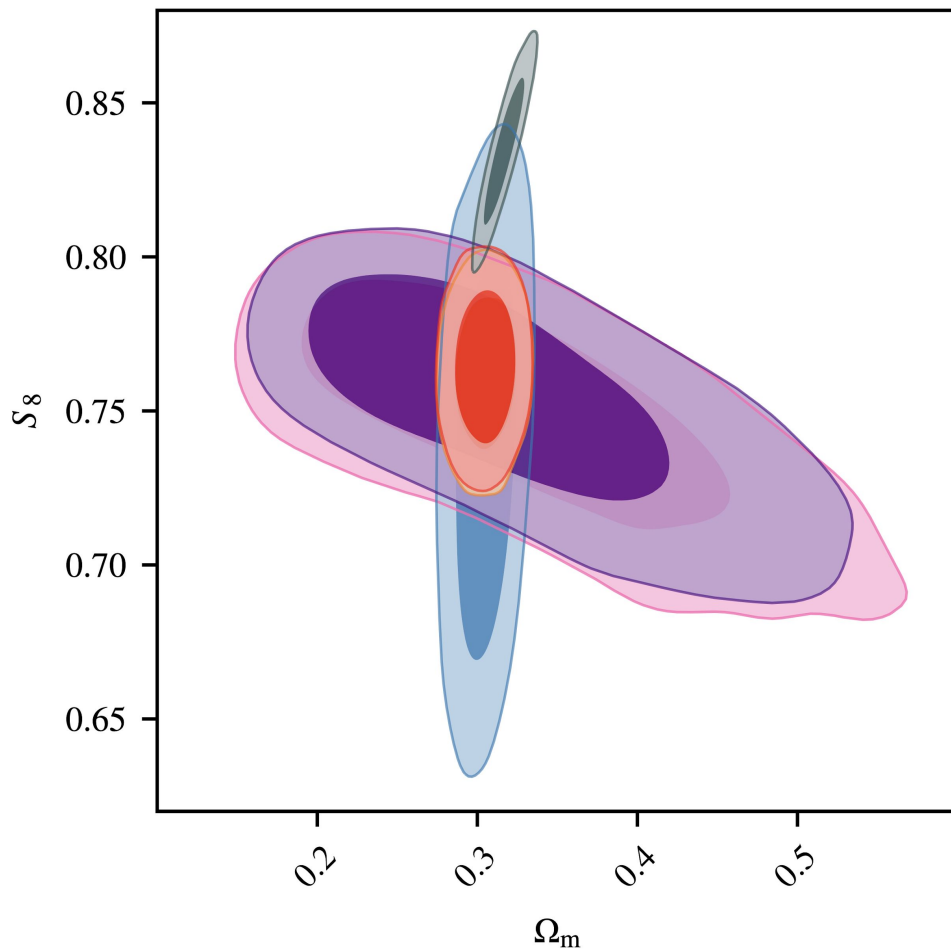
Theoretical Predictions includes fully non-linear galaxy bias model

3x2pt: Cosmic Shear + Clustering + Galaxy-Galaxy Lensing ¹⁹



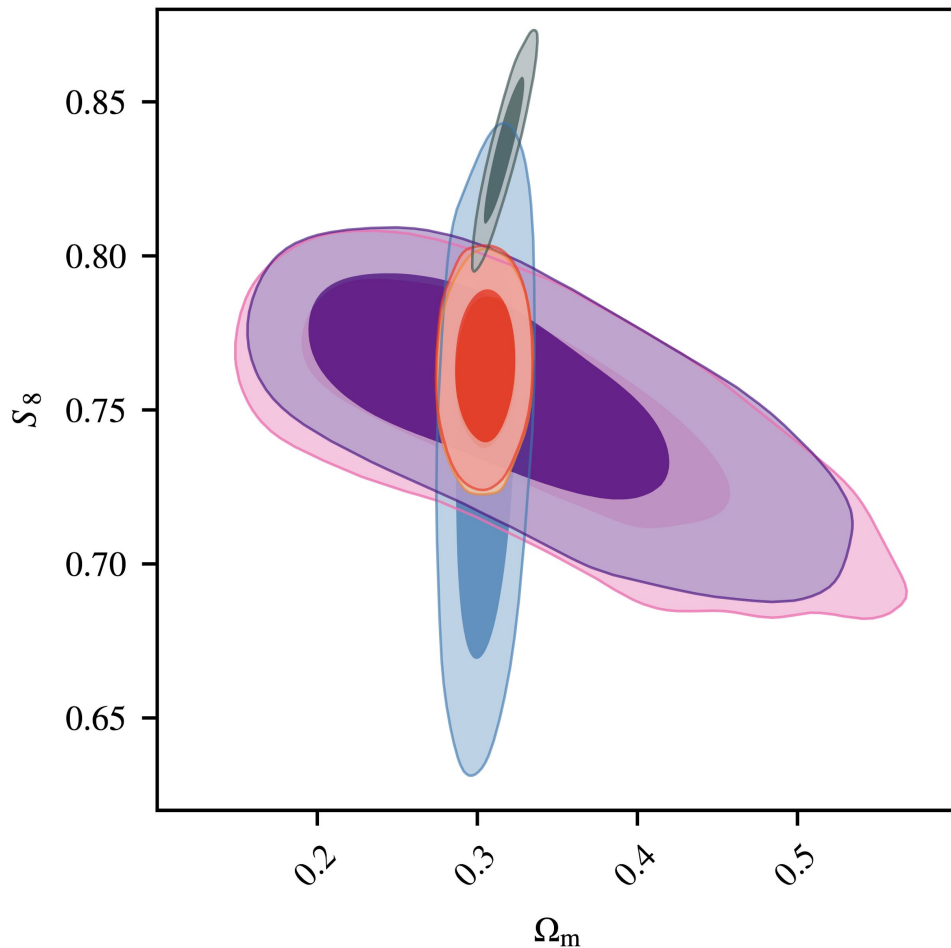
Heymans, Tröster et al. 2020

Consistency between Probes



- Cosmic shear
- Galaxy clustering
- Cosmic shear + GGL
- Cosmic shear + galaxy clustering
- KiDS-1000 3×2 pt
- Planck 2018 TTTEEE+lowE

Consistency between Probes and *Planck*



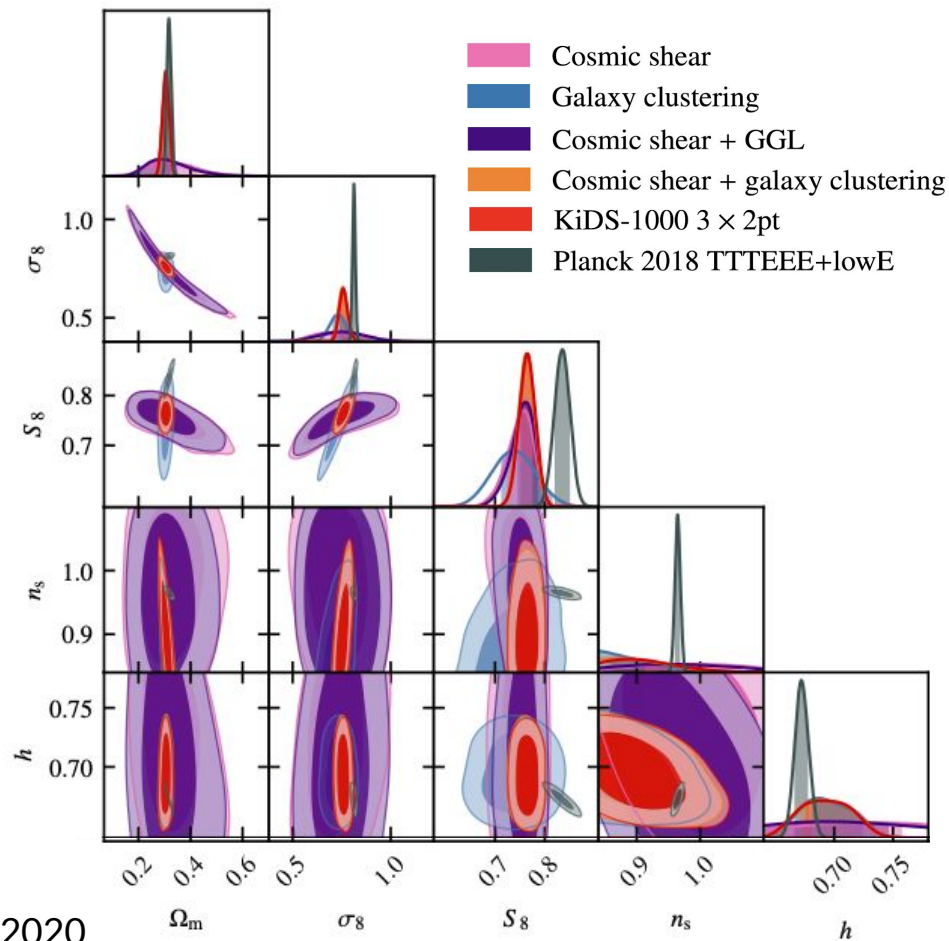
$$S_8 = \sigma_8 \sqrt{\Omega_m/0.3} = 0.766^{+0.020}_{-0.014}$$

KiDS-3x2pt and *Planck* differ in S_8
with a significance of 3.1σ

- Cosmic shear
- Galaxy clustering
- Cosmic shear + GGL
- Cosmic shear + galaxy clustering
- KiDS-1000 3×2 pt
- Planck 2018 TTTEEE+lowE

Quantifying S_8 only leads to a $\sim 3.1\sigma$ tension.

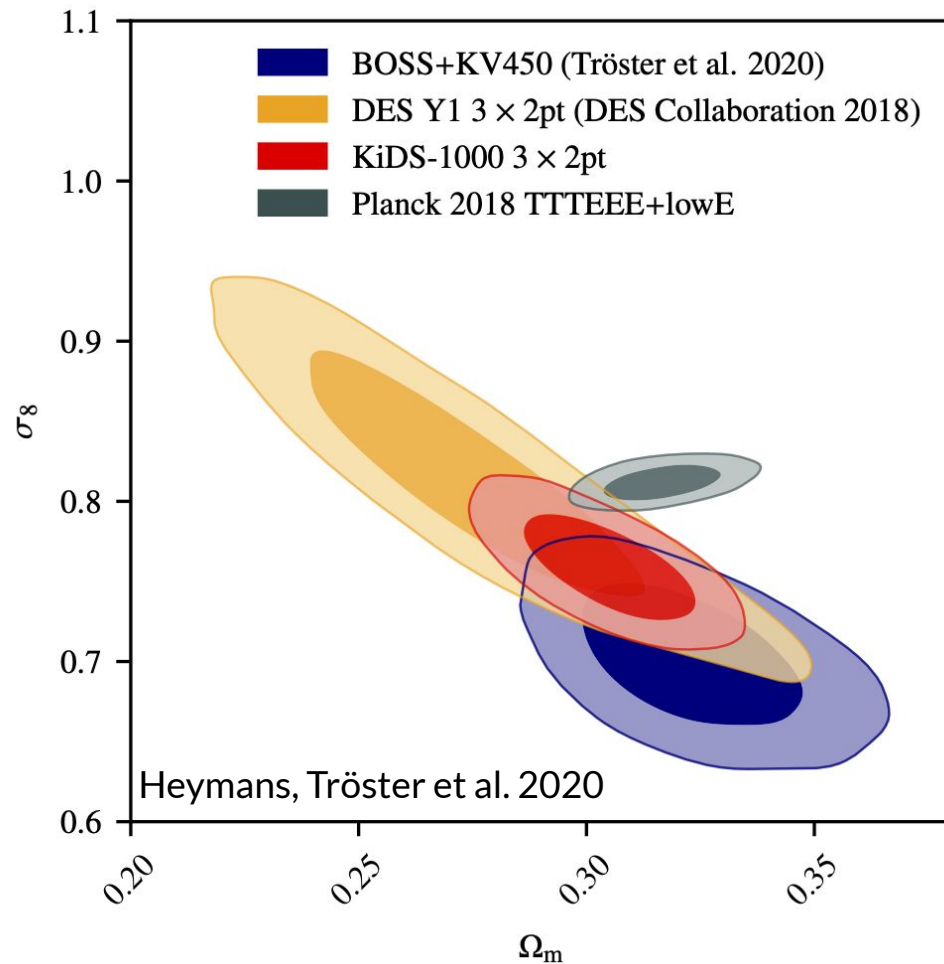
Including additional parameters, which KiDS is mainly insensitive to, dilutes the overall tension to the $\sim 2\sigma$ level.



Summary

$$S_8 = \sigma_8 \sqrt{\Omega_m/0.3} = 0.766^{+0.020}_{-0.014}$$

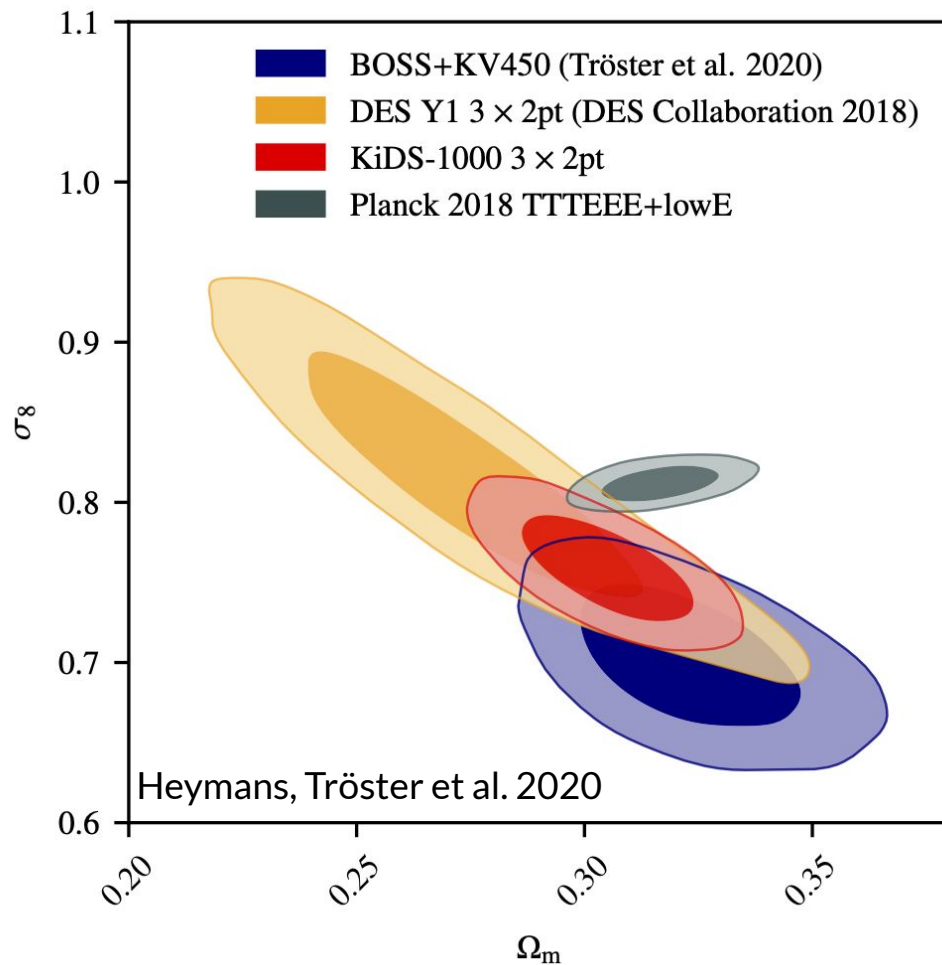
- $\sim 3\sigma$ “Tension” in S_8 is driven by differences in σ_8 . The Universe is less “clumpy” than *Planck* predicts.
- This result is validated using
 - mock KiDS and BOSS galaxy surveys
 - KiDS image simulations and null-tests
 - spectroscopic-photometric clustering analysis
 - All identified systematic uncertainties folded through as nuisance parameters



Summary

$$S_8 = \sigma_8 \sqrt{\Omega_m/0.3} = 0.766^{+0.020}_{-0.014}$$

- $\sim 3\sigma$ “Tension” in S_8 is driven by differences in σ_8 . The Universe is less “clumpy” than *Planck* predicts.



Thanks to KiDS and all our funders



European Research Council
Established by the European Commission



MARIE CURIE



Netherlands Organisation
for Scientific Research



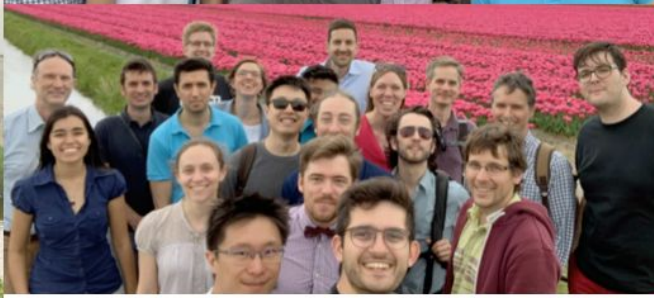
Alexander von Humboldt
Stiftung/Foundation



KiDS Pre-lockdown



KiDS in-lockdown



DFG
Deutsche Forschungsgemeinschaft



Australian Government
Australian Research Council



Looking forward to your questions!



Image Credit: Giblin, Kuijken and the KiDS team. Original Background Credit: ESO/Beletsky